

Non-thermal Electron Acceleration in Low Mach Number Collisionless Shocks

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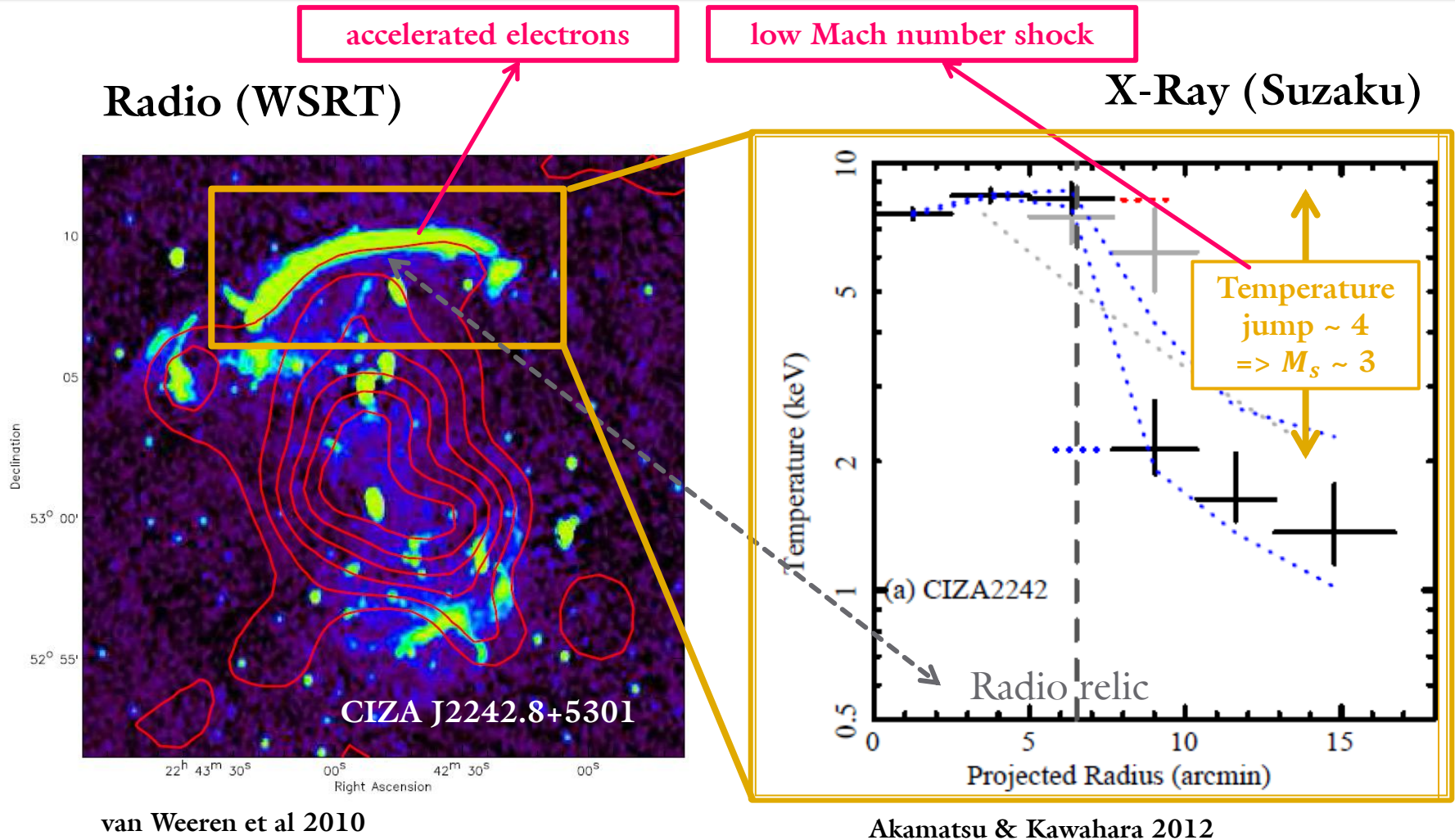
In collaboration with
Lorenzo Sironi and Ramesh Narayan
ApJ 794 ,153, [arXiv1406.5190](https://arxiv.org/abs/1406.5190)
ApJ in presss, [arXiv1409.7393](https://arxiv.org/abs/1409.7393)

Outline

- Motivation
- Simulation setup
- Shock structure and energy spectra
- Acceleration mechanism
- Parameter dependence
- Astrophysical implications

Observational Evidence

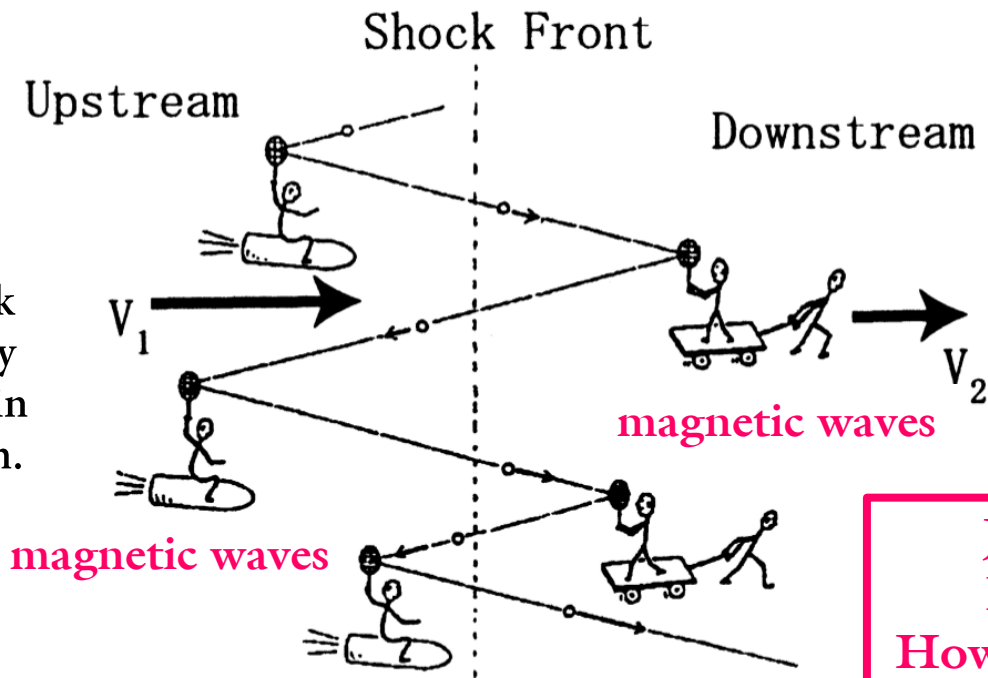
Radio relics at outskirts of galaxy clusters



Particle acceleration at shocks

First-order Fermi acceleration

Particles gain energy after each shock crossing .



Particles cross shock due to scatterings by the **magnetic waves** in up- and downstream.

Just a paradigm.
Relies on waves.
How do the waves arise?

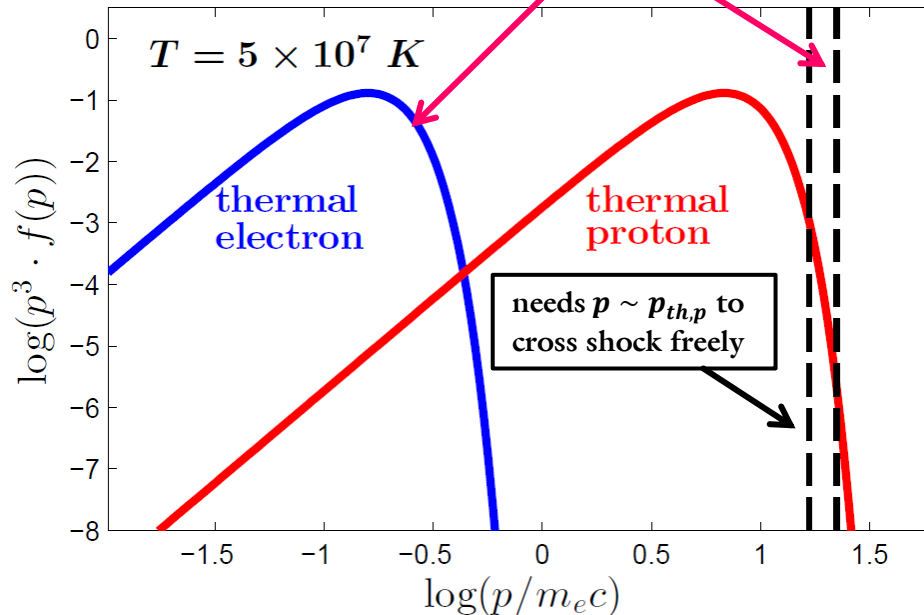
(original sketch by Scholer)
Credit: Hoshino 2001

Theoretical Challenge

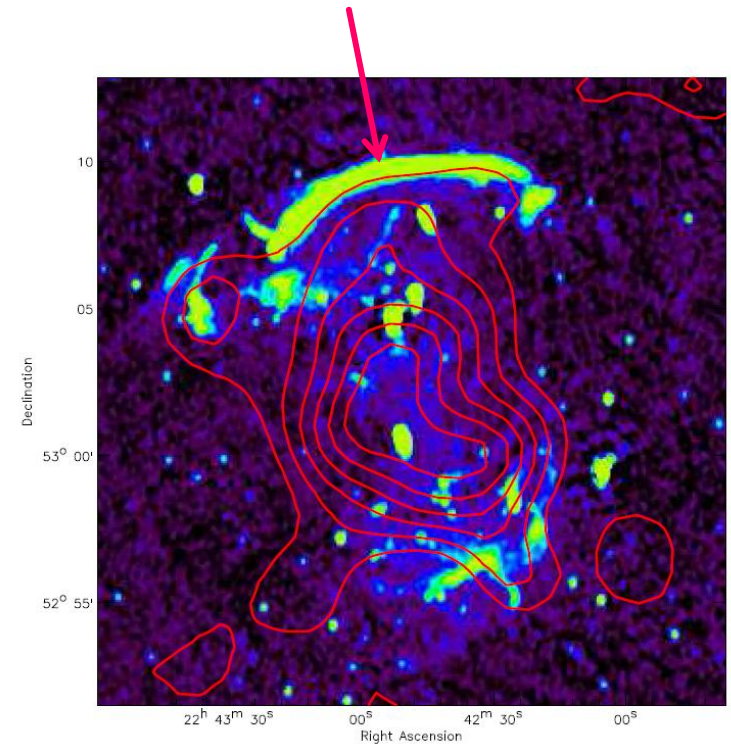
electron injection problem

Traditional injection model:
assumes particles need to cross shock from downstream

electron acceleration efficiency too poor.



In tension with observations!



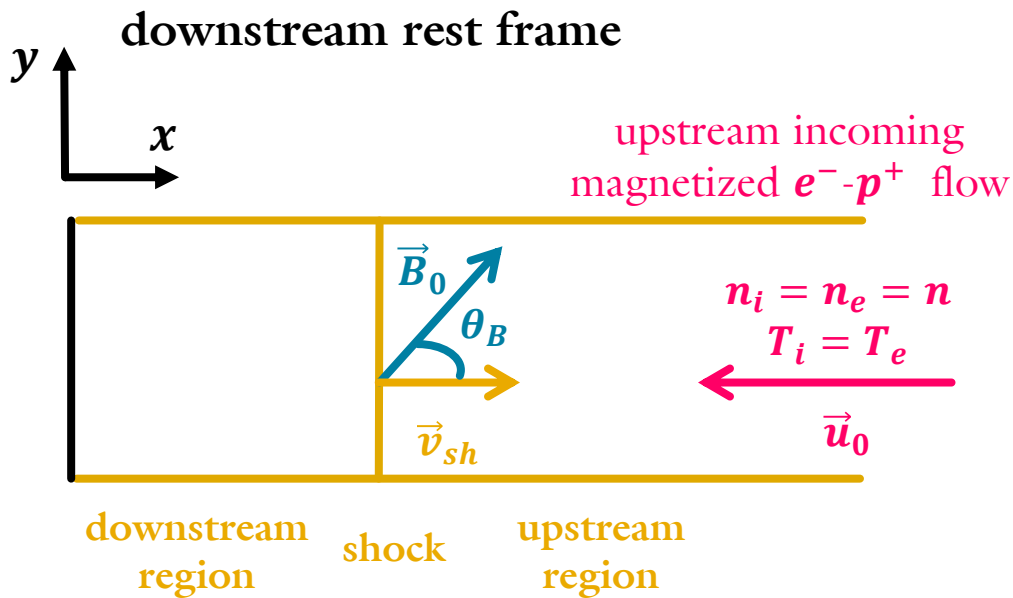
This work: studies electron acceleration in low Mach number shock from first principles, identifies an **electron Fermi-type acceleration** mechanism, which is mediated by **electron self-generated waves** and operates efficiently over a wide range of parameter space.

Simulation Setup

Self-consistent particle-in-cell plasma simulations

Method: Particle-in-cell (PIC) method, simulates collisionless plasma from first principles.

Code: TRISTAN-MP (Spitkovsky 2005)



Mach number: $M_s = \frac{u_0 + v_{sh}}{C_s}$

plasma beta : $\beta_p = \frac{nk_B(T_i + T_e)}{B_0^2/8\pi}$

Simulation runs:

$M_s = 3$ fixed

T_e varies from $10^{7.5}\text{K}$ to 10^9K

β_p varies from 6 to 60

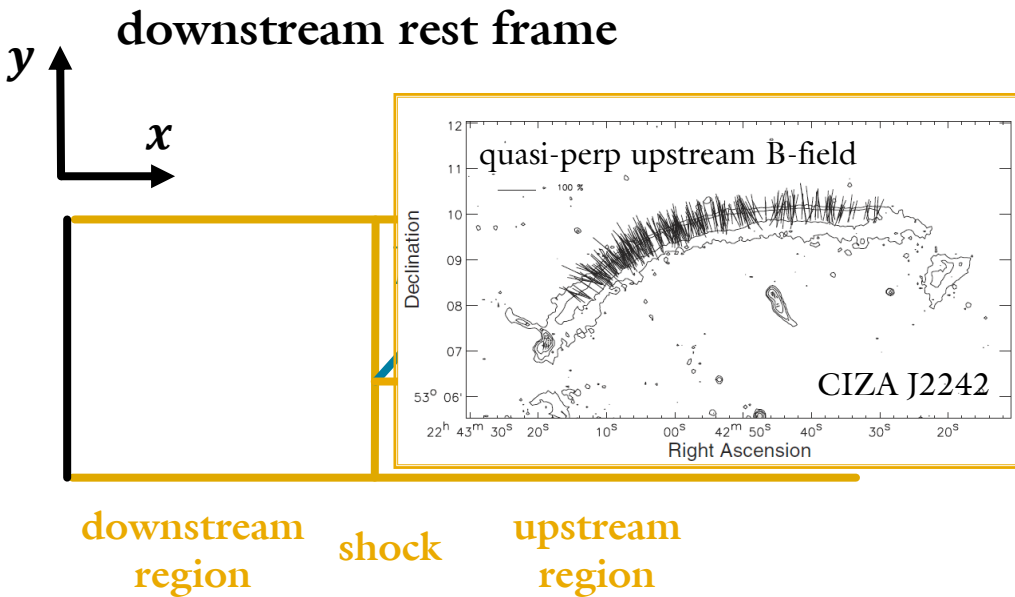
θ_B varies from 13° to 80°

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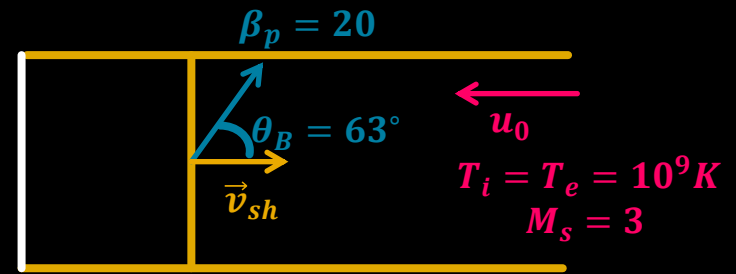
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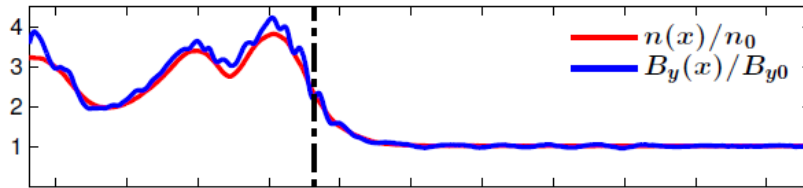
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Shock Structure

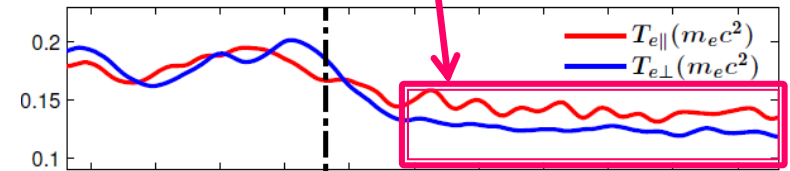


downstream shock upstream

density/B-field profile

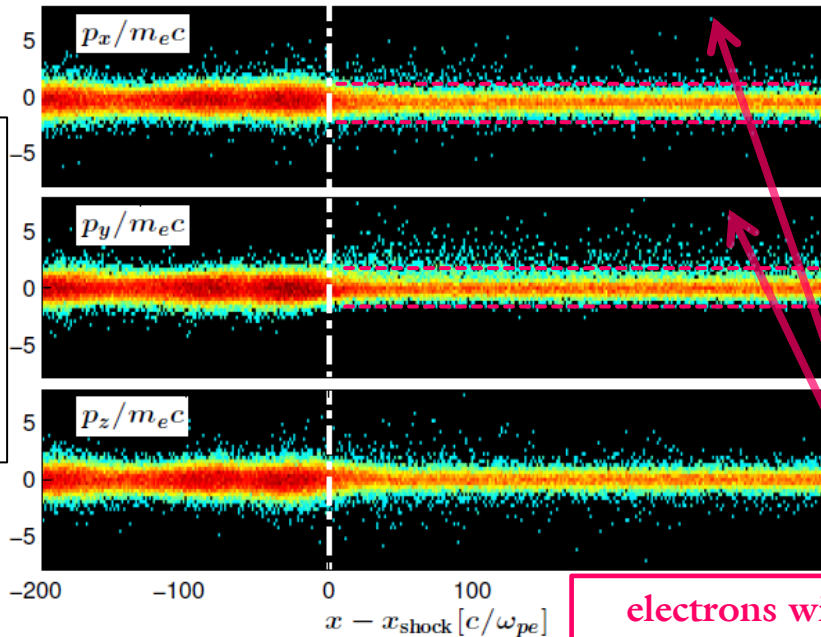


electron temperature



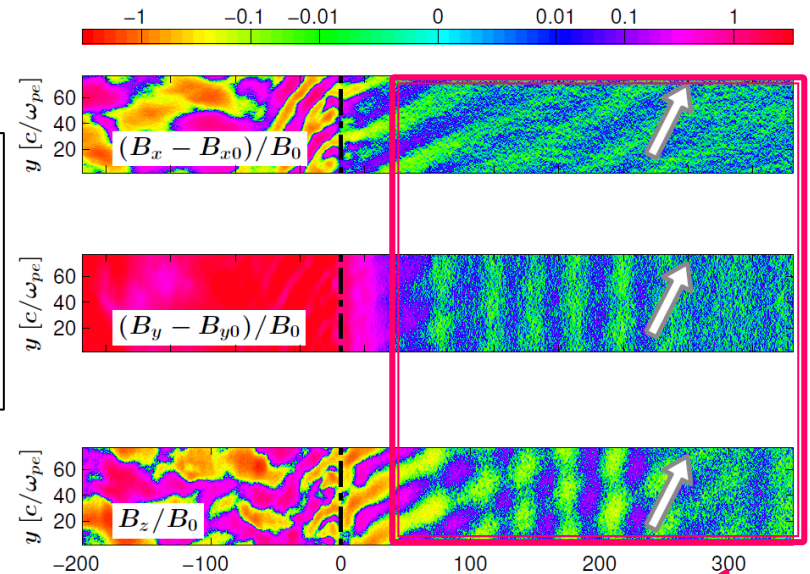
electron temperature anisotropy in the upstream

electron phase space



electrons with very large momentum in the upstream

magnetic waves

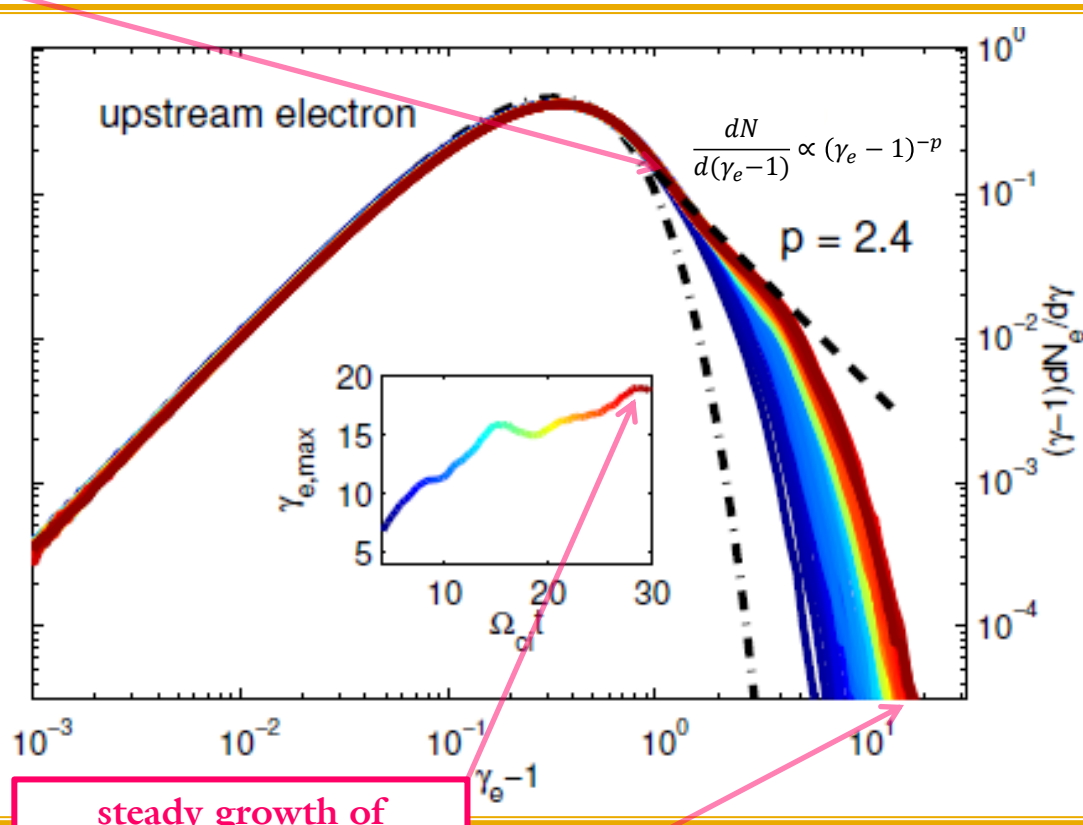


upstream waves move towards the shock

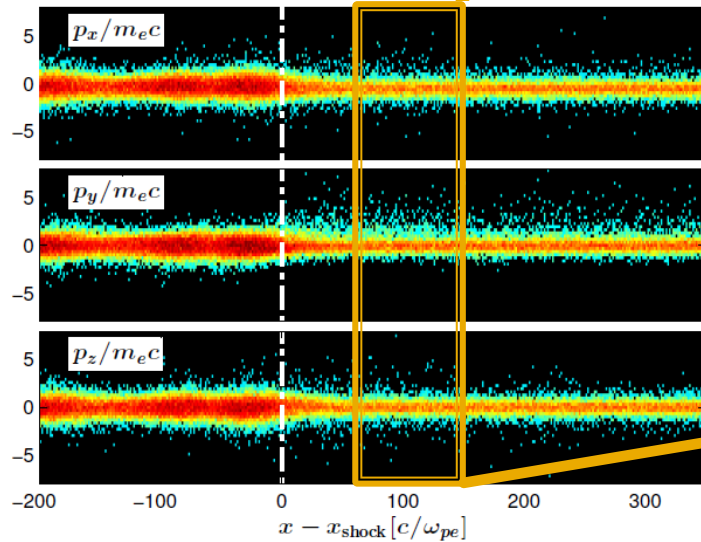
\vec{v}_{wave}

Electron Spectral Evolution

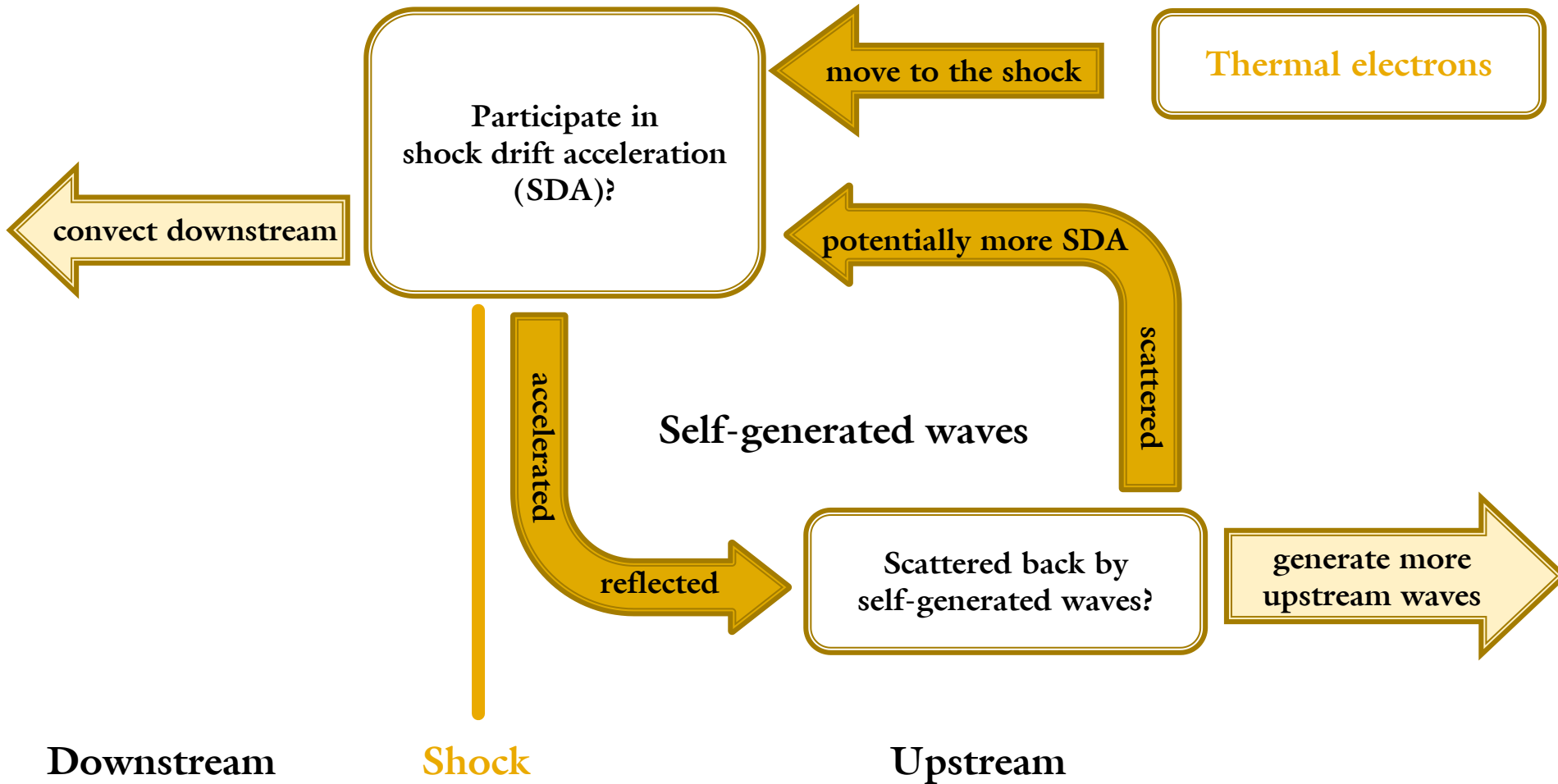
about 10% upstream electrons are accelerated to non-thermal energy



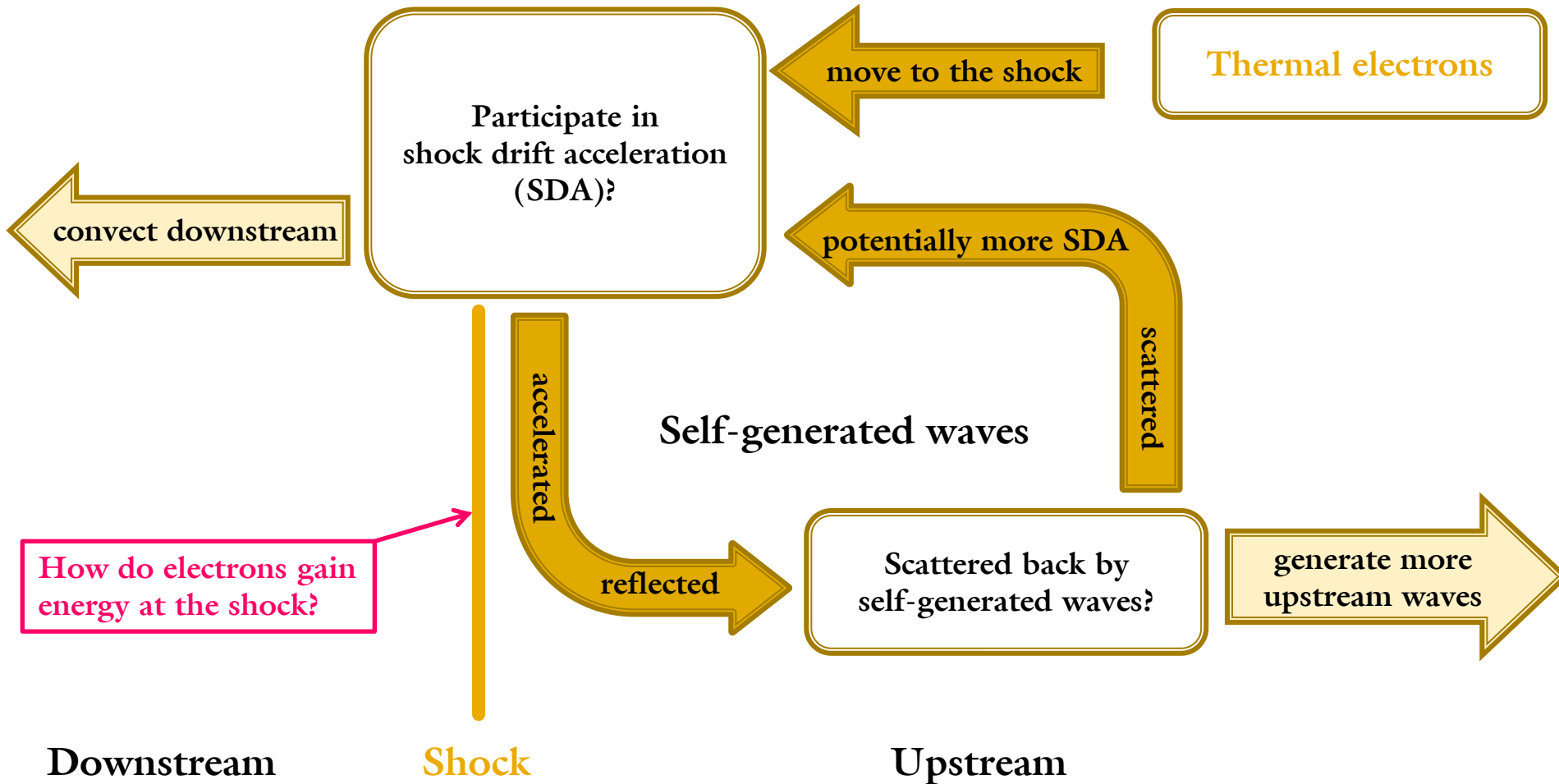
steady growth of maximum energy, efficient acceleration



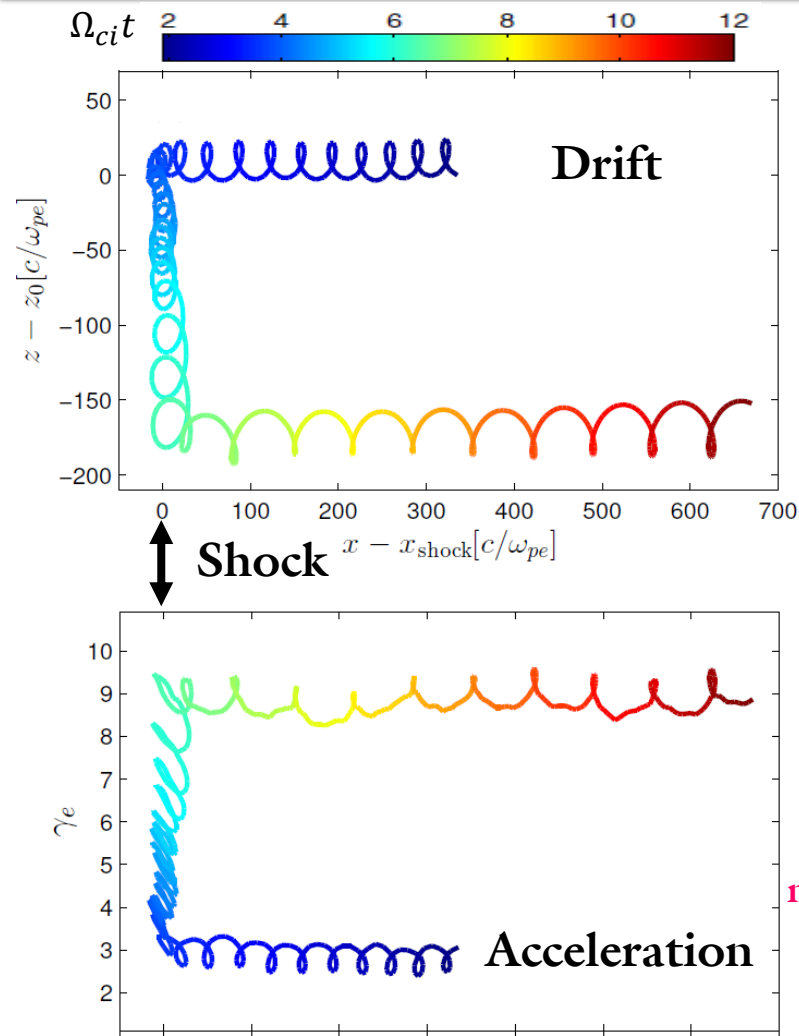
The Big Picture



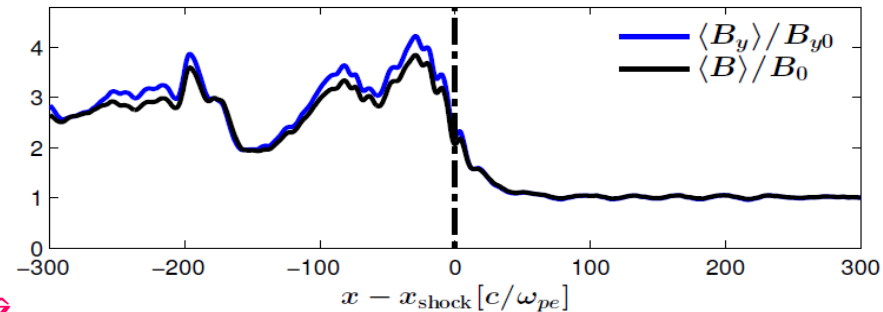
The Big Picture



Shock Drift Acceleration

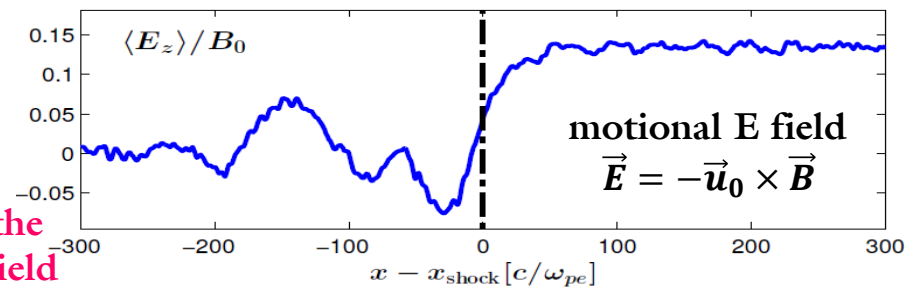


grad-B force
 ←
 drift along $-\hat{z}$



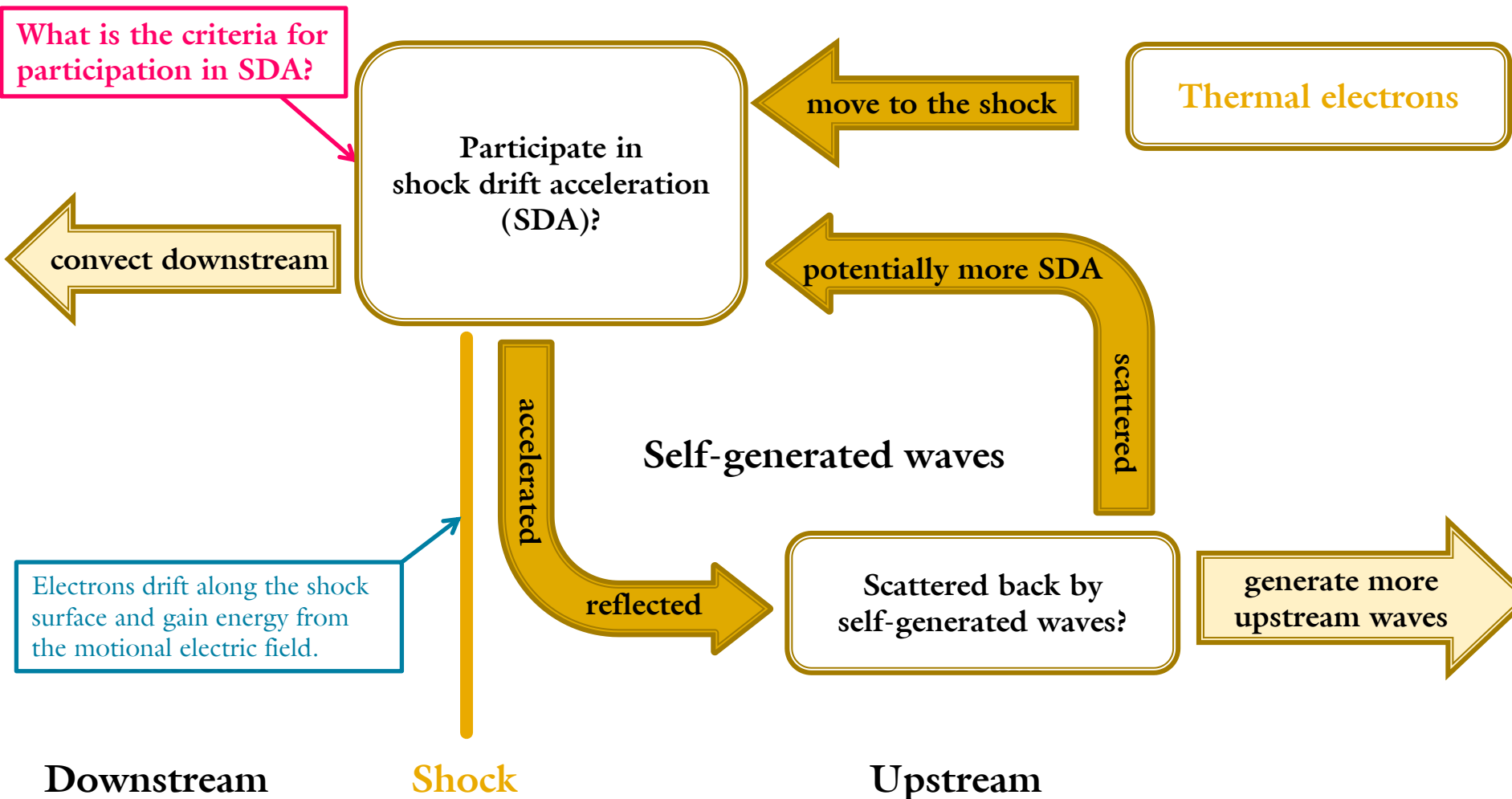
$$\vec{v}_{\nabla B} = \mu/q(\vec{B} \times \nabla B)/B^2$$

drift along the
 motional E field
 ←
 energy gain



$$\Delta\gamma_e m_e c^2 = q \int E_z dz$$

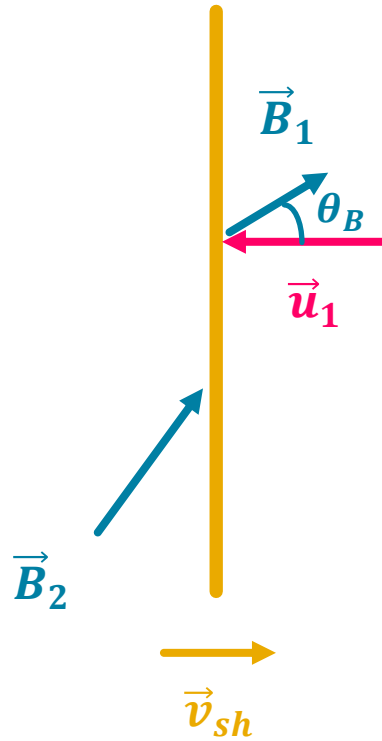
The Big Picture



Frame Transformation

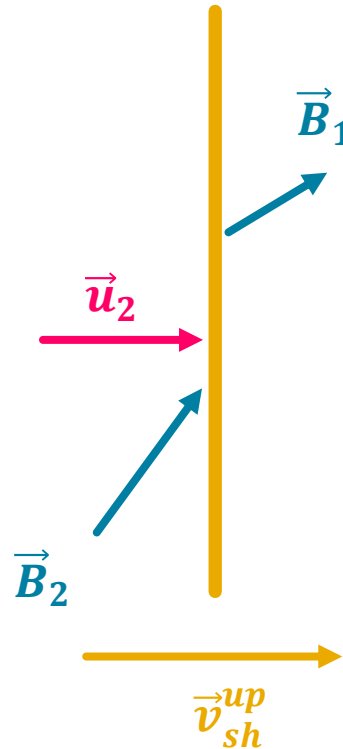
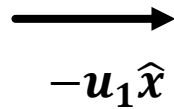
downstream rest frame

also the simulation frame,
most straightforward!



upstream rest frame

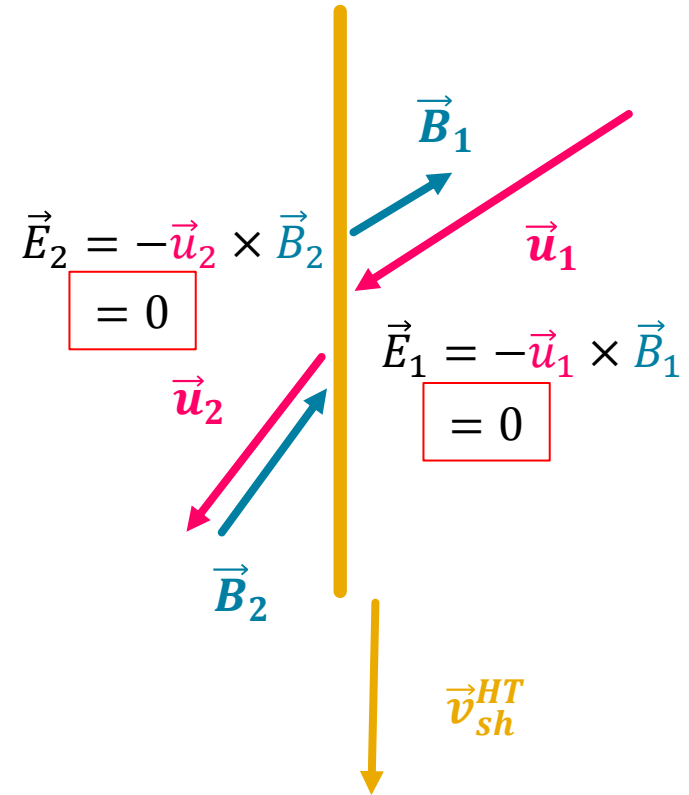
upstream plasma velocity
space independent of bulk
motion



de Hoffman-Teller frame

No motional electric field!

$$-u_t c \hat{B}_1 = -v_{sh}^{up} / \cos \theta_B \hat{B}_1$$



Injection into SDA

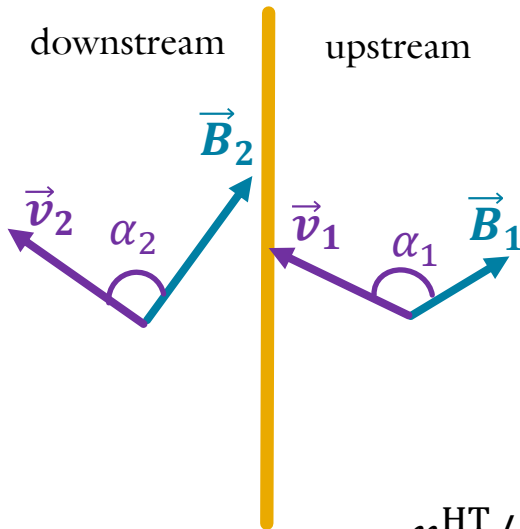
governed by mirror reflection at the shock

HT frame

No motional electric field

downstream

upstream



Assumptions:

Energy conservation

Magnetic moment conservation

If transmitted

$$\sin \alpha_2^2 = \frac{B_2}{B_1} \sin \alpha_1^2 \leq 1$$

Reflection Criteria

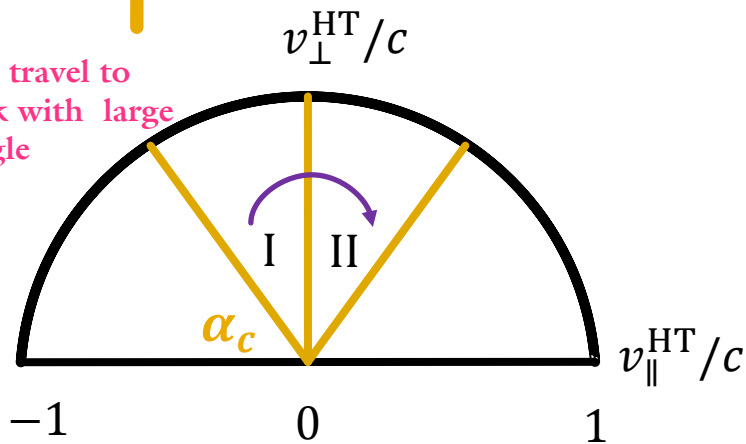
$$v_{1\parallel}^{\text{HT}} < 0$$

$$\alpha_1^{\text{HT}} > \alpha_c \equiv \sin^{-1} \sqrt{\frac{B_2}{B_1}}$$

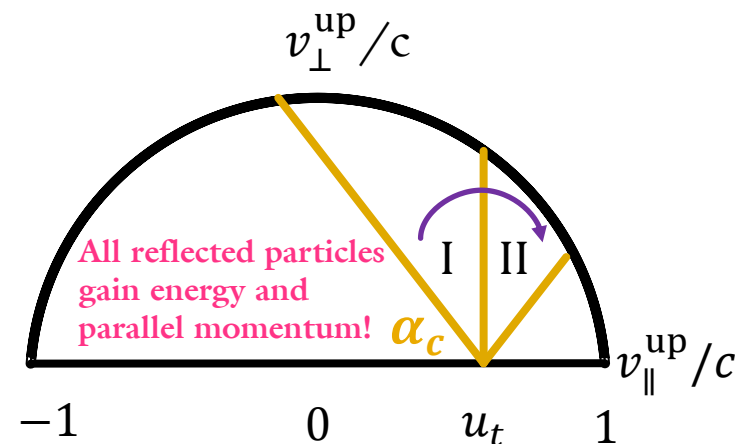
After reflection

$$v_{1\parallel}^{\text{HT}} \Rightarrow -v_{1\parallel}^{\text{HT}}$$

Needs to travel to the shock with large pitch angle



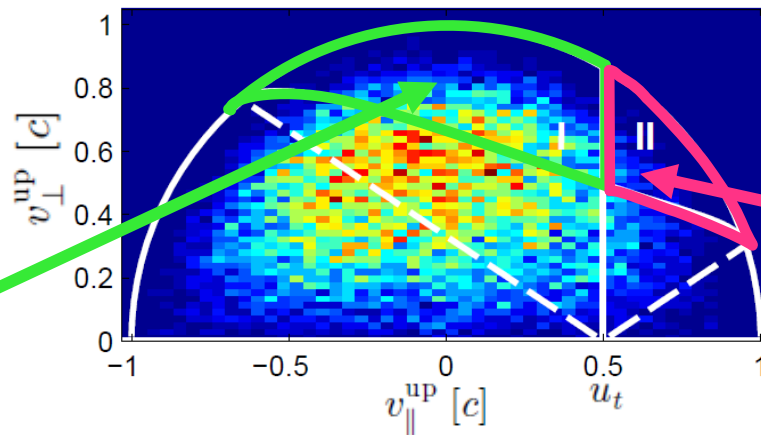
$u_t \hat{B}_1$



Injection into SDA

mirror reflection criteria verified by particle tracing

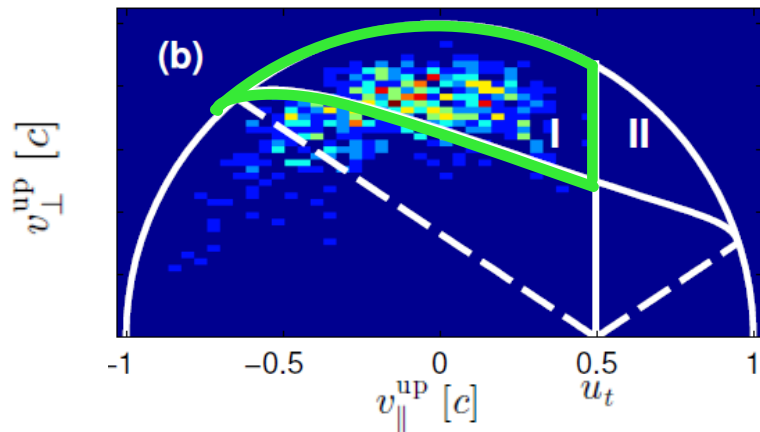
representative particles at initialization



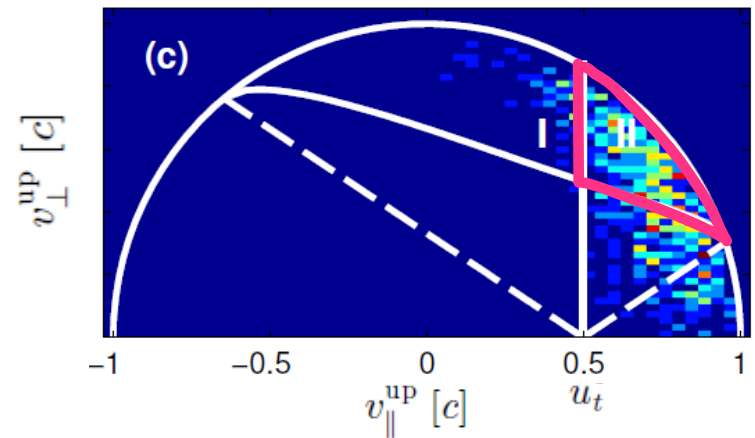
region in velocity space allowed for reflection

region in velocity space occupied by reflected particles

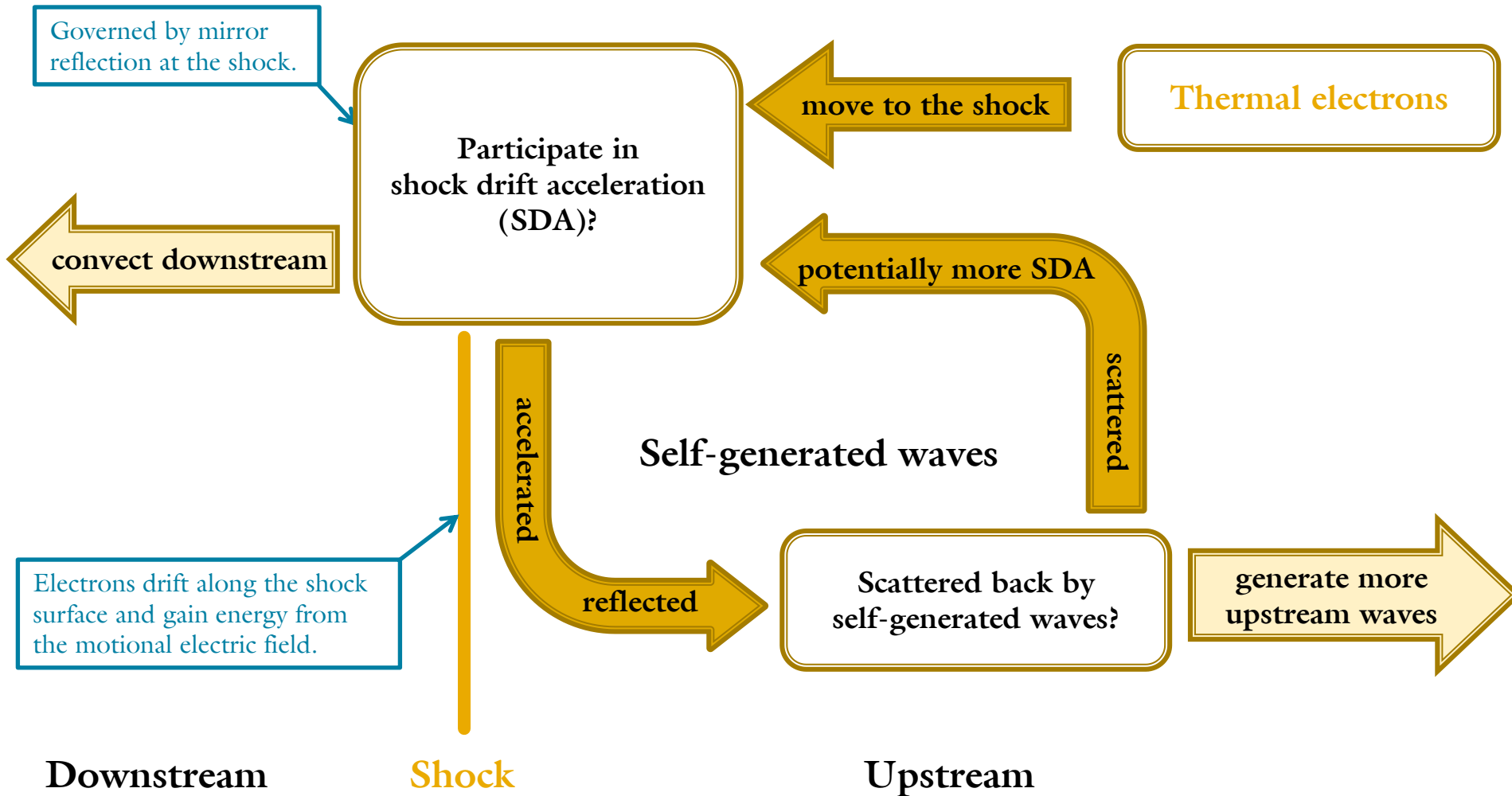
reflected particles **before** reflection



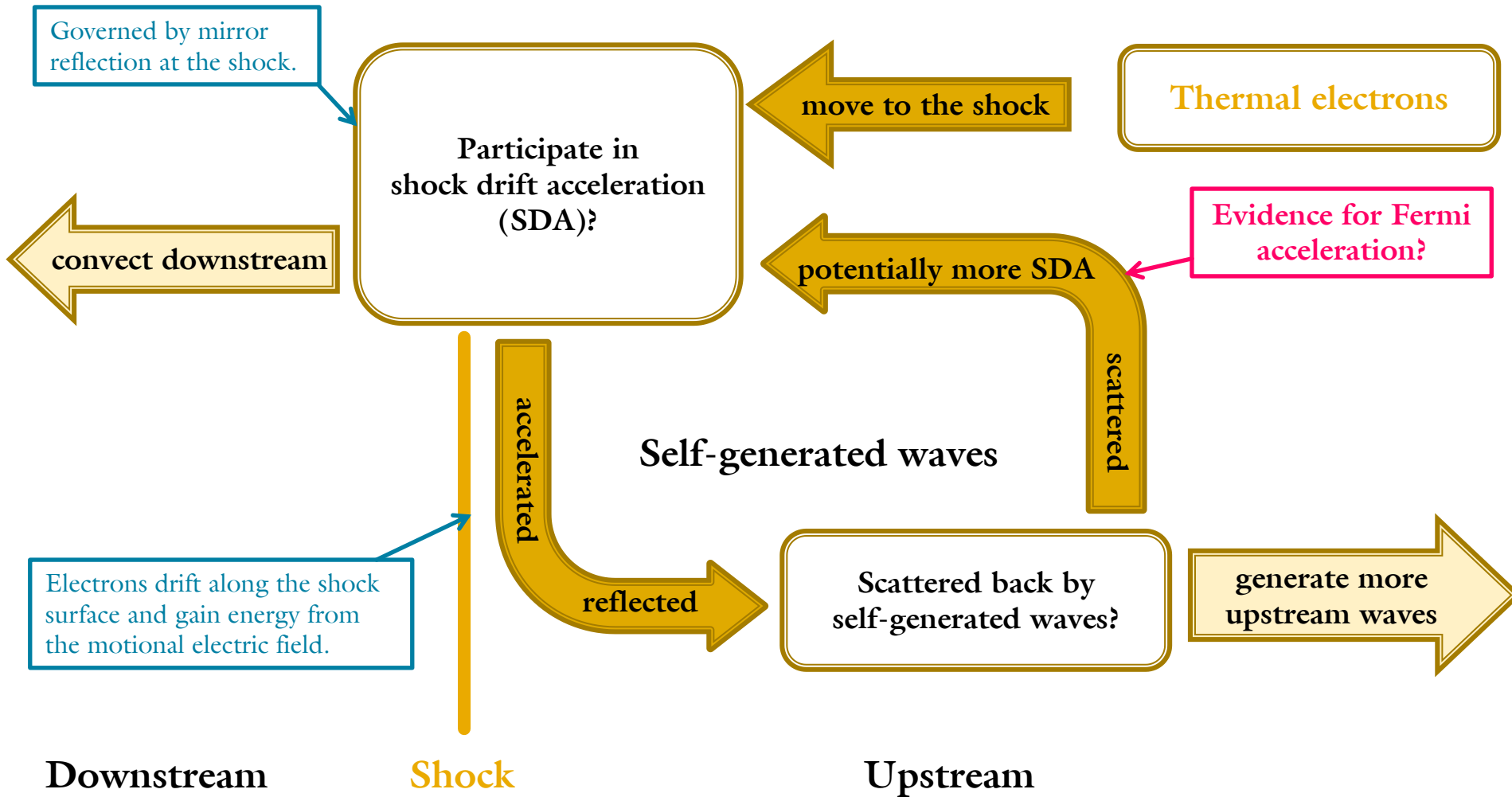
reflected particles **after** reflection



The Big Picture

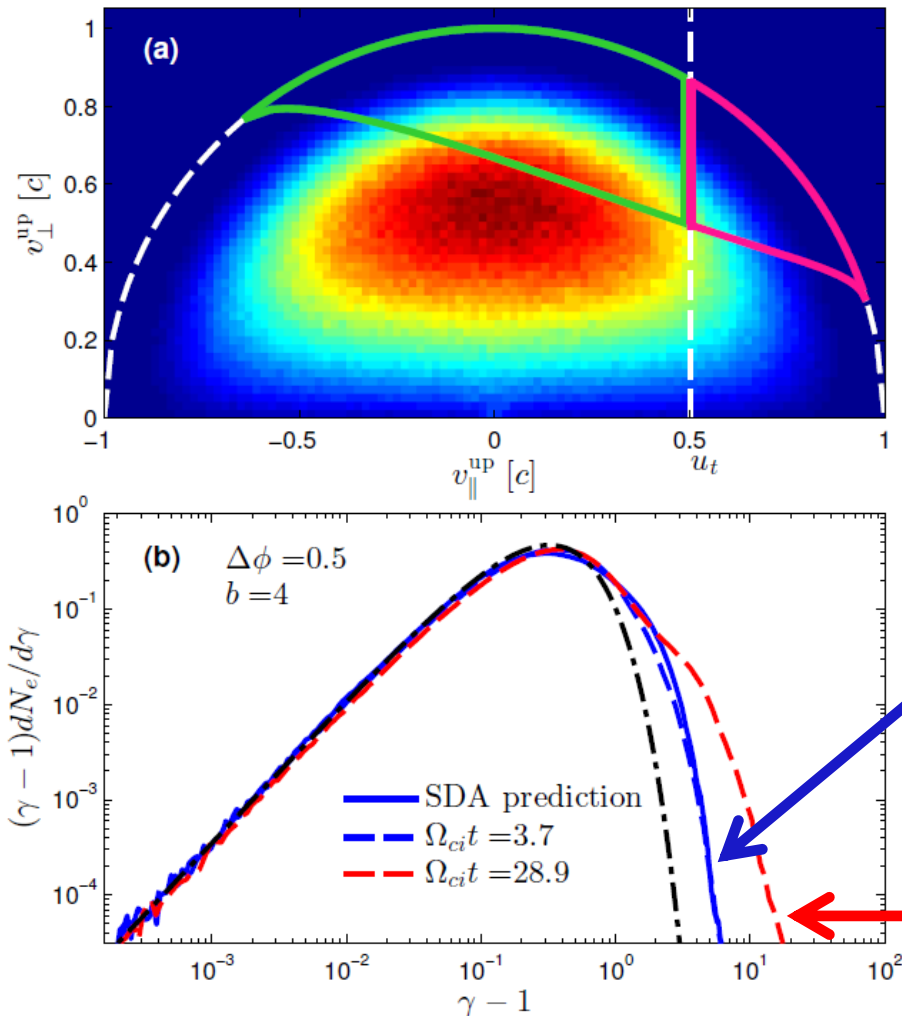


The Big Picture



Beyond single-cycle SDA

single-cycle SDA cannot explain long-term spectral evolution

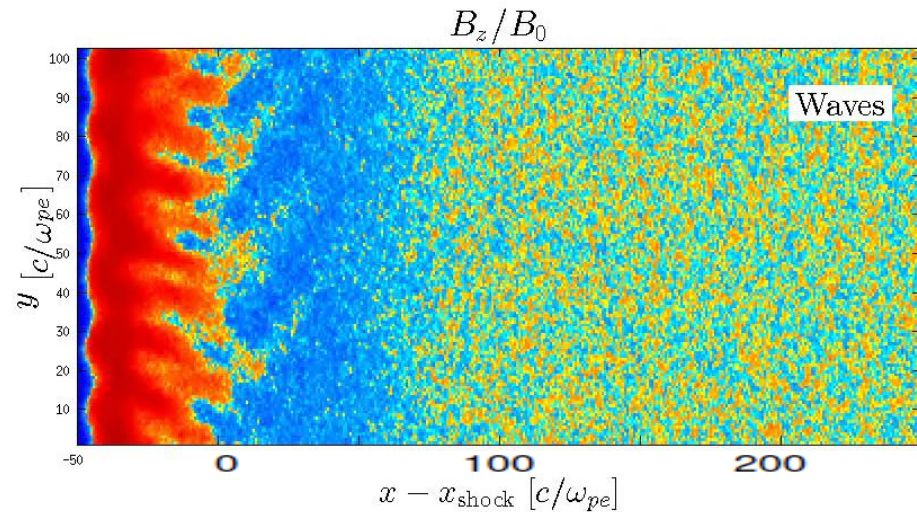
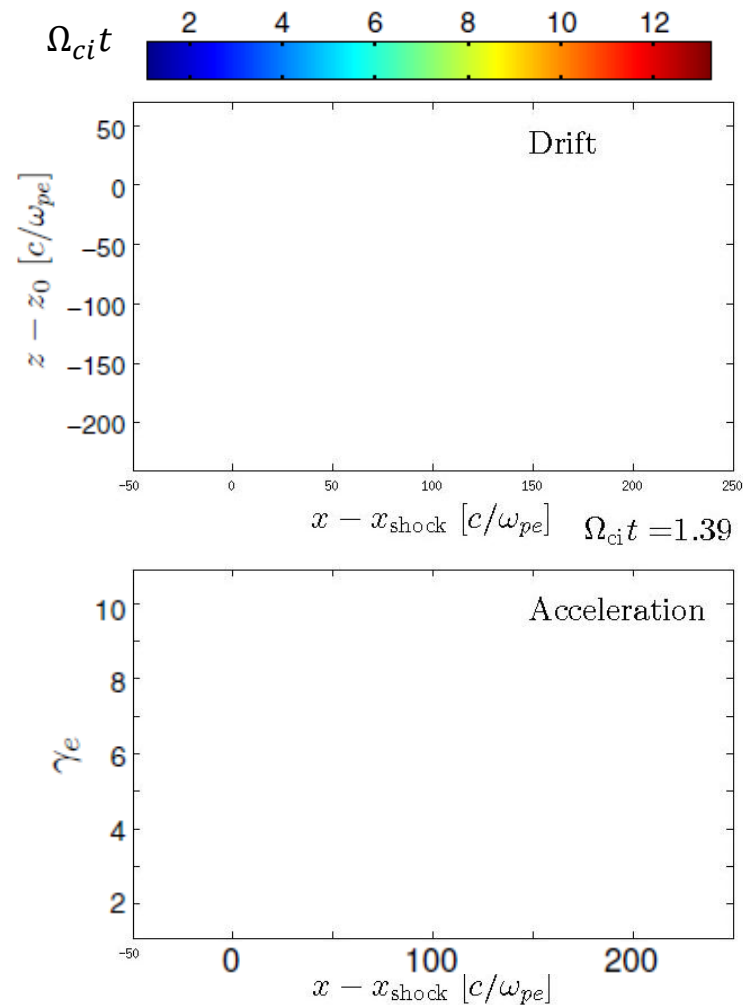


Single-cycle SDA agrees with spectrum at early times.

Single-cycle SDA is not enough to explain the long term spectral evolution.

Fermi-like Acceleration

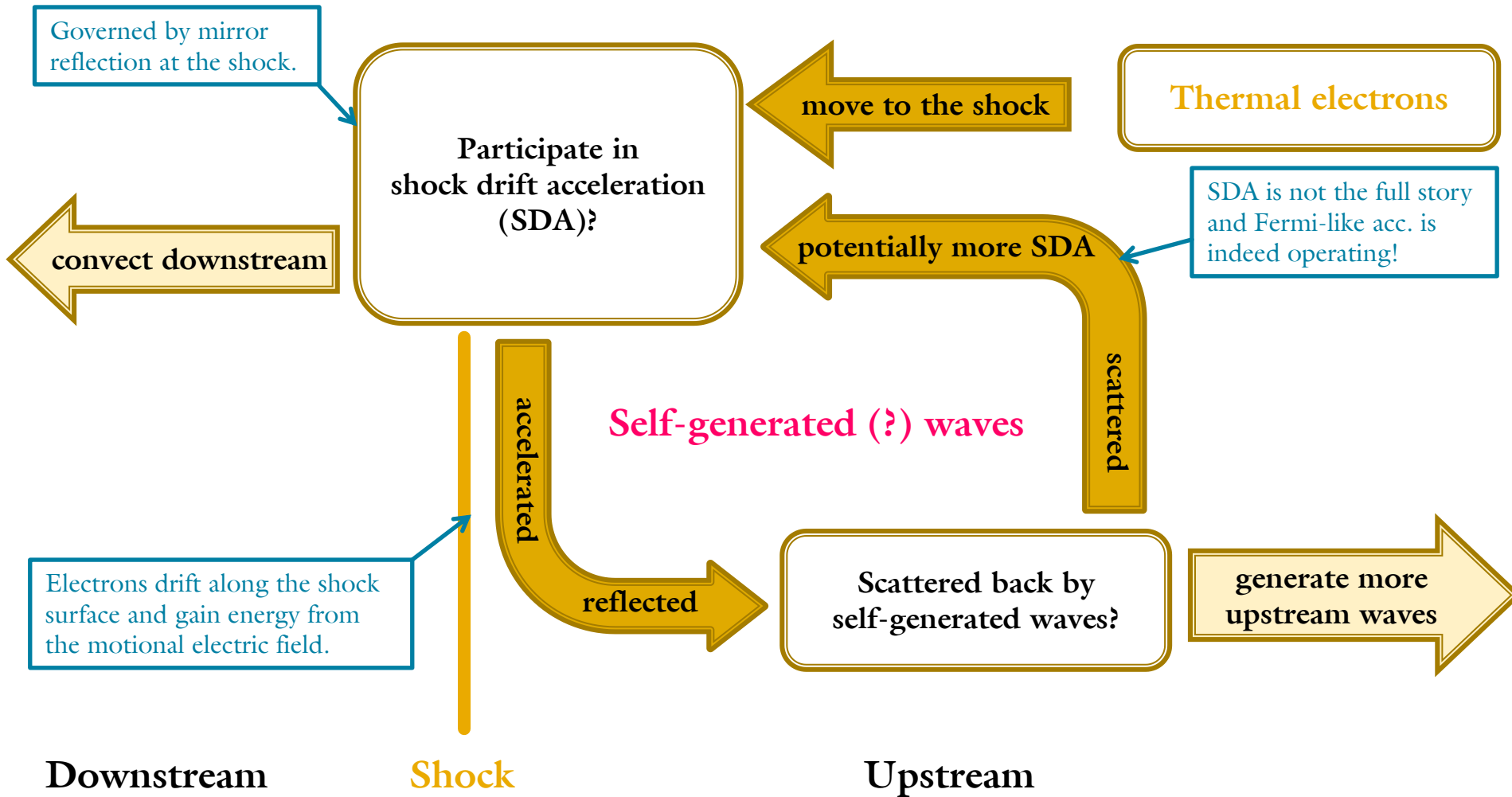
multiple cycles of SDA mediated by upstream waves



video link:

<http://youtu.be/asXmDyNSL8s>

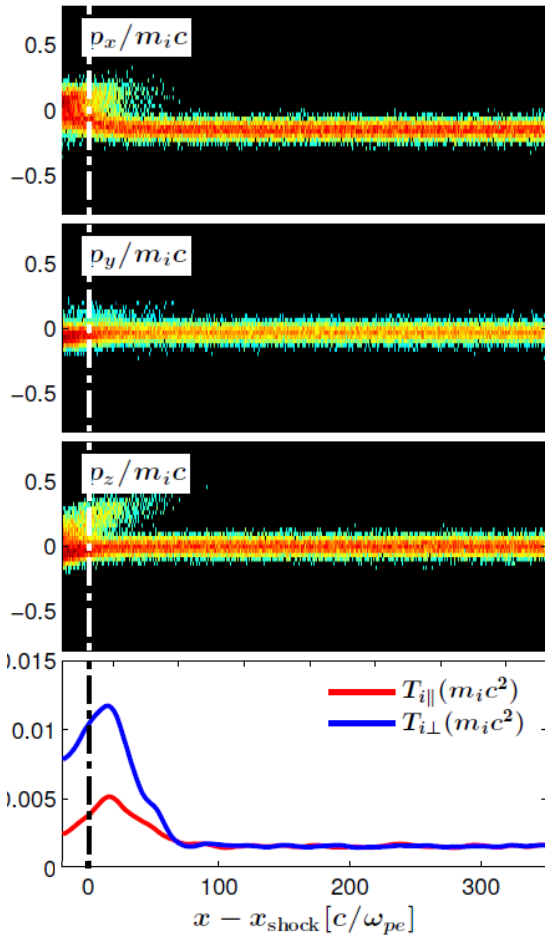
The Big Picture



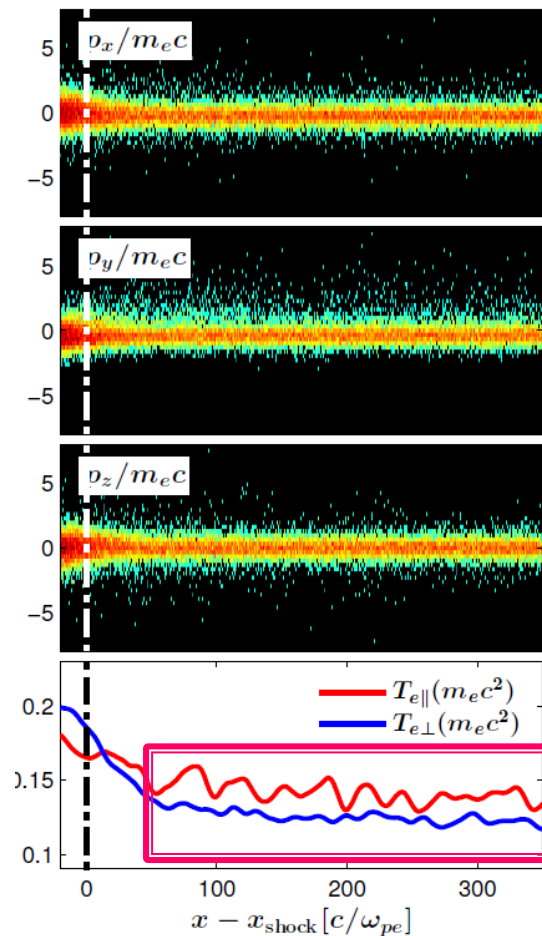
Waves generated by e^- or p^+ ?

p^+ isotropic vs. e^- temperature anisotropy provides source of free energy

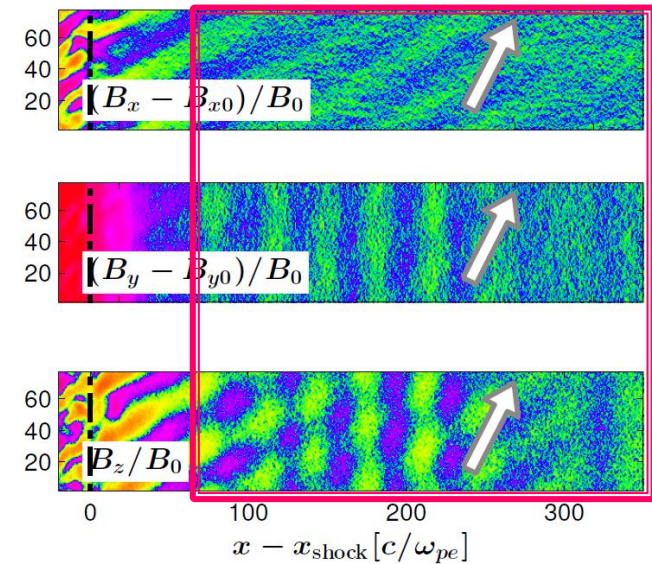
Ions



Electrons



Waves

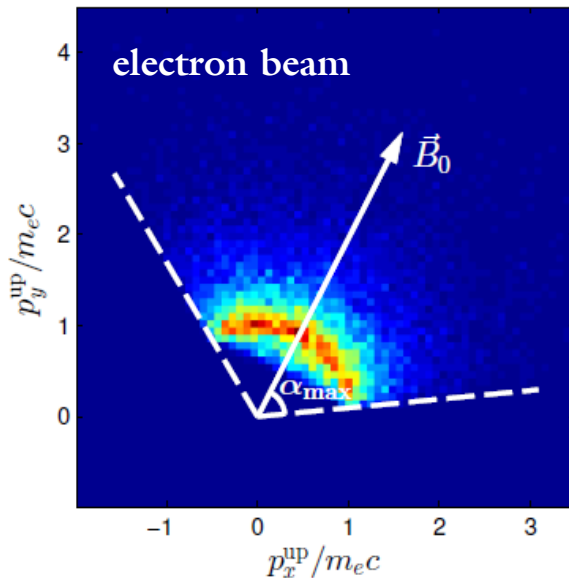


e^- self-generate the waves

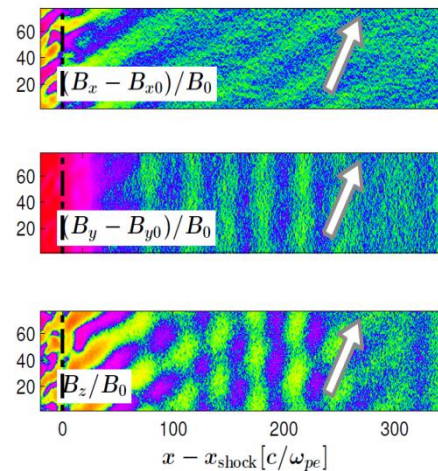
controlled beam-plasma instability experiment reproduces the wave patterns

Controlled Beam-Plasma Instability Experiment

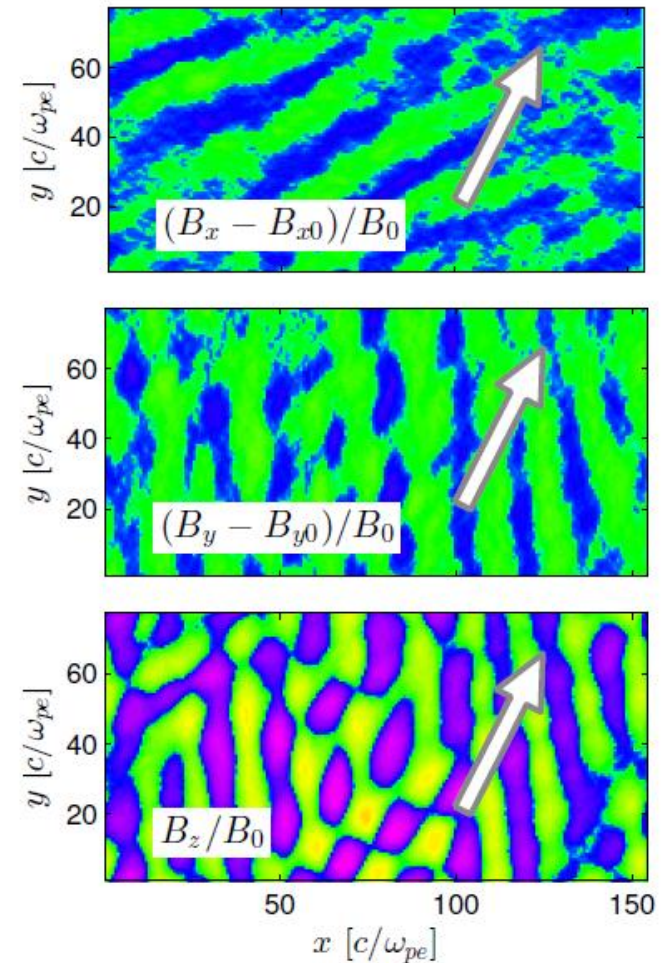
- Upstream rest frame
- Periodic boundary conditions
- Maxwellian magnetized background plasma
- Electron beam modeled by SDA injection theory



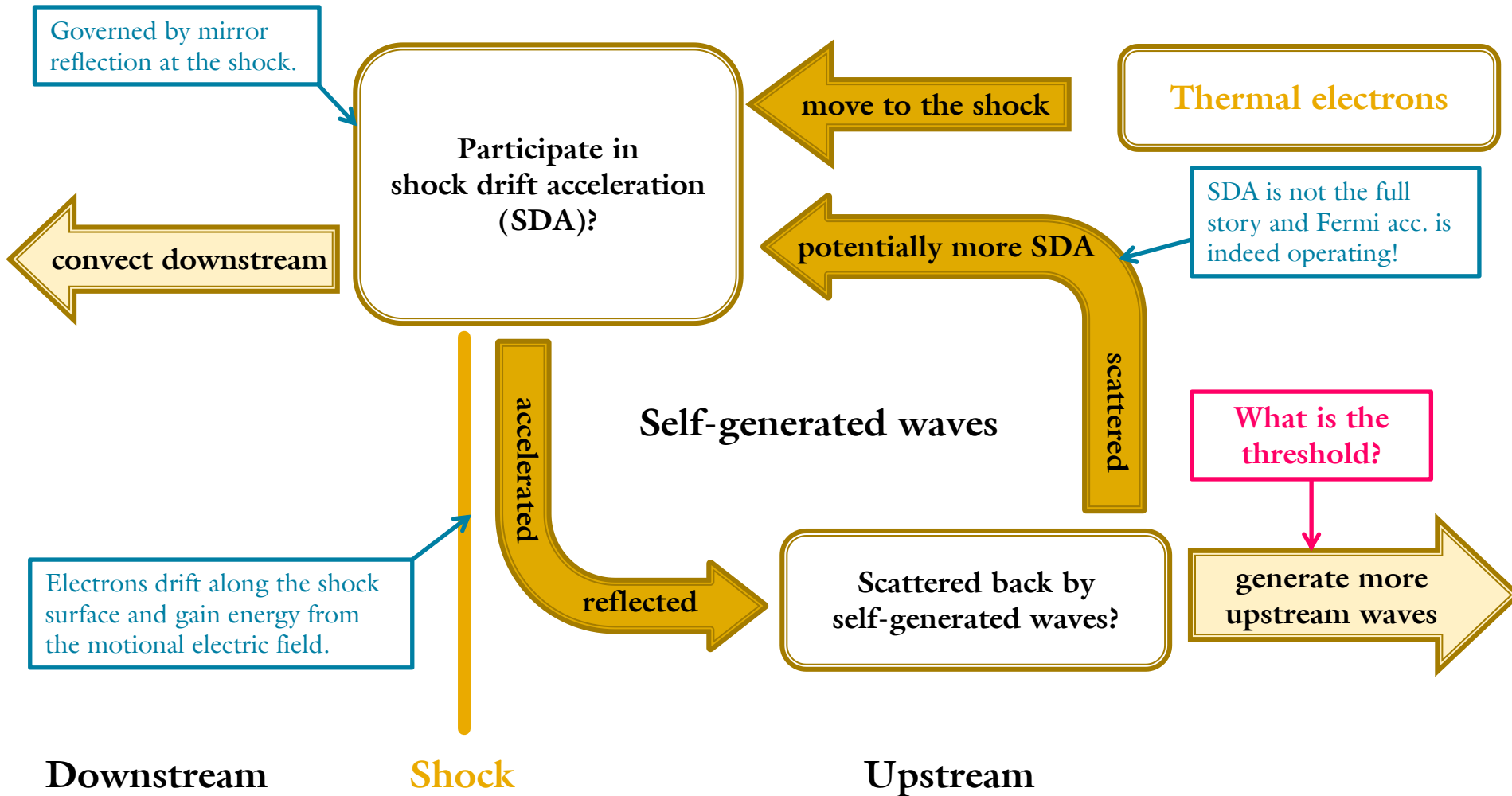
shock simulation



controlled experiment



The Big Picture



Threshold for the wave growth

e^- pressure anisotropy $>$ background magnetic pressure

Electron firehose instability
(oblique mode)

Pressure due to electron temperature anisotropy.

$$\frac{n_e k T_{e\parallel} - n_e k T_{e\perp}}{B_0^2 / 8\pi} > 1.3$$

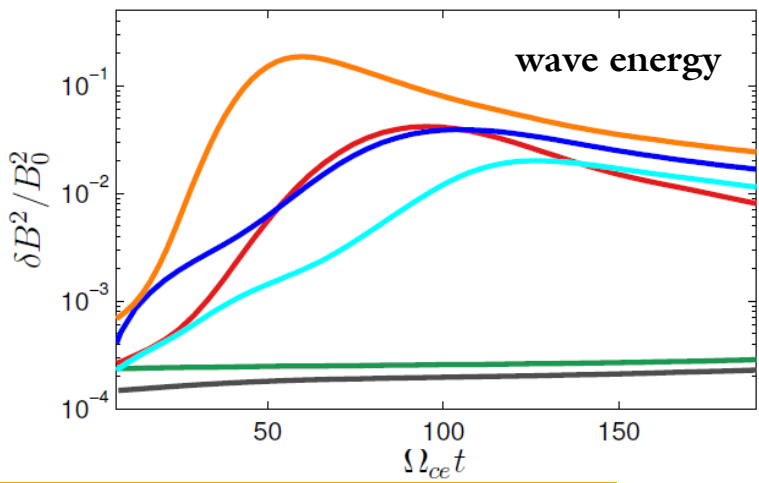
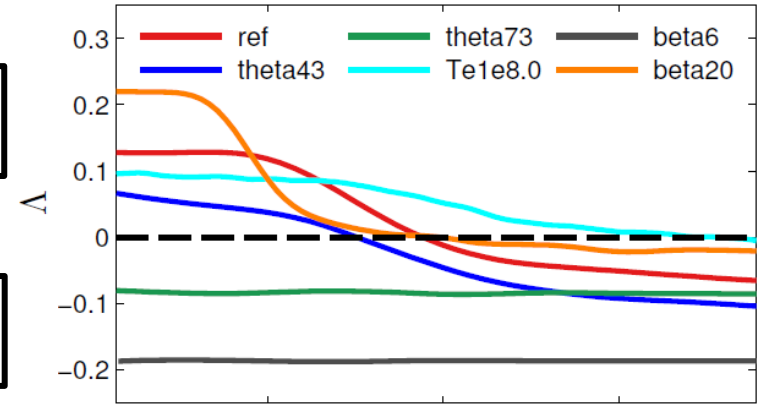
Background magnetic pressure

above threshold, waves grow

below threshold, no waves grow

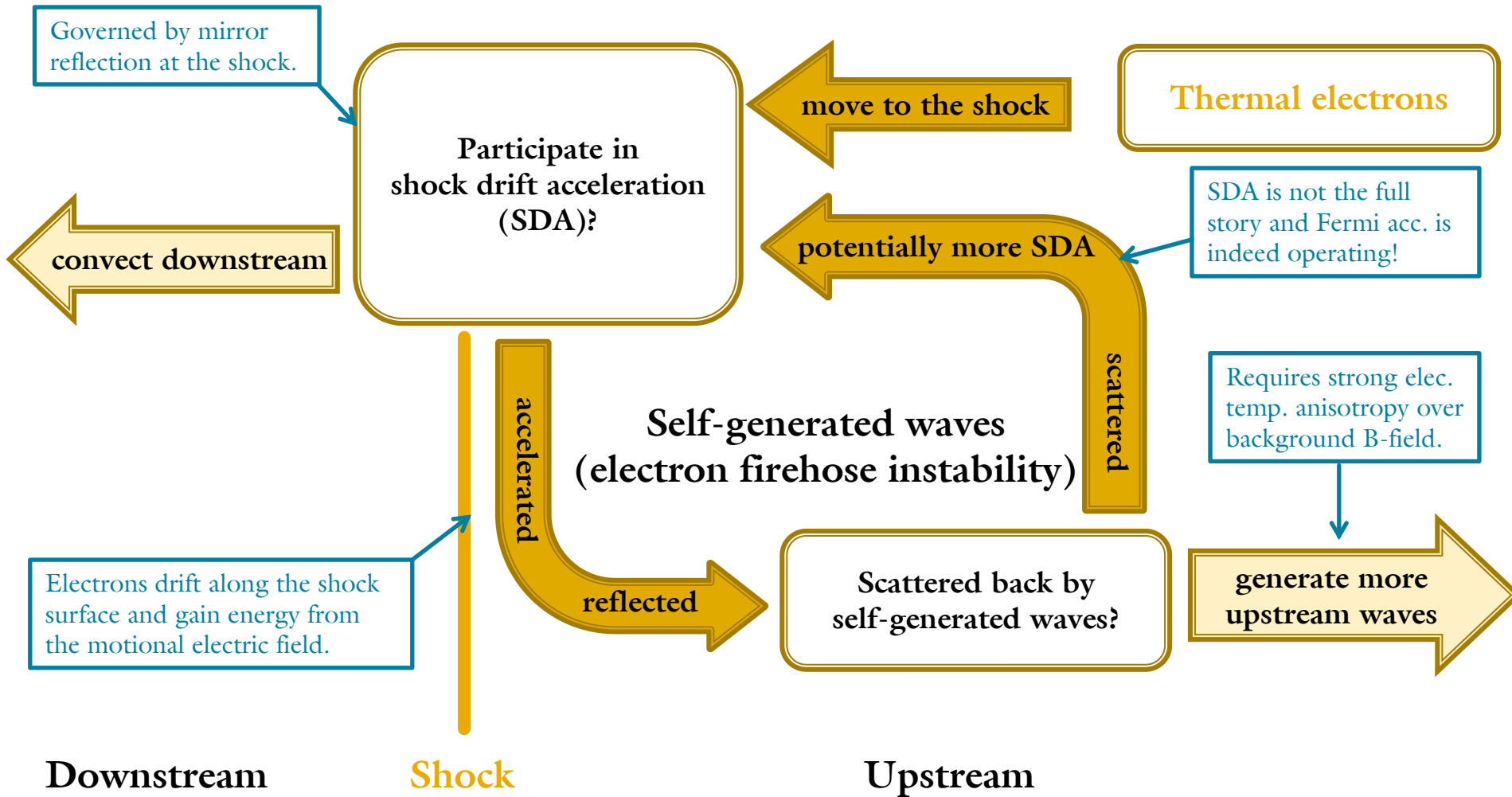
Threshold parameter

$$\Lambda \equiv 1 - \frac{T_{e\perp}}{T_{e\parallel}} - 1.3 \frac{B_0^2 / 8\pi}{n_e k T_{e\parallel}} > 0$$

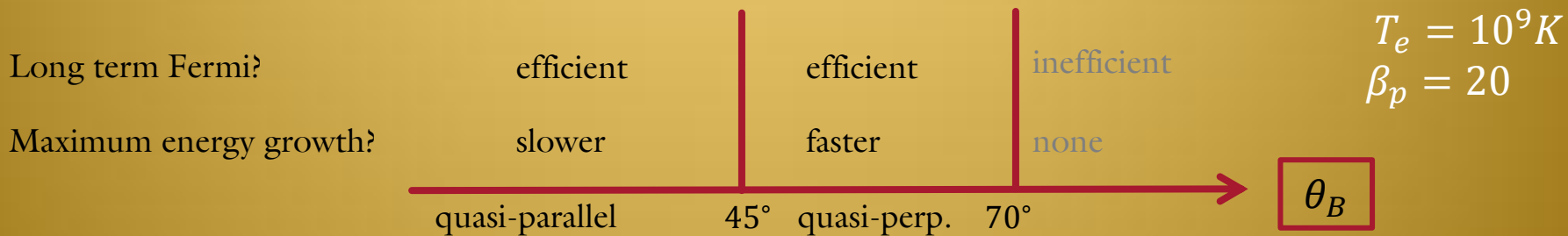


Gary & Nishimura 2003
Camporeale & Burgess 2008

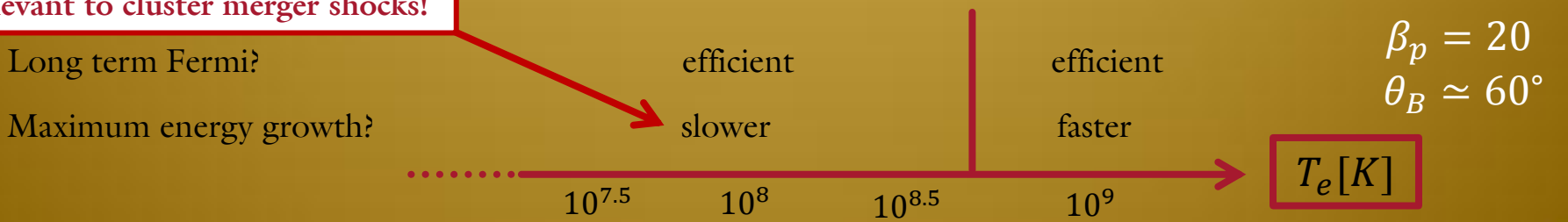
The Big Picture



Parameter Dependence

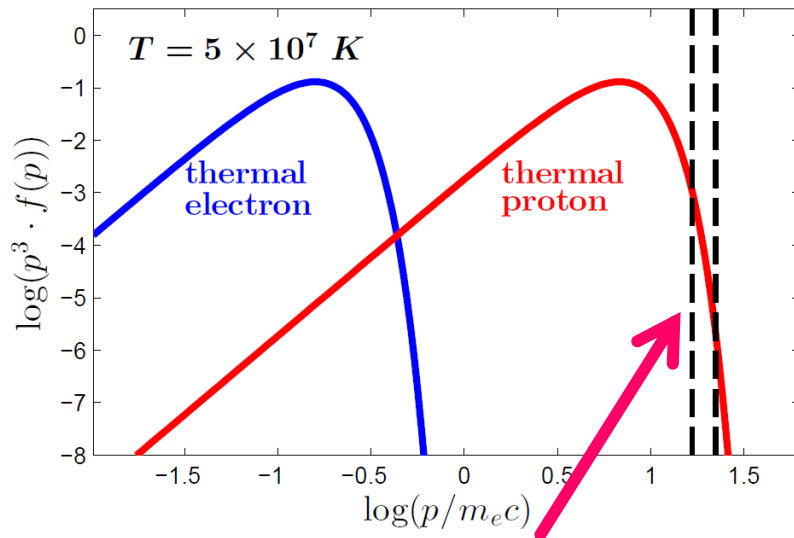


high plasma beta, low temperature
low Mach number shocks are
relevant to cluster merger shocks!



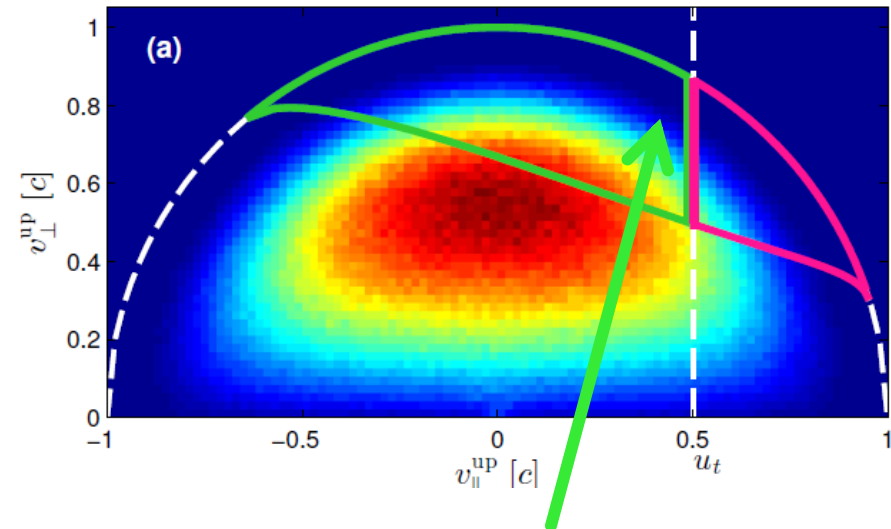
Electron injection problem

direct injection via thermal leakage



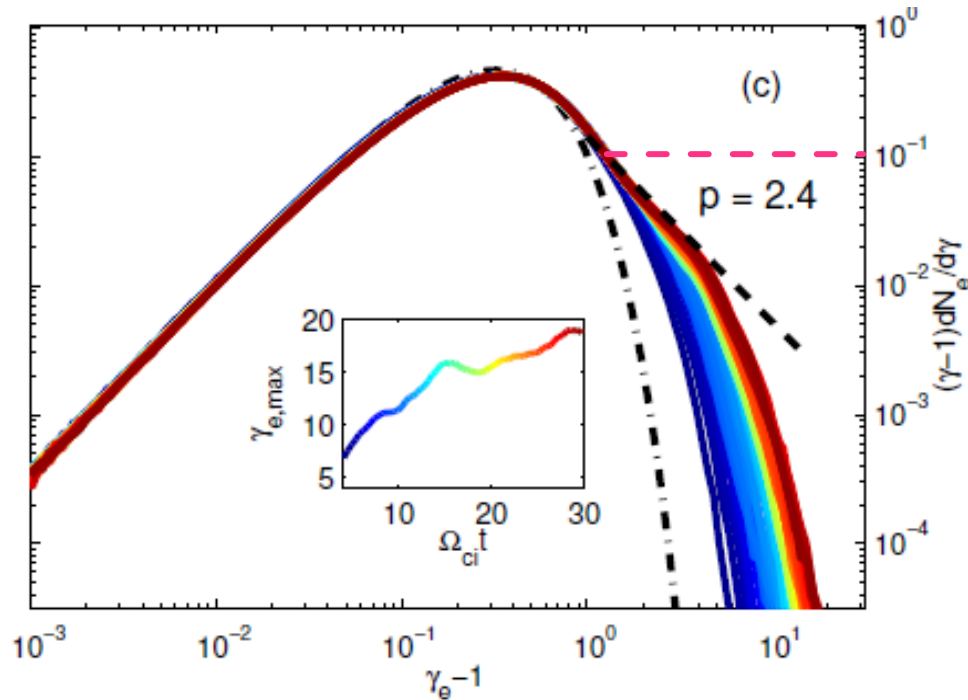
- Relies on scattering by downstream waves.
- Needs $p_e > p_{th,p}$.
- Electron acceleration efficiency extremely poor.

injection into SDA

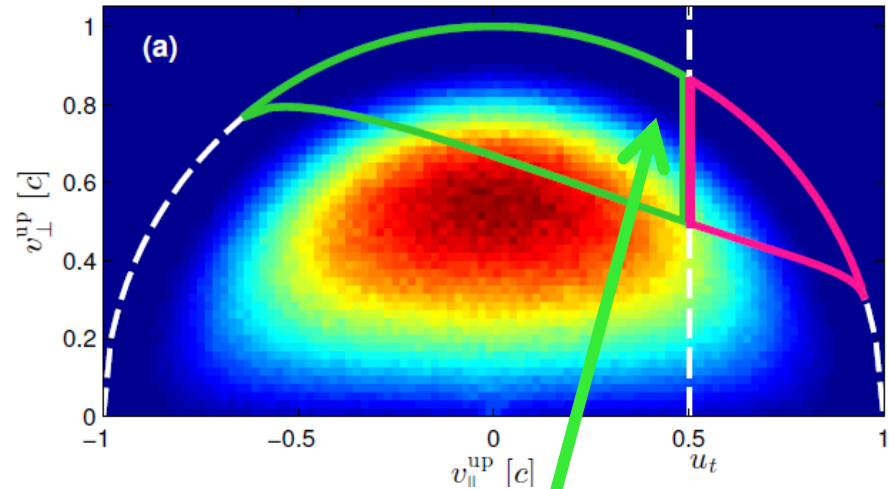


- Shock itself reflects electrons.
- No need for $p_e > p_{th,p}$.
- Thermal electrons can be easily accelerated.

Electron injection problem

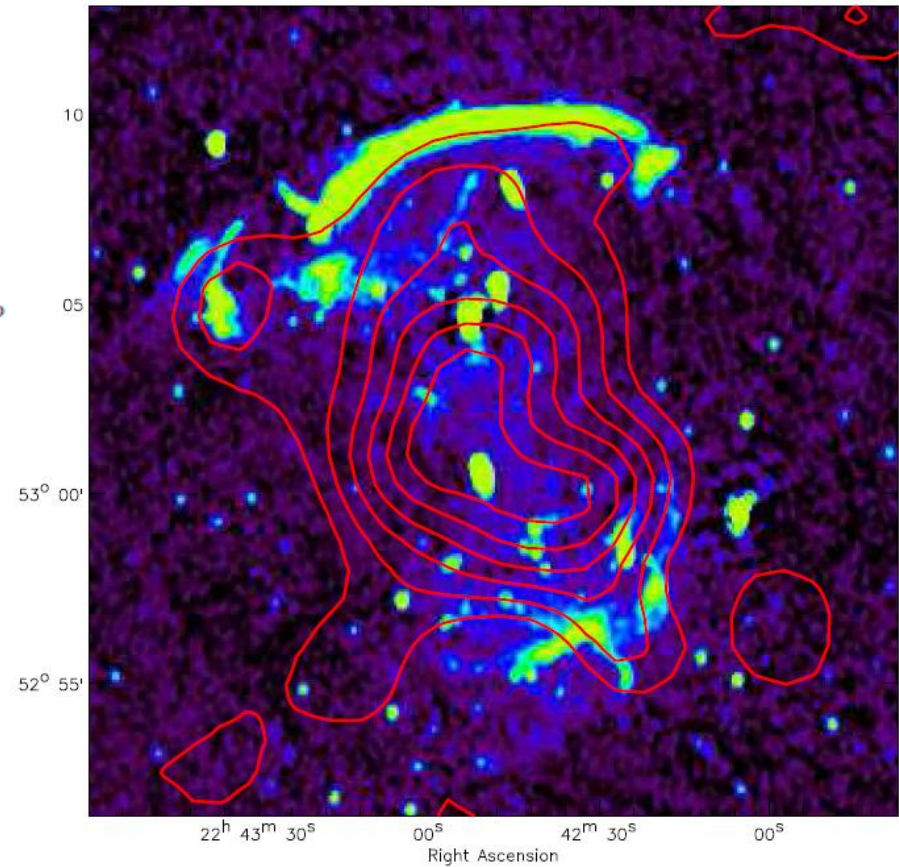
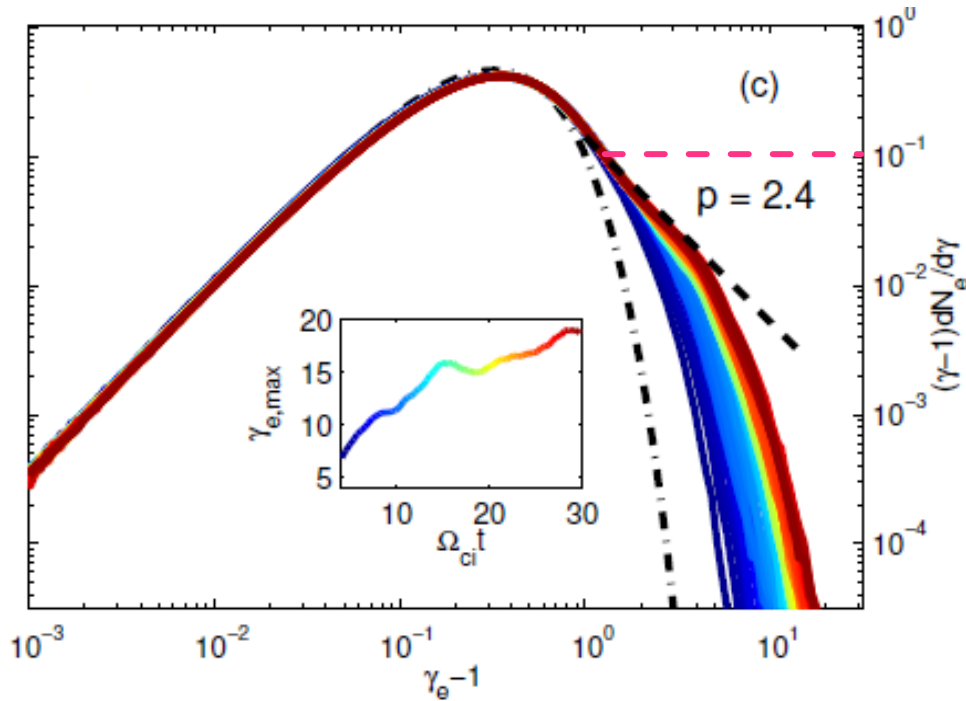


Injection into SDA



- Shock itself reflects electrons.
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- Thermal electrons can be easily accelerated.

Electron injection problem



no more tension with observations!

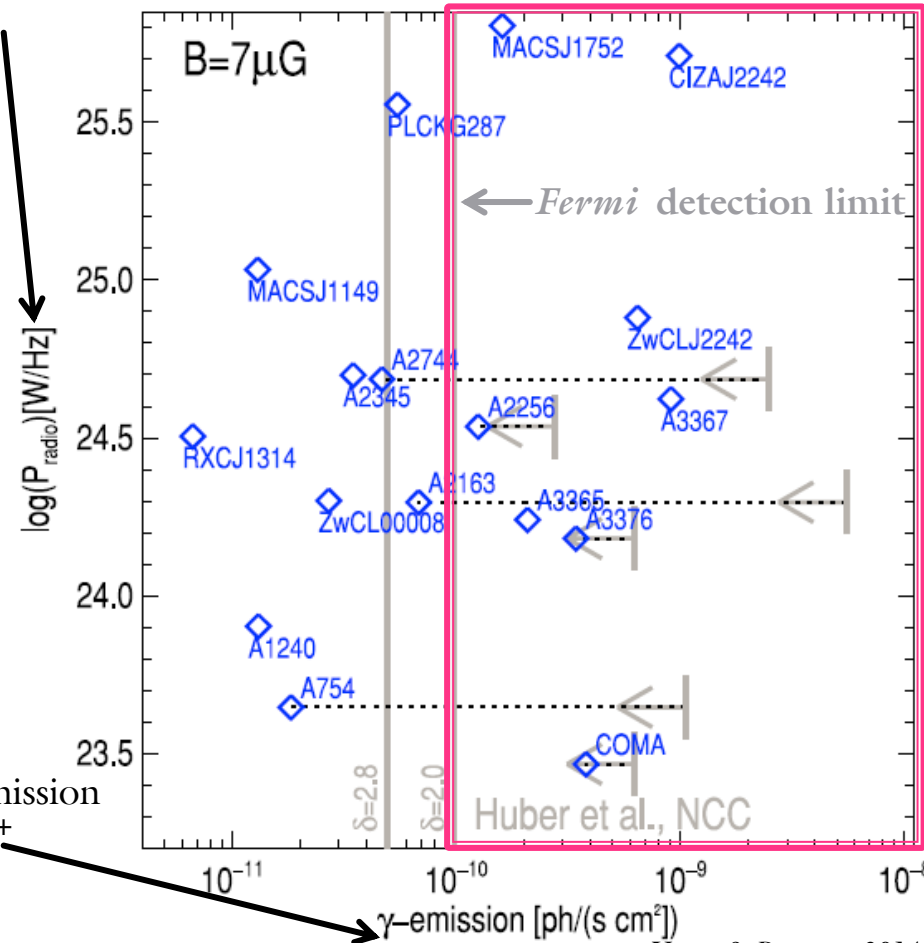
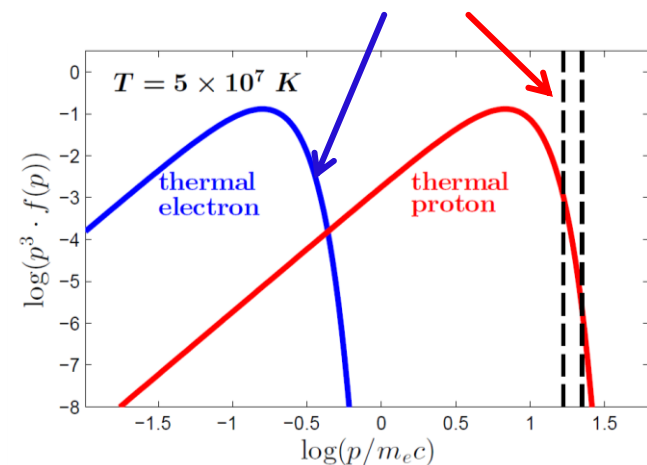
Missing γ -ray from accelerated ions?

CRp/CRe ratio \ll previously thought

thermal leakage model:
a lot more p^+ than e^-
could be accelerated

observed radio emission
from accelerated e^-

Above *Fermi* limit, yet no detection!

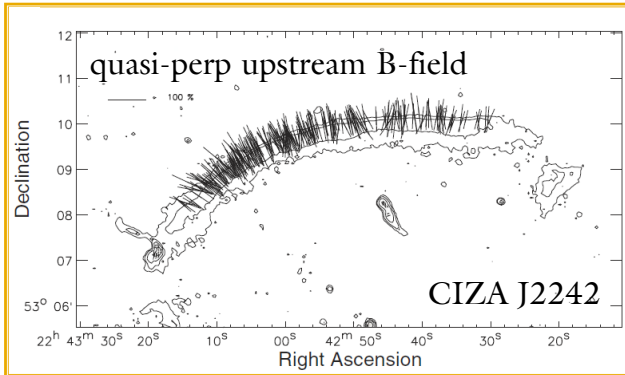


estimated γ -ray emission
from accelerated p^+

Vazza & Brüggén 2014

Missing γ -ray from accelerated ions?

CRp/CRE ratio \ll previously thought

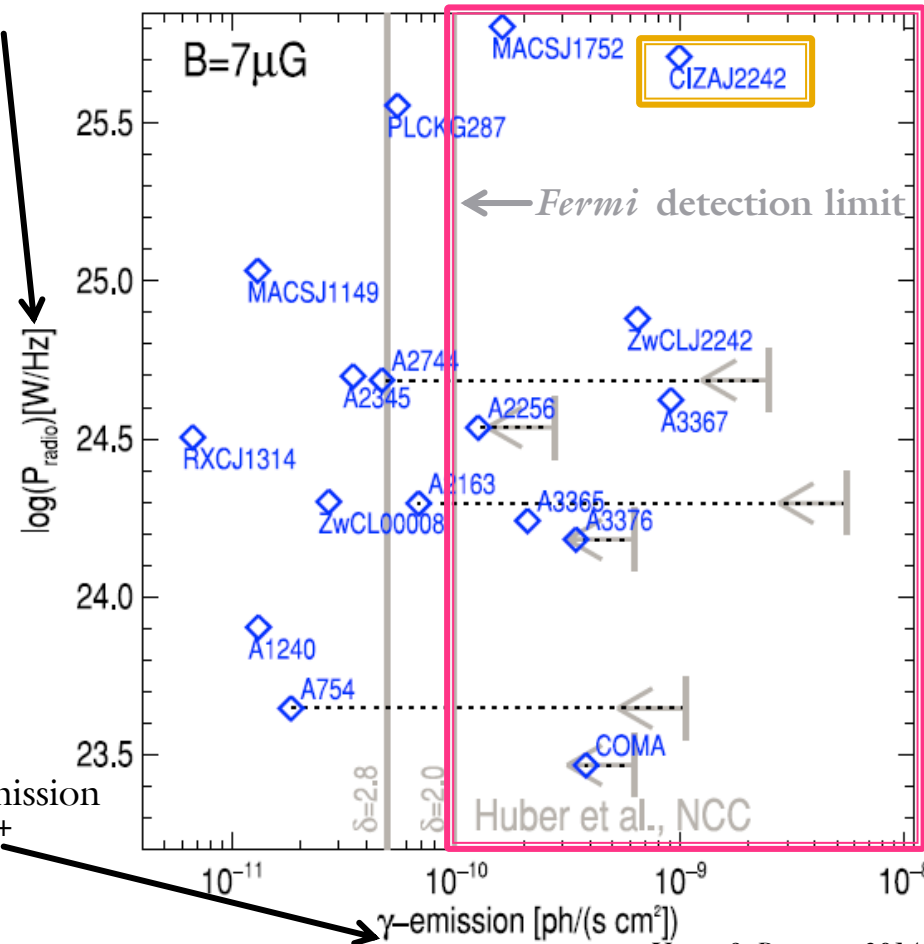


observed radio emission from accelerated e^-

- We find electrons are efficiently accelerated over a wide range of obliquities.
- Ion acceleration is poor in quasi-perp shocks. (Caprioli & Spitkovsky 2014)
- The ratio of ion- to electron-acceleration efficiency is much smaller than previously thought.

estimated γ -ray emission from accelerated p^+

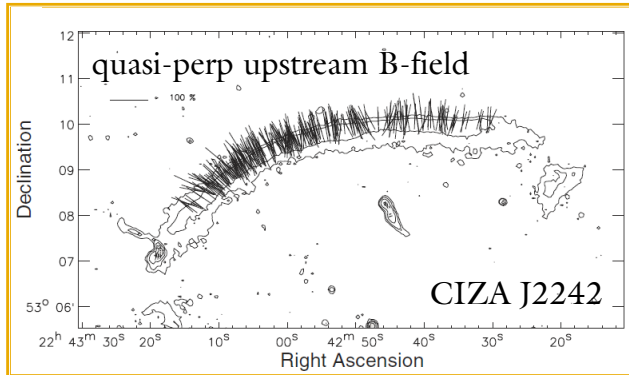
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Vazza & Brüggen 2014

Missing γ -ray from accelerated ions?

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- Ion acceleration is poor in quasi-perp shocks. (Caprioli & Spitkovsky 2014)
- The ratio of ion- to electron-acceleration efficiency is much smaller than previously thought.

The expected γ -ray emission from ions is actually lower, consistent with null detections of *Fermi*.

Summary

- Electron acceleration to non-thermal energies at low Mach number shocks are revealed through radio relics, but the mechanism has been poorly understood.
- We study electron acceleration in low Mach number shocks from first principles using the PIC method.
- We identify a Fermi-like electron acceleration mechanism.
- The electrons are efficiently injected into shock drift acceleration (SDA) by mirror reflection at the shock front, they gain energy by multiple cycles of SDA mediated by scattering with self-generated upstream waves associated with oblique electron firehose instability.
- We find that the mechanism can operate over a large range of parameter space, especially relevant for cluster merger shocks.
- Our mechanism naturally explains the bright radio emission from radio relics, and alleviates tension with null detection of gamma-ray from merger shocks through *Fermi*.

2014arXiv1406.5190

2014arXiv1409.7393