
Nonlinear Gyrokinetic Description of Alfvénic Turbulence

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Outline

- Introduction
- Gyrokinetics and Polarization Density
- Alfvén Eigenmodes in Tokamaks
- Weak Turbulence Theory of Toroidal Alfvén Eigenmodes
- Relations to MHD Turbulence

Basics of Shear- Alfvén Waves

$$\delta E_{\parallel} = 0 \quad \Rightarrow \quad \frac{1}{c} \partial_t \delta A_{\parallel} = -\nabla_{\parallel} \delta \phi$$

ideal MHD

From $\vec{\nabla} \cdot \vec{\delta J} = 0 \quad \Rightarrow \quad \frac{m_i n_0 c}{B^2} \partial_t \underline{\underline{\nabla_{\perp}^2 \delta \phi}} = \nabla_{\parallel} J_{\parallel} = \nabla_{\parallel} \nabla_{\perp}^2 \delta A_{\parallel}$

$$v_A^{-2} \omega^2 = k_{\parallel}^2$$

MHD : from $\vec{\nabla} \times$ of $\rho \frac{d}{dt} \vec{v} = \vec{J} \times \vec{B}$

i.e. “ Vorticity ”

Gyrokinetics : “ Polarization Density ”

$$n_{\text{pol}} = \nabla_{\perp} \cdot \left(\frac{m_i n_0 c}{B^2} \nabla_{\perp} \delta \phi \right)$$

Gyrokinetics and Polarization Density

Ref. Hahm-Lee-Brizard, Phys Fluids '88

- Vorticity Equation can be derived from

$$\frac{dn_e}{dt} = \frac{1}{|e|} \nabla_{\parallel} j_{\parallel e}$$

$$\frac{d}{dt} N_{\text{gyrocenter}} = -\frac{1}{|e|} \nabla_{\parallel} j_{\parallel i}$$

← From moments of Gyrokinetic Equation

$$\frac{d}{dt} "N_{\text{pol}}" = \frac{1}{|e|} \nabla_{\parallel} J_{\parallel}$$

- What is Physical Meaning of Polarization Density?
- What is Gyrokinetics?

Conventional (old-fashioned) Derivation of Non-linear Gyrokinetic Equation

- Closely follow Guiding Center transformation by
P.J. Catto, Plasma Phys. **20**, 719 (1977)
- Resulting equation
Frieman and Chen, PF **25**, 502 (1982)
Lee, PF **26** 556 (1983)
- Purpose: illustrate basic physics and
mathematical complexity
involved in this conventional method.

Consider uniform $\mathbf{B} = B\hat{\mathbf{b}}$ to emphasize nonlinear effects

- Goal: from

$$\left[\frac{\partial}{\partial t} + \mathbf{v} \cdot \frac{\partial}{\partial \mathbf{x}} + \frac{q}{m} \left(\mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{B} \right) \cdot \frac{\partial}{\partial \mathbf{v}} \right] f(\mathbf{x}, \mathbf{v}, t) = 0 \quad \text{6D Vlasov Eqn}$$

get

$$\left(\frac{\partial}{\partial t} + \frac{d\mathbf{R}}{dt} \cdot \frac{\partial}{\partial \mathbf{R}} + \frac{dv_{\parallel}}{dt} \frac{\partial}{\partial v_{\parallel}} \right) \langle f \rangle(\mathbf{R}, \mu, v_{\parallel}, t) = 0 \quad \text{5D GK Eqn}$$

with

$$\frac{d\mu}{dt} = 0 \quad \text{and} \quad \frac{\partial}{\partial \theta} \langle f \rangle = 0$$

$\mu \simeq v_{\perp}^2 / (2B)$: magnetic moment, an adiabatic invariant at lowest order

- Assumption:

- $\omega \ll \Omega_{ci}$
- $k_{\parallel} \ll k_{\perp} \sim \rho_i^{-1}$
- $\delta f / f_0 \sim \delta n / n_0 \sim e\delta\phi / T_e \ll 1$

Guiding Center Transformation à la Catto

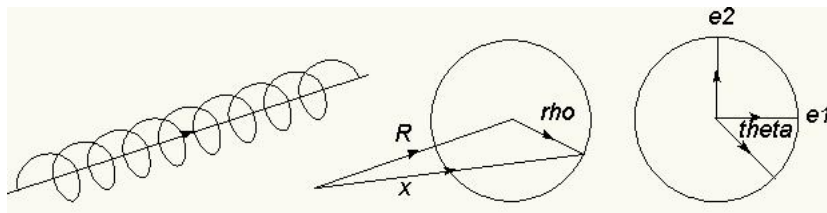
$(\mathbf{x}, \mathbf{v}) \rightarrow (\mathbf{R}, v_{\parallel}, \mu, \theta)$, θ : gyrophase-angle

$$\mathbf{R} = \mathbf{x} - \boldsymbol{\rho}, \boldsymbol{\rho} = \frac{\hat{\mathbf{b}} \times \mathbf{v}}{\Omega}, \Omega = \frac{eB}{mc}$$

$$v_{\parallel} = \hat{\mathbf{b}} \cdot \mathbf{v}, \mu = v_{\perp}^2 / (2B)$$

θ defined by

$$\begin{cases} \mathbf{v} = v_{\parallel} \hat{\mathbf{b}} + v_{\perp} \hat{\mathbf{e}}_{\perp} \\ \hat{\mathbf{e}}_{\perp} = -\hat{\mathbf{e}}_2 \cos \theta - \hat{\mathbf{e}}_1 \sin \theta \\ \hat{\mathbf{e}}_{\rho} = \hat{\mathbf{e}}_1 \cos \theta - \hat{\mathbf{e}}_2 \sin \theta \end{cases}$$



Note that for uniform \mathbf{B} ,

$$d^3 \mathbf{x} d^3 \mathbf{v} = \underbrace{B}_{\text{phase-space volume}} d\mu d\theta dv_{\parallel} d^3 \mathbf{R}$$

B : “phase-space volume”

Then, we would like to express $\frac{\partial}{\partial \mathbf{x}}$ and $\frac{\partial}{\partial \mathbf{v}}$ in G.C. space i.e., in terms of μ , v_{\parallel} , \mathbf{R} , and θ ;

$$\frac{\partial}{\partial \mathbf{x}} = \frac{\partial \mathbf{R}}{\partial \mathbf{x}} \cdot \frac{\partial}{\partial \mathbf{R}} + \frac{\partial \mu}{\partial \mathbf{x}} \frac{\partial}{\partial \mu} + \frac{\partial v_{\parallel}}{\partial \mathbf{x}} \frac{\partial}{\partial v_{\parallel}} + \frac{\partial \theta}{\partial \mathbf{x}} \cdot \frac{\partial}{\partial \theta}$$

$$\frac{\partial}{\partial \mathbf{v}} = \frac{\partial \mathbf{R}}{\partial \mathbf{v}} \cdot \frac{\partial}{\partial \mathbf{R}} + \frac{\partial \mu}{\partial \mathbf{v}} \frac{\partial}{\partial \mu} + \frac{\partial v_{\parallel}}{\partial \mathbf{v}} \frac{\partial}{\partial v_{\parallel}} + \frac{\partial \theta}{\partial \mathbf{v}} \cdot \frac{\partial}{\partial \theta}$$

→ important to check what quantities are held constant when taking partial derivatives

Since

$$\left. \frac{\partial}{\partial \mathbf{x}} \mu \right|_{\mathbf{v}=\text{const}} = 0, \quad \left. \frac{\partial}{\partial \mathbf{x}} v_{\parallel} \right|_{\mathbf{v}=\text{const}} = 0, \quad \left. \frac{\partial}{\partial \mathbf{x}} \right|_{\mathbf{v}=\text{const}} \theta = 0, \quad \text{and } \mathbf{R} = \mathbf{x} - \frac{\hat{\mathbf{b}} \times \mathbf{v}}{\Omega}$$

$\frac{\partial}{\partial \mathbf{x}}$ → only the 1st term on the R.H.S. survives ⇒

$$\frac{\partial}{\partial \mathbf{x}} = \mathbf{I} \cdot \frac{\partial}{\partial \mathbf{R}} = \frac{\partial}{\partial \mathbf{R}}$$

Also, noting that

$$\left. \frac{\partial}{\partial \mathbf{v}} \right|_{\mathbf{x}=\text{const}} v_{\parallel} = \left. \frac{\partial}{\partial \mathbf{v}} \right|_{\mathbf{x}=\text{const}} \mathbf{v} \cdot \hat{\mathbf{b}} = \hat{\mathbf{b}}, \quad \frac{\partial}{\partial \mathbf{v}} \mu = \mathbf{v}_{\perp} / B$$

$$\frac{\partial}{\partial \mathbf{v}} \mathbf{R} = \frac{\partial}{\partial \mathbf{v}} \left(\mathbf{x} - \frac{\hat{\mathbf{b}} \times \mathbf{v}}{\Omega} \right) \rightarrow - \frac{\partial}{\partial \mathbf{v}} \left(\frac{\hat{\mathbf{b}} \times \mathbf{v}}{\Omega} \right) = \frac{\mathbf{I} \times \hat{\mathbf{b}}}{\Omega}$$

$$\frac{\partial}{\partial \mathbf{v}} = \hat{\mathbf{b}} \frac{\partial}{\partial v_{\parallel}} + \frac{\mathbf{v}_{\perp}}{B} \frac{\partial}{\partial \mu} - \frac{\hat{\mathbf{b}} \times \hat{\mathbf{e}}_{\perp}}{v_{\perp}} \frac{\partial}{\partial \theta} + \frac{\mathbf{I} \times \hat{\mathbf{b}}}{\Omega} \frac{\partial}{\partial \mathbf{R}}$$

\Rightarrow

$$\mathbf{v} \cdot \frac{\partial}{\partial \mathbf{x}} = v_{\parallel} \hat{\mathbf{b}} \cdot \frac{\partial}{\partial \mathbf{R}} + \mathbf{v}_{\perp} \cdot \frac{\partial}{\partial \mathbf{R}} \quad (1)$$

$$\frac{q}{m} \mathbf{E} \cdot \frac{\partial}{\partial \mathbf{v}} = \frac{q}{m} \left(E_{\parallel} \cdot \frac{\partial}{\partial v_{\parallel}} + \frac{\mathbf{E} \cdot \mathbf{v}_{\perp}}{B} \frac{\partial}{\partial \mu} - \frac{\mathbf{E} \cdot \hat{\mathbf{b}} \times \mathbf{v}_{\perp}}{v_{\perp}^2} \frac{\partial}{\partial \theta} \right) + \frac{c \mathbf{E} \times \mathbf{B}}{B^2} \cdot \frac{\partial}{\partial \mathbf{R}} \quad (2)$$

$$\begin{aligned} \frac{q \mathbf{v} \times \mathbf{B}}{mc} \cdot \frac{\partial}{\partial \mathbf{v}} &= 0 + 0 - \Omega \frac{\mathbf{v} \times \mathbf{B} \cdot \mathbf{B} \times \mathbf{v}_{\perp}}{B^2 v_{\perp}^2} \frac{\partial}{\partial \theta} + \Omega \frac{(\mathbf{v} \times \hat{\mathbf{b}}) \times \hat{\mathbf{b}}}{\Omega} \cdot \frac{\partial}{\partial \mathbf{R}} \\ &= \Omega \frac{\partial}{\partial \theta} - \mathbf{v}_{\perp} \cdot \frac{\partial}{\partial \mathbf{R}} \end{aligned} \quad (3)$$

We also want to express $\phi(\mathbf{x})$ and $\mathbf{E}(\mathbf{x})$ in terms of $(\mathbf{R}, \mu, \mathbf{v}_{\parallel}, \theta)$

$$\phi(\mathbf{x}) = \phi(\mathbf{R} + \boldsymbol{\rho}(\theta)) \Rightarrow$$

$$\frac{\partial \phi}{\partial \theta} = \frac{\partial \mathbf{x}}{\partial \theta} \Big|_{\mathbf{R}} \cdot \frac{\partial \phi}{\partial \mathbf{x}} = \frac{\partial \boldsymbol{\rho}}{\partial \theta} \cdot \frac{\partial \phi}{\partial \mathbf{x}} = \frac{\mathbf{v}_{\perp}}{\Omega} \cdot \frac{\partial \phi}{\partial \mathbf{x}} = -\frac{\mathbf{E} \cdot \mathbf{v}_{\perp}}{\Omega}$$

\therefore the 2nd term of RHS of Eq. (4.2)

$$\frac{q}{m} \frac{\mathbf{E} \cdot \mathbf{v}_{\perp}}{B} \frac{\partial}{\partial \mu} = -\frac{1}{c} \left(\frac{q}{m}\right)^2 \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu}$$

Collecting all terms in Eqs. (1)-(3),

$$\left[\frac{\partial}{\partial t} + v_{\parallel} \hat{\mathbf{b}} \cdot \frac{\partial}{\partial \mathbf{R}} + c \frac{\mathbf{E} \times \mathbf{B}}{B^2} \cdot \frac{\partial}{\partial \mathbf{R}} - \frac{q}{m} \nabla_{\parallel} \phi \frac{\partial}{\partial v_{\parallel}} + \Omega \frac{\partial}{\partial \theta} - \frac{q\Omega}{mB} \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu} - \Omega \frac{\mathbf{v}_E \cdot \mathbf{v}_{\perp}}{v_{\perp}^2} \frac{\partial}{\partial \theta} \right] f = 0 \quad (4)$$

$$-i\omega \quad ik_{\parallel} v_{\parallel} \quad \mathbf{k}_{\perp} \cdot \mathbf{v}_E \quad k_{\parallel} v_{\parallel} \left(\frac{e\phi}{T_e} \right) \quad \Omega \quad \underbrace{(i) \quad (ii)}_{\text{ugly!}}$$

- Term (i) can be shown to be the 1st order correction to μ i.e.,

$$\frac{d\mu}{dt} = \frac{d\mu^{(0)}}{dt} + \frac{d\mu^{(1)}}{dt} \Rightarrow \frac{d}{dt} \left(\frac{v_{\perp}^2}{2B} \right)^{(1)} = \frac{\mathbf{v}_{\perp}^{(0)}}{B} \cdot \frac{d}{dt} \mathbf{v}_{\perp}^{(1)}(\theta)$$

where

$$\frac{d}{dt} \mathbf{v}_{\perp}^{(1)} = \frac{q}{m} (\mathbf{v}_{\perp}^{(1)} \times \mathbf{B} + \mathbf{E}^{(1)}) \Rightarrow \mathbf{v}_{\perp}^{(0)} \cdot \frac{d}{dt} \mathbf{v}_{\perp}^{(1)} = \frac{q}{m} \mathbf{E}_{\perp}^{(1)} \cdot \mathbf{v}_{\perp}^{(0)}$$

- Term (ii) similarly, 1st order correction to the gyrophase θ , i.e., gyration speed is slightly nonuniform due to $\mathbf{E}_{\perp}^{(1)}$,
 → Not of primary physical interest
- Now, we perform perturbation theory:

with

$$\Omega \gg \omega \sim k_{\parallel} v_{\parallel}, \quad \frac{\omega}{\Omega} \sim \frac{e\delta\phi}{T} \ll 1, \quad k_{\parallel} \ll k_{\perp} \sim \rho_i^{-1}$$

- Eq. (4) \Rightarrow

$$\underbrace{\Omega \frac{\partial f}{\partial \theta}}_{\text{Largest term}} + \left(\frac{\partial}{\partial t} + v_{\parallel} \hat{\mathbf{b}} \cdot \frac{\partial}{\partial \mathbf{R}} + c \frac{\mathbf{E} \times \mathbf{B}}{B^2} \cdot \frac{\partial}{\partial \mathbf{R}} - \frac{q}{m} \nabla_{\parallel} \phi \frac{\partial}{\partial v_{\parallel}} - \frac{q\Omega}{mB} \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu} \right) f = 0 \quad (5)$$

Let $f = f^{(0)} + f^{(1)} + \dots$, with expansion parameter $\delta \sim \frac{\omega}{\Omega} \sim \frac{k_{\parallel} v_{\parallel}}{\Omega} \sim \frac{|e|\phi}{T_e}$

- 0-th order $\Rightarrow \Omega \frac{\partial}{\partial \theta} f^{(0)} = 0 \Rightarrow f^{(0)}$ is independent of θ ,
 $\therefore f = \langle f \rangle + f_{AC}$, $\langle \dots \rangle = \frac{1}{2\pi} \oint d\theta \{ \dots \}$ gyrophase average
with $f^{(0)} = \langle f \rangle$, $f^{(1)} = f_{AC} \ll f^{(0)} = \langle f \rangle$
- 1-st order \Rightarrow

$$\underbrace{\Omega \frac{\partial}{\partial \theta} f^{(1)}}_{(a)} + \left(\frac{\partial}{\partial t} + v_{\parallel} \hat{\mathbf{b}} \cdot \frac{\partial}{\partial \mathbf{R}} + c \frac{\mathbf{E} \times \mathbf{B}}{B^2} \cdot \frac{\partial}{\partial \mathbf{R}} - \frac{q}{m} \nabla_{\parallel} \phi \frac{\partial}{\partial v_{\parallel}} - \underbrace{\frac{q\Omega}{mB} \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu}}_{(b)} \right) f^{(0)} = 0 \quad (6)$$

(a) and (b) can be combined into

$$\Omega \frac{\partial}{\partial \theta} \left[f_{AC} - \frac{q\phi}{mB} \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu} \langle f \rangle \right]$$

- Taking gyro-phase average of Eq. (6): $\langle \dots \rangle = \frac{1}{2\pi} \oint d\theta \dots$

$$\langle \Omega \frac{\partial}{\partial \theta} \{ \dots \} \rangle = 0 \Rightarrow$$

$$\left[\frac{\partial}{\partial t} + v_{\parallel} \hat{\mathbf{b}} \cdot \frac{\partial}{\partial \mathbf{R}} + \frac{c}{B} \hat{\mathbf{b}} \times \nabla \langle \phi \rangle - \frac{q}{m} \hat{\mathbf{b}} \cdot \frac{\partial}{\partial \mathbf{R}} \langle \phi \rangle \frac{\partial}{\partial v_{\parallel}} \right] \langle f \rangle = 0 \quad (7)$$

Finally, the electrostatic NLGK vlasov equation in uniform \mathbf{B}

- $\langle \phi \rangle$ contains the Finite Larmor Radius (FLR) effect!

although it's gyrophase-averaged

$$\phi(\mathbf{x}) = \phi(\mathbf{R} + \boldsymbol{\rho}) = \sum_{\mathbf{k}} \phi_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{x}} = \sum_{\mathbf{k}} \phi_{\mathbf{k}} e^{i\mathbf{k}_{\perp} \cdot \mathbf{R}} e^{ik_{\perp} \rho \sin \theta}$$

Fourier-Bessel Expansion:

$$e^{ik_{\perp} \rho \sin \theta} = \sum J_n(k_{\perp} \rho) e^{in\theta}$$

$$\langle e^{ik_{\perp} \rho \sin \theta} \rangle = \frac{1}{2\pi} \oint d\theta \sum_n^n J_n(k_{\perp} \rho) e^{in\theta} = J_0(k_{\perp} \rho)$$

∴

$$\langle \phi \rangle = \sum_{\mathbf{k}} J_0(k_{\perp} \rho) \phi_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{R}}$$

- Widespread Misconception: “Gyrokinetic Theory throws away the gyrophase-dependent information”
- Part of Reasons: Conventional (old-fashioned) derivation is rather opaque (much more complex in general geometry in nonuniform \mathbf{B})
Illustration in this note is a bit “modernized” version than the original papers up to mid 80’s.
 - Hard to identify the role or necessity of θ –dependent information
 - Also, most attention was paid to the nonlinear GK-“Vlasov” Equations.

Gyrokinetic Poisson Equation

- Maxwell's Eqns are still fine!
but was NOT written in g.c. coordinates (\mathbf{R})
- So we need to express $n_i(\mathbf{x})$ in terms of $\langle f \rangle(\mathbf{R}, \mathbf{v}_{\parallel}, \mu)$

$$(\mathbf{R}, \mathbf{v}_{\parallel}, \mu, \theta) \Rightarrow (\mathbf{x}, \mathbf{v})$$

“Pull-Back” Transformation for GK Maxwell's Eqn
(ES \Rightarrow Poisson)

$$(\mathbf{x}, \mathbf{v}) \Rightarrow (\mathbf{R}, \mathbf{v}_{\parallel}, \mu, \theta)$$

“Push-Forward” Transformation for GK-Vlasov

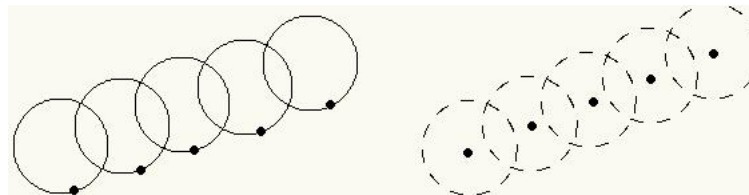
$$\nabla^2 \phi = -4\pi e [n_i(\mathbf{x}) - n_e(\mathbf{x})]$$

- $n_i(\mathbf{x})$: typically obtained from GK Eqn
- $n_e(\mathbf{x})$: from adiabatic response for pure - ITG
or from drift-kinetic or bounce-kinetic
or from some other fluid eqns for more realistic case
“GK” required for ETG

$$\begin{aligned} n_i(\mathbf{x}) &= \int d^3\mathbf{v} f_i(\mathbf{x}, \mathbf{v}, t) \\ &= \int d^3\mathbf{x}' d^3\mathbf{v} f_i(\mathbf{x}', \mathbf{v}) \delta(\mathbf{x}' - \mathbf{x}) \\ &= \int d^3\mathbf{R} d\mu dv_{\parallel} d\theta B f_i(\mathbf{R}, \mu, v_{\parallel}, \theta) \delta(\mathbf{R} + \boldsymbol{\rho} - \mathbf{x}) \end{aligned} \quad (8)$$

not quite the same

$$\int d^3\mathbf{R} d\mu dv_{\parallel} B \langle f \rangle (\mathbf{R}, \mu, v_{\parallel})$$



Since

$$f_i(\mathbf{R}, \mu, v_{\parallel}, \theta) = \langle f \rangle + f_{AC}(\mathbf{R}, \mu, v_{\parallel}, \theta),$$

we need to know “ f_{AC} ” as well.

Back to Eq. (6):

$$\Omega \frac{\partial}{\partial \theta} \left[f_{AC} - \frac{q\phi}{mB} \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu} \langle f \rangle \right] + \left(\frac{d}{dt} \langle f \rangle \right) = 0$$

and Eq. (7)

$$\left. \frac{d}{dt} \right|^{(0)} \langle f \rangle = 0$$

\Rightarrow

$$\Omega \frac{\partial}{\partial \theta} \left[f_{AC} - \frac{q\phi}{mB} \frac{\partial \phi}{\partial \theta} \frac{\partial}{\partial \mu} \langle f \rangle \right] + \left(\frac{d}{dt} - \left. \frac{d}{dt} \right|^{(0)} \right) \langle f \rangle = 0 \quad (9)$$

$$\frac{d}{dt} - \left. \frac{d}{dt} \right|^{(0)} \propto \left(\phi - \langle \phi \rangle \right)$$

integrating Eq. (9)

$$f_{AC}(\theta) \simeq \frac{q}{mB} (\phi - \langle \phi \rangle) \frac{\partial}{\partial \mu} \langle f \rangle \quad (10)$$

Polarization Density

Eq. (8) \Rightarrow

$$n_i(\mathbf{x}) = \underbrace{\int d^3\mathbf{R} d\mu dv_{\parallel} d\theta B \langle f \rangle \delta(\mathbf{R} + \boldsymbol{\rho} - \mathbf{x})}_{n_{i,gc}(\mathbf{x})} + \underbrace{\int d^3\mathbf{R} d\mu dv_{\parallel} d\theta B \frac{q}{mB} (\phi - \langle \phi \rangle) \frac{\partial \langle f \rangle}{\partial \mu} \delta(\mathbf{R} + \boldsymbol{\rho} - \mathbf{x})}_{n_{pol}(\mathbf{x})}$$

- $n_{i,gc}(\mathbf{x})$: G.C. density at particle position
- $n_{pol}(\mathbf{x})$: Polarization Density, one can evaluate exactly for $\langle f \rangle \propto e^{-\mu B/T}$,
i.e., “Maxwellian in $\mu \propto v_{\perp}^2$ ”

$$\rightarrow (1 - \Gamma_0) \delta\phi$$

$$\rightarrow \text{For long wavelength ; } \propto \vec{\nabla}_{\perp} \cdot \left(\frac{n}{B^2} \vec{\nabla}_{\perp} \delta\phi \right)$$

Nonlinear Gyrokinetics for Large Scale Computation

- Direct simulation of actual size fusion plasmas in realistic geometry using the primitive nonlinear plasma equations (Vlasov-Maxwell), is far beyond the computational capability of foreseeable future.
- For turbulence problems in fusion plasmas, the temporal scales fluctuations much longer than the period of a charged particle's cyclotron motion, while the spatial scales and gyro-orbits are much smaller than the macroscopic length scales: → details of the charged particle's gyration motion are **not** of physical interest → Develop reduced dynamical equations which capture the essential features
- After decoupling of gyro-motion, gyrokinetic equation describes evolution of gyro-center distribution function, independent of the gyro-phase, θ , defined over a five-dimensional phase space (\mathbf{R} , v_{\parallel} , μ). → save enormous amounts of computing time by having a time step greater than the gyro-period, and by reducing the number of dynamical variables.
- In gyrokinetic approach, gyro-phase is an ignorable coordinate, magnitude of the perpendicular velocity enters as a parameter in terms of an adiabatic invariant μ
- Nonlinear gyrokinetic equations are now widely used in turbulence simulations.⁶

Modern Nonlinear Gyrokinetics

- Starting from the original Vlasov-Maxwell system (6D), pursue “**Reduction of dimensionality**” for both computational and analytic feasibility.
- Keep intact the underlying symmetry/conservation of the original system.
- Perturbation analysis consists of near-identity coordinate transformation which “**decouples**” the gyration from the slower dynamics of interest in the single particle Lagrangian, rather than a direct “gyro-phase average” of Vlasov equation.
- This procedure is **reversible**:
The gyro-phase dependent information can be recovered when it is needed.

Phase Space Lagrangian Derivation of Nonlinear Gyrokinetics

[since Hahm, PF **31**, 2670 '88, followed by Brizard, Sugama,...]

- **Conservations Laws are Satisfied.**
- Various expansion parameters appear at different stages
→ Flexibility in variations of ordering for specific application
- Guiding center drift calculations in equilibrium field **B**:
Expansion in $\delta_B = \rho_i / L_B \sim \rho_i / R$.
- Perturbative analysis consists of near-identity transformations to new variables which remove the **gyro-phase** dependence in perturbed fields $\delta\mathbf{A}(\mathbf{x})$, $\delta\phi(\mathbf{x})$ where $\mathbf{x} = \mathbf{R} + \boldsymbol{\rho}$:
Expansion in $\varepsilon_\phi = e[\delta\phi - (v_{||}/c)\delta A_{||}]/T_e \sim \delta B_{||}/B_0$.
- Derivation more transparent, less amount of algebra

Hierarchy of Nonlinear Governing Equations

Nonlinear equations: From fundamental, primitive to reduced, simplified	Steps for reduction	Physics lost due to reduction
Vlasov-Klimontovich equation [Klimontovich 67]	Remove high frequency terms ($\geq \omega_{ci}$)	
Gyrokinetic Equation: Conservative [Dubin 83, Hahm 88a,b, Brizard 89, 06]		High frequency phenomena
Gyrokinetic Equation: Conventional [Frieman 82]	Neglect velocity space nonlinearity	Conservation of energy between particles and fields, of phase-space volume, nonlinear trapping of particles along B .
Gyrofluid Equation [Beer 96, Dorland 93]	Take moments in velocity space	Some nonlinear kinetic effects including inelastic Compton scattering [Mattor 92], accuracy in damping rates of zonal flow [Rosenbluth 96] and damped mode [Sugama97]
Fluid Equations [Yagi 94, Scott 97, Zeiler 97, Xu 02, Simakov 05]	Expansion in finite Larmor radius terms; Ordering for collisional plasmas	Most kinetic effects associated with long mean free paths and finite size orbits.

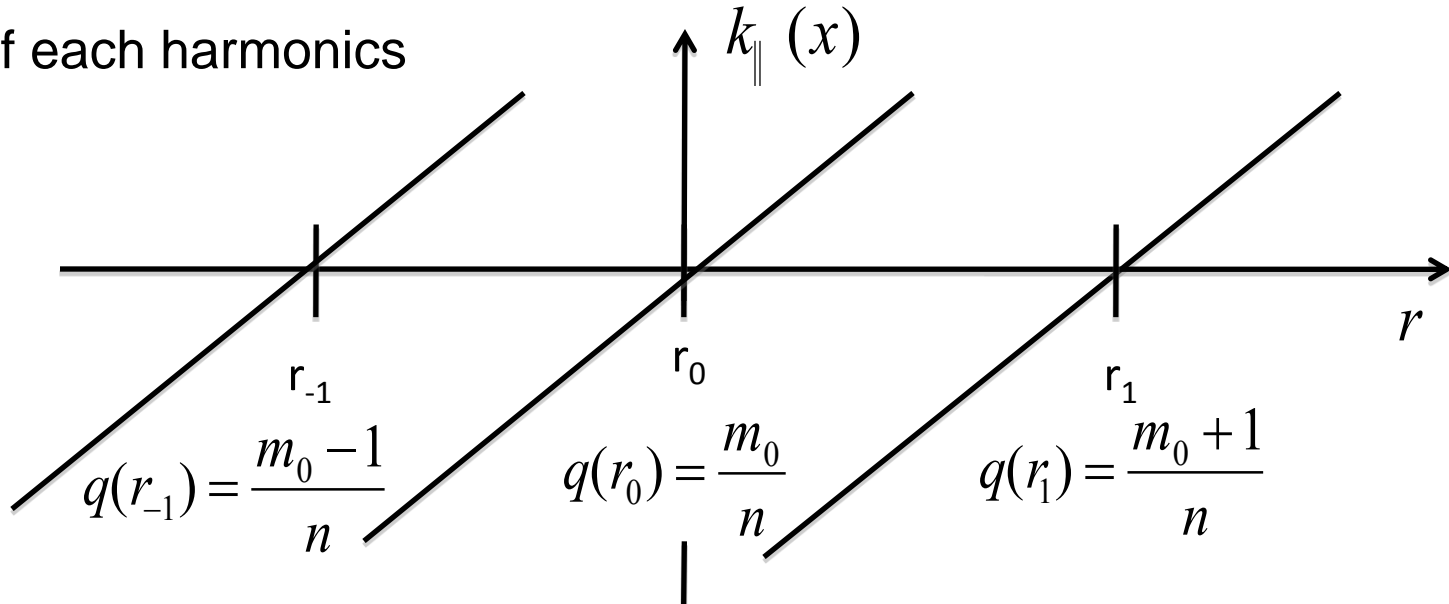
from Diamond-Itoh-Itoh-Hahm : PPCF 47, R35, (2005)

Shear-Alfvén Continuum in Sheared Magnetic Field

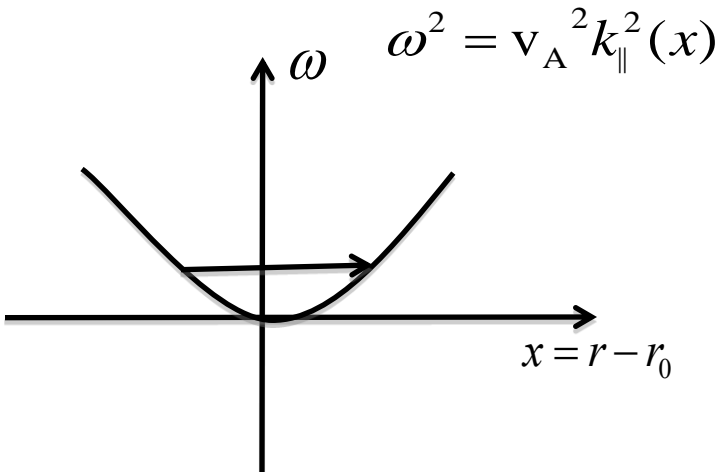
- When driving is weak, shear-Alfvén wave DR. $\omega^2 = k_{\parallel}^2 v_A^2$
 - In sheared magnetic field, with $k_{\parallel} = \frac{nq(r) - m}{qR} \cong \frac{k_{\theta}}{L_s} (x - x_{m,n})$
- For given n, m , linear D.R. is satisfied at least one radial position for any reasonably small values in k_{\parallel}
 - as $k_{\parallel}(x)$ is varied as a function of x , ω assumes “continuum” of values, rather than an “eigenvalue” (discretized)
 - Alfvén continuum → initial wave packet will phase-mix and decay algebraically in time.
- Then what’s the consequence of toroidal geometry?
 - i.e., coupling between neighboring poloidal harmonics

Linear Coupling of Poloidal Harmonics

- k_{\parallel} of each harmonics



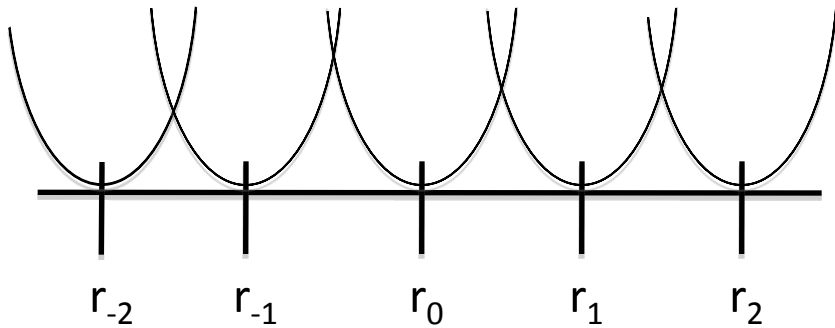
- Shear-Alfvén Continuum of each poloidal harmonics (in Slab)



For $\forall \omega$, dispersion relation satisfied at one x .

Toroidicity-Induced Alfvén GAP

$$\omega^2 = k_{\parallel}^2 v_A^2, \text{ with } k_{\parallel} \propto \frac{k_{\theta}}{L_s} (r - r_m)$$



➔ Each harmonics considered in sheared slab

$$v_A^2 \propto B^2 = \text{const}$$

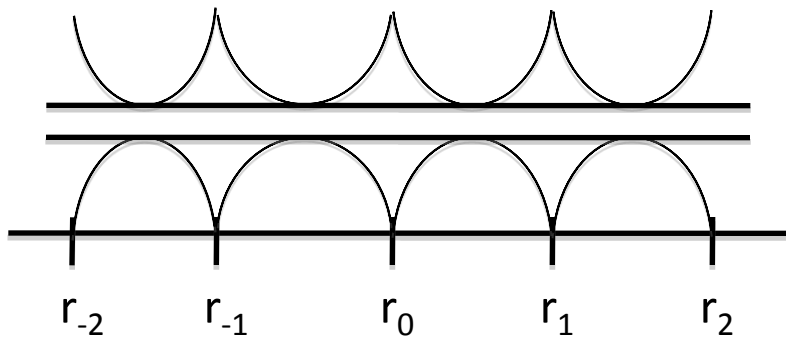
$$B = \frac{B_0}{1 + \frac{r}{R_0} \cos \theta} \text{ induces toroidal GAP}$$

→ (involves solving Mathieu Equation)

Continuum Dispersion Relation Not Satisfied for

$$\frac{v_A}{2qR} (1 - \varepsilon) < \omega < \frac{v_A}{2qR} (1 + \varepsilon)$$

$$\omega^2 = \omega_A^2(r)$$



Toroidal Alfvén Eigen modes

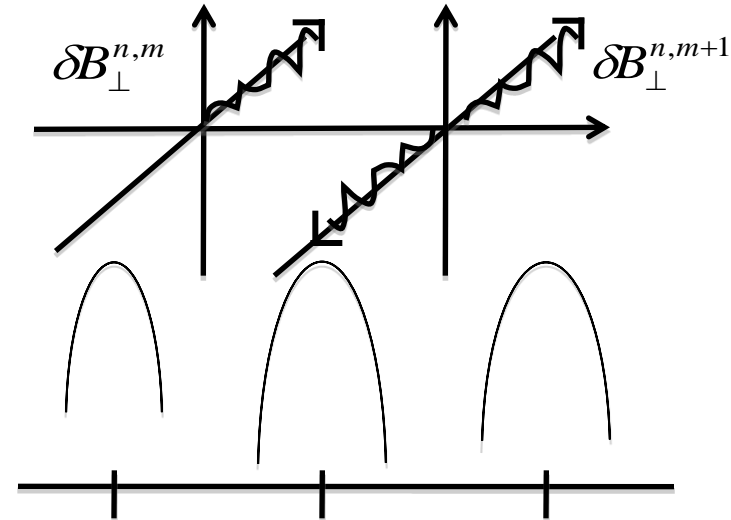
- At the midpoint between two adjacent rational surfaces

$$k_{\parallel} = \frac{1}{2qR} \sim \text{GAP occurs near } \omega \simeq \frac{v_A}{2qR}$$

- “Standing wave formation” from superposition of

$$\delta B_{\perp}^{n,m} e^{i\text{co-prop}} \text{ and } \delta B_{\perp}^{n,m+1} e^{i\text{counter-prop}}$$

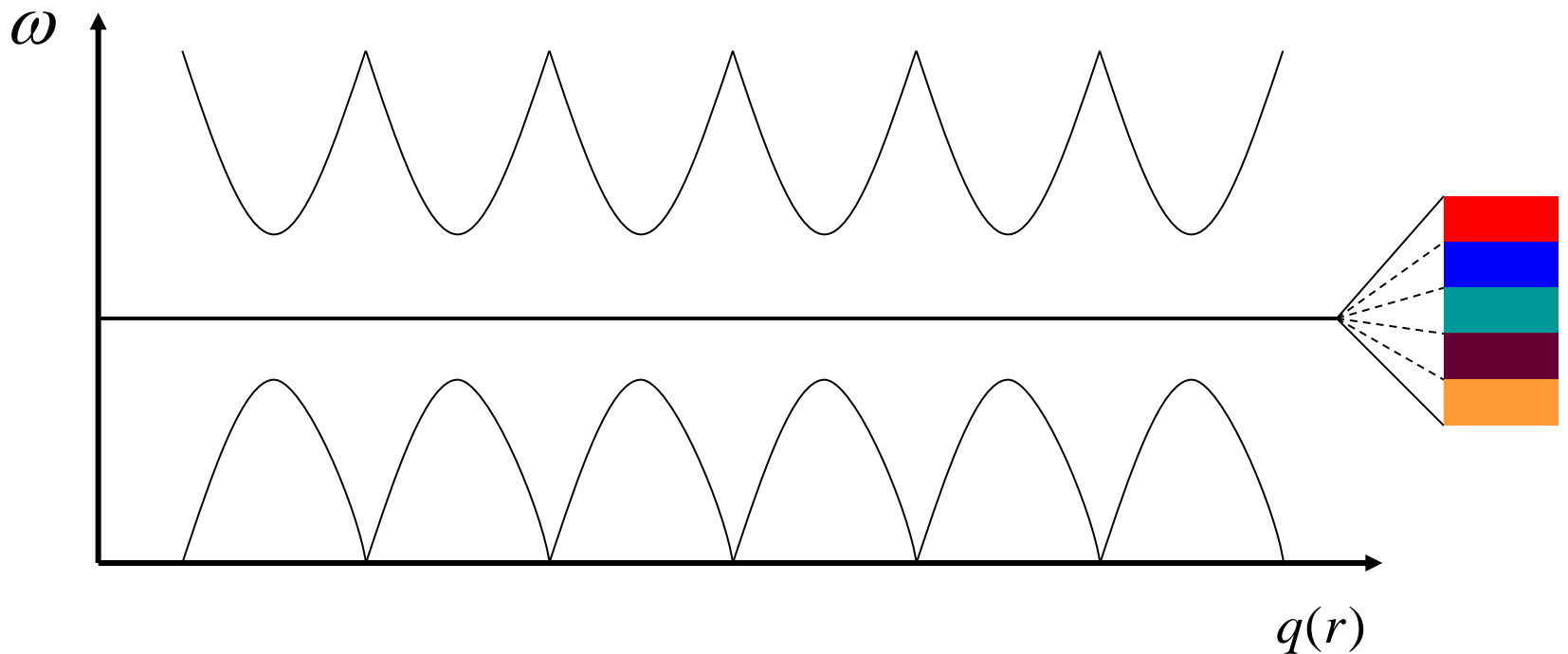
- This “TAE” modes can be excited via resonance with energetic ions.



- ... AE zoo accommodates TAE, BAE, GAE, CAE, HAE, EAE, LSAE, RSAE, ..., and
Nonconventional AE !

Translational Invariance

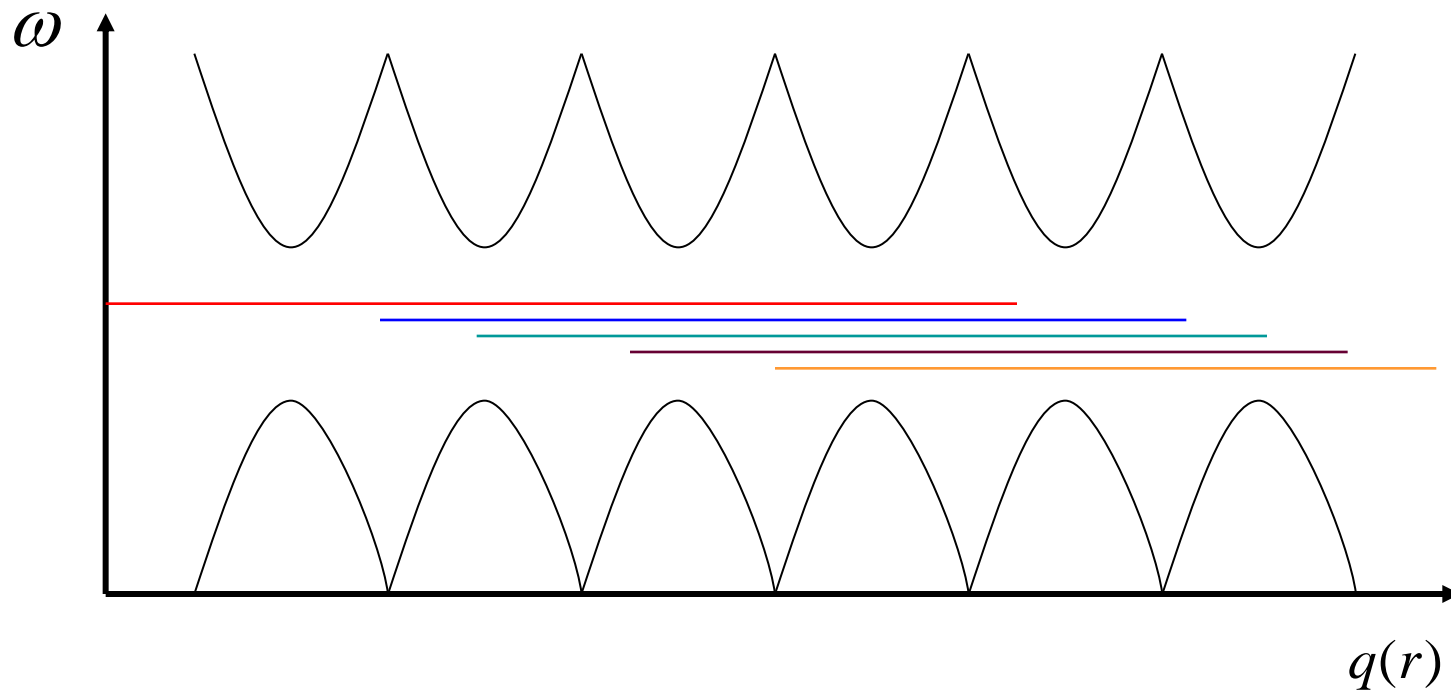
- If $\frac{v_A}{qR_0}$ is uniform in r , $\sim N(q(a) - q(0))$ modes have the same eigenfrequency (degenerate).



For single-N,

Quasi-Translational Invariance

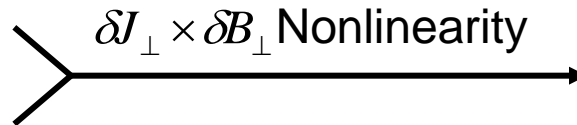
- With equilibrium variation, degeneracy is broken. Each TAE's has slightly different eigenfrequency.
“High-N TAE” still contains many poloidal harmonics.



Nonlinear Saturation Mechanism for High-N TAE

“ Ion Compton Scattering ”

TAE ω



“ BEAT WAVE ”

$$\omega'' = \omega - \omega'$$

$$k''_{\parallel} = k_{\parallel} - k'_{\parallel}$$

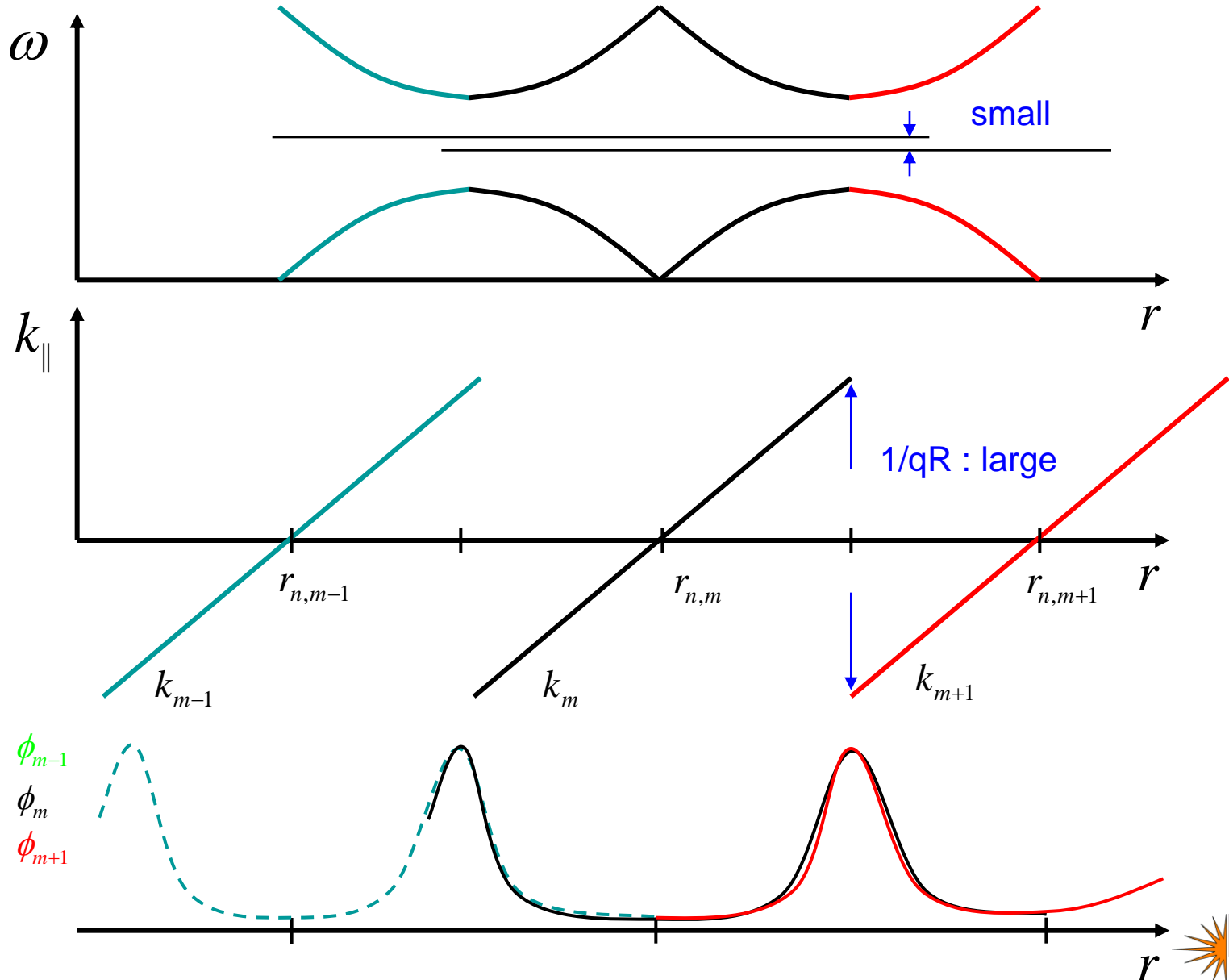
TAE ω'

- If $\frac{\omega''}{k''_{\parallel}} \sim v_{Ti} \ll v_A$, **via Compton Scattering**, fluctuation energy is transferred to lower frequency mode, eventually absorbed by linearly stable mode near lower continuum.



“ Look for Nonlinear Coupling Channel ” via beat wave with low phase velocity which gives $\frac{\omega''}{k''_{\parallel}} \sim v_{Ti}$!

Generation of Low Phase Velocity Beat Wave



Nonlinear Interaction of TAE's

→ Sound Wave-like Density Fluctuation

	Test mode (TAE)	Background mode (TAE)	Beat wave (Sound wave)
Wave Vector	\vec{k}	\vec{k}'	$\vec{k}'' = \vec{k} - \vec{k}'$
Frequency	ω	ω'	$\omega'' = \omega - \omega'$ $\lesssim \varepsilon\omega$ <small>small</small>
k_{\parallel} @ gap	$\frac{1}{2qR}$	$-\frac{1}{2qR}$	$\frac{1}{qR}$ <small>large</small>
• Single – n :			
Toroidal Mode Number	n	n	0
Poloidal Mode Number	m	$m+1$	-1
• For multi – n :			
Toroidal Mode Number		n'	$n - n' \ll n$
Poloidal Mode Number		$m' + 1$	$m - m' - 1 \ll m$

* This talk : $10^1 \leq N \ll 10^2$

Weak Turbulence Nonlinear Analysis [Hahm&Chen, PRL '95]

- 3rd Order Perturbation Theory : $\frac{\gamma_L}{\omega_A} \ll 1$

- 1st order : Test TAE (\vec{k}) ideal MHD

$$\frac{\partial}{\partial t} \psi_{\vec{k}}^{(1)} = -\hat{\mathbf{n}} \cdot \nabla \phi_{\vec{k}}^{(1)}$$

$$\frac{\partial}{\partial t} \nabla_{\perp}^2 \phi_{\vec{k}}^{(1)} = -v_A^2 \hat{\mathbf{n}} \cdot \nabla \nabla_{\perp}^2 \psi_{\vec{k}}^{(1)}$$

- 2nd order : Nonlinear Interaction of two TAE's, $\begin{pmatrix} \vec{k} \\ \vec{k}' \end{pmatrix}$.
 → Sound-wave-like density fluctuation, \vec{k}'' .

$$\left(\frac{\partial}{\partial t} + v_{\parallel} \hat{\mathbf{n}} \cdot \nabla \right) \delta f_{\vec{k}''}^{(2)} = \left\{ \left(\frac{\delta \vec{B}}{B_0} \cdot \frac{\partial}{\partial t} \vec{v}_E \right)_{\vec{k}''} + \frac{|e|}{M_i} \hat{\mathbf{n}} \cdot \nabla \phi_{\vec{k}''} \right\} \frac{\partial f_0}{\partial v_{\parallel}}$$

$\delta J_{\perp} \times \delta B_{\perp}$ NL

“ $\omega'' = k'' v_{\parallel}$ Resonance important.”

- 3rd order : Nonlinear Evolution of Test TAE (\vec{k}),
 in the presence of density fluctuation $\delta n_{\vec{k}''}^{(2)}$
 and other TAE's $\phi_{\vec{k}'}^{(1)}$.

$$v_A^2 \hat{\mathbf{n}} \cdot \nabla \nabla_{\perp}^2 \psi_{\vec{k}} + \frac{\partial}{\partial t} \nabla_{\perp}^2 \phi_{\vec{k}} + \sum_{\vec{k}'} \nabla \cdot \left(\frac{\delta n_{\vec{k}''}^{(2)}}{n_0} \frac{\partial}{\partial t} \nabla_{\perp} \phi_{\vec{k}'}^{(1)} \right) = 0.$$

Nonlinear Evolution of TAE in the presence of density fluctuation

$$v_{A0}^2 \left(\hat{n} \cdot \nabla \nabla_{\perp}^2 \psi_{\vec{k}} + \sum_{\vec{k}'} \nabla \psi_{\vec{k}'} \times \hat{n} \cdot \nabla \nabla_{\perp}^2 \psi_{\vec{k}''} \right) + \frac{\partial}{\partial t} \nabla_{\perp}^2 \phi_{\vec{k}} + \sum_{\vec{k}'} \nabla \phi_{\vec{k}'} \times \hat{n} \cdot \nabla \nabla_{\perp}^2 \phi_{\vec{k}''} + \sum_{\vec{k}'} \nabla \cdot \left(\frac{\delta n_{\vec{k}''}^{(2)}}{n_0} \right) \frac{\partial}{\partial t} \nabla_{\perp} \phi_{\vec{k}'} = 0$$

- Recall; Vorticity Equation is $\vec{\nabla}_{\perp} \cdot \delta \vec{J}_{\perp}^{\text{pol}} + \nabla_{\parallel} \cdot \delta J_{\parallel} = 0$

where “ $\delta \vec{J}_{\perp}^{\text{pol}} = \frac{1}{B_0} \hat{n} \times \rho \frac{d\vec{v}_E}{dt}$ ” depends on “number density.”

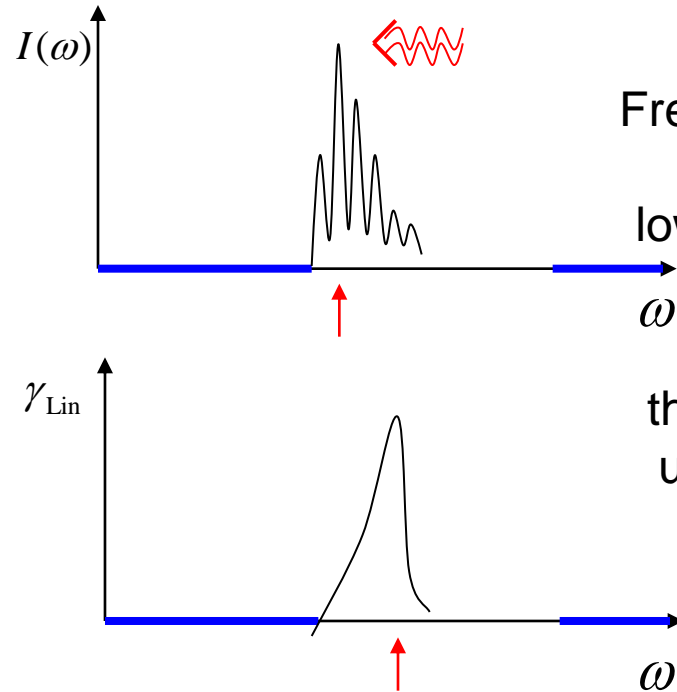
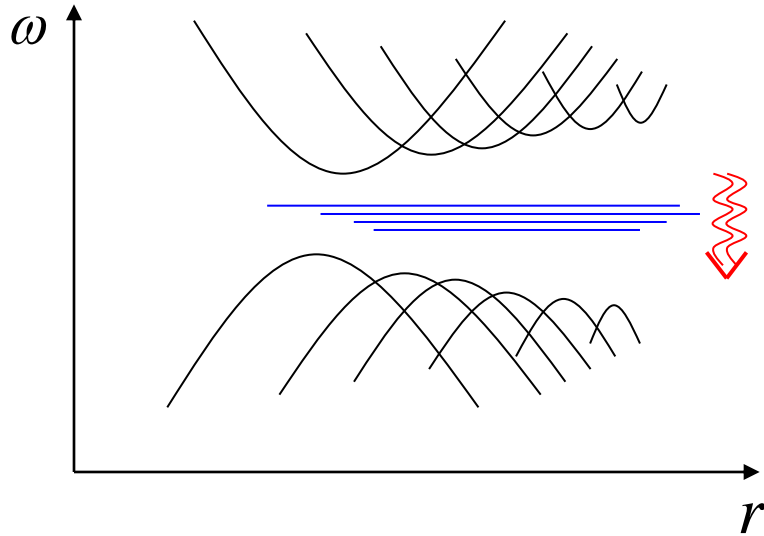
- Other nonlinearities are subdominant for $k'_{\perp} k''_{\perp} \rho_s^2 \ll \frac{\omega'}{\Omega_{ci}}$

* Multiplying $\phi_{\vec{k}}^*$, take imaginary part of the radial average, we get

$$\frac{\partial}{\partial t} I_{\vec{k}} = \gamma_L(\vec{k}) I_{\vec{k}} - \sum_{\vec{k}'} M_{\vec{k}, \vec{k}'} I_{\vec{k}'} I_{\vec{k}}$$

$$I_{\vec{k}} \equiv \left\langle |\nabla_{\perp} \phi_{\vec{k}}|^2 \right\rangle, \quad M_{\vec{k}, \vec{k}'} \equiv \frac{\omega'}{2} \frac{\chi_e^2 \text{Im} \chi_i}{|\chi_e + \chi_i|^2} \frac{M_i}{B_0}$$

“ Frequency Chirping from Ion Compton Scattering ”



Frequency chirps
towards
lower continuum
from
the linearly most
unstable mode.

Fluctuation Energy is transferred to
Lower Frequency due to Ion Compton Scattering.

At Nonlinear Saturation

$$\left(\frac{\delta B_r}{B_\phi} \right)^2 \simeq \frac{1}{4\pi} \left(1 + \frac{T_e}{T_i} \right)^2 \left(\frac{\bar{\gamma}_L}{\omega_A} \right) \left(\frac{r}{R_0} \right)^4$$

• Magnitude : $\frac{\bar{\gamma}_L}{\omega_A} < \frac{\bar{\gamma}_A^{\text{Max}}}{\omega_A} \lesssim 10^{-2}$

$$\frac{r}{R_0} \sim 10^{-1} \quad \Rightarrow \quad \frac{\delta B_r}{B_\phi} \lesssim 10^{-3}$$

• Scaling : $\frac{\delta B_r}{B_\phi} \propto \left(\frac{\bar{\gamma}_L}{\omega_A} \right)^{\frac{1}{2}}$: weak turbulence theory

($\Leftrightarrow \delta B_r \propto \gamma_L^2$; à la single wave trapping) Berk, Breizman, Fu, W. Park(not H. Park), Wu, White, Rosenbluth ...

Our Mechanism :

- More relevant for High-N Multi-mode Overlapping Case. Also for stronger drive.

Single Wave Trapping (Berk-Breizmann et al.)

- Resonant particles are trapped in the potential well produced by TAE.
- The potential well will last an auto-correlation time, ω_{ac}^{-1} .
- A trapped particle will transverse a closed orbit in, ω_b^{-1} , mixing hot-particle distribution function.

→ Require : $\underline{\omega_b > \omega_{ac}}$

Nonlinear Saturation : $\omega_b \sim \gamma_{Lin}$

- It also requires a well-defined potential well in space,

$$\underline{\Delta X_{\text{particle, excursion}} < \Delta X_{N,M}}$$

,where $\Delta X_{N,M}$: distance between neighboring rational surfaces.

Weak Turbulence Theory

Requires

$$\omega_b < \omega_{ac}$$

and

$$\Delta X_{\text{ptl. exc.}} > \Delta X_{N,M}$$

Chirikov Criterion

Validity Regimes

* For Quasi-Linear Theory (Weak Turbulence Expansion)

$$\bullet \Delta X_{\text{ptl. exc.}} > \Delta X_{N,M}$$

$$\Rightarrow \frac{\delta B_r}{B_\phi} > \frac{1}{8q^4 \hat{s}} \left(\frac{r_0^2}{R_0 \rho_\alpha} \right) \cdot \frac{1}{N^4}$$

$$\bullet \omega_b < \omega_{ac}$$

$$\Rightarrow \frac{\delta B_r}{B_\phi} > \frac{1}{8q^4 \hat{s}} \left(\frac{r_0^2}{R_0 \rho_\alpha} \right) \cdot \frac{4\varepsilon^2 \hat{s}^2}{N^2}$$

* For Single Wave Trapping

$$\Delta X_{\text{ptl. exc.}} < \Delta X_{N,M}, \quad (\text{Then, } \omega_b > \omega_{ac} \rightarrow 0)$$

PLOT : in $\left(\frac{\delta B_r}{B_\phi}, N \right)$ space

Approximate equilibrium parameters for ITER PRETOR.

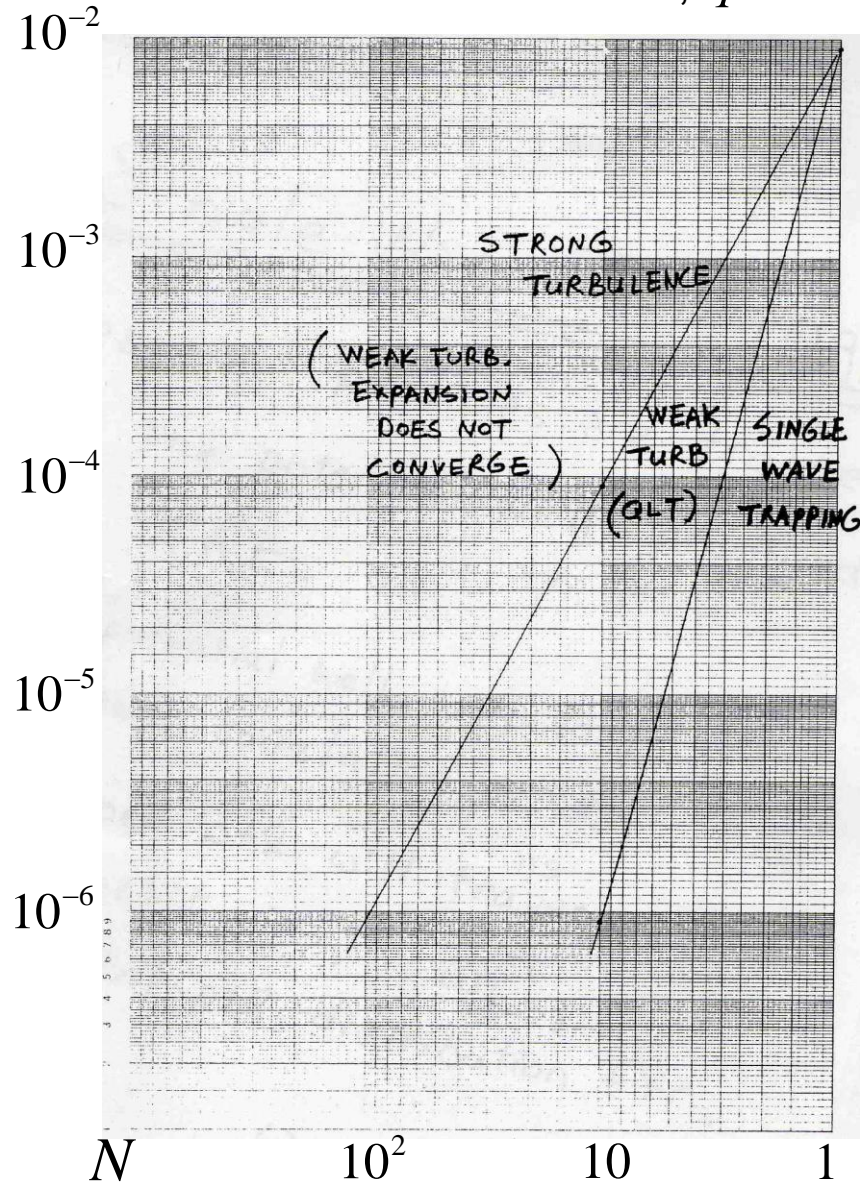
[Candy, Rosenbluth, NF '95]

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Validity Regimes For ITER PRETOR Parameters

$$(\delta B_r / B_0)$$

$r/a \simeq 0.8, q \simeq 2$ dense mode packing



Incompressible MHD Turbulence

Tutorial at Theorié-Fest, Aix-en-Provence, 200? by P.H. Diamond
connects :

- **Weak Turbulence Kinetic Theory** : Sagdeev & Galeev '67
- **Scaling Derivation** : PHD & Craddock: Comments in Plasma Phys '90
Lazarian & Vishniac: Ap. J. '99

and

- **Lengthy Crank** : Goldreich & Sridhar: Ap. J. '95

Common Theme :

Nonlinear Interaction of Low Frequency Beat Wave
with
Particles or **Eddys**

	Weak Turbulence Theory of Incompressible MHD Turb.	WTT of Toroidal Alfvén Eigenmodes in Tokamaks
High Freq. Shear-Alfvén Waves interact nonlinearly with	Eddys	Thermal Ions
allowed by	Low Freq. Beat Wave produced by Counter-Prop ⁿ	Low Frequency Beat Wave produced by “standing” TAEs (formed linearly by Counter-Prop ⁿ of neighboring harmonics)
resulting in	Alfvén Effect (down by $\frac{\Delta\omega_k}{k_{\parallel}v_A}$) and Scale-dependent Anisotropy (Intermediate Turbulence : G&S Ap. J. '97 \simeq Anisotropic I-K '65)	Down-shift of Frequency k_{\parallel} determined by equilibrium \vec{B} geometry k_{\perp} by linear drive (energetic particles)
Theory breaks down	when $k_{\parallel}v_A = v_{\perp} / \ell_{\perp}$ for small scales first	with non-overlapping island (violation of Chirikov Criterion) in phase-space, for large scales first
turning into	Critically Balanced Cascade (G&S '95 '97)	Single-Wave Trapping → Hole-Clump Pair in phase-space. Berk-Breizmann Paradigm

Conclusions

Alfvén waves have been studied from different angles.

- MHD Turbulence Community : Fixation with k-spectra
...
- MFE Theory Community : Fixation with linear instability
zoology
TAE, GAE, RSAE, LSAE, ...
- Nonlinear Gyrokinetic Theory-based
Extensions may bridge some gaps.