

Multi-Scale Interaction in Diffusive Shock Acceleration

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T also KAIIST

Outline

i.) Short Commentary on DSA Theory

- Very Brief Overview
- Two Critical Issues in DSA → Multi-Scale Interaction

ii.) #1: Confinement on Mesoscales ($kr_g < 1$)

primary
focus

- Inverse “Cascade” of $\langle \tilde{B}^2 \rangle_k$ via Modulational Instability
 - Alfvén Wave Packet Dynamics in Shocklet Scatterer Field
 - Stochastic Refraction / “Inverse Cascade” Induced Amplification of Precursor Density Fluctuations
- Quasi-steady and Unsteady States

P. D. and Malkov
Ap. J. '07

iii.)
relation
to (ii.)

#2: Enhanced Acceleration Efficiency (the knee and beyond...)

- Classical Fermi → "Pinball Process"
i.e. scattering by shocklets in convergent precursor flow
- Confinement → Trapping and Loss Islands in Phase Space

fractional
kinetics

→ DSA not entirely "D"
↑ ↑

Malkov and P.D.
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iv.) Where Next?

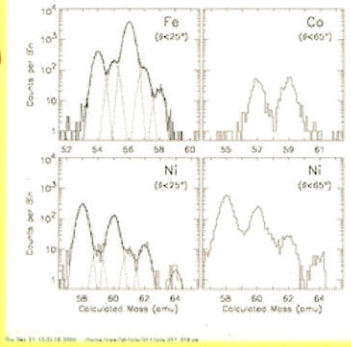
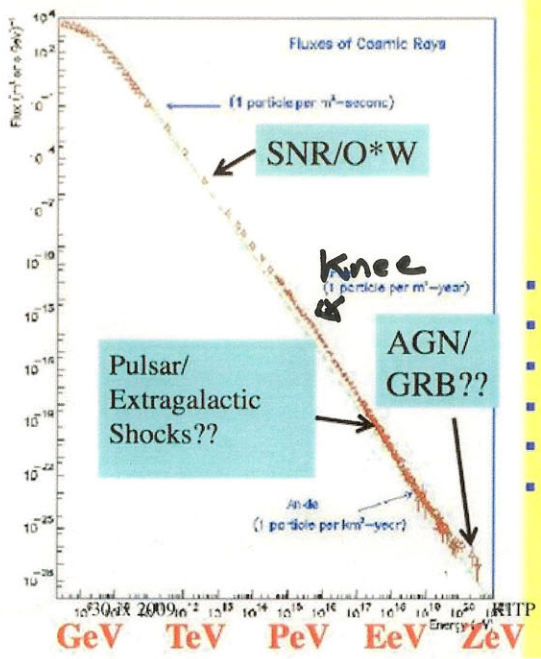
- "Radiation Hydro" Theory of Drury Instability
- Analysis of System States

and ?

N.B.: Observations at the Outset

- 2 fluctuation populations $\left\{ \begin{array}{l} kr_g \sim 1 \rightarrow \text{resonant Alfvén waves} \\ \text{(small scale)} \\ kr_g \ll 1 \rightarrow \text{shocklets} \\ \text{(mesoscales)} \end{array} \right.$
- ∇P_{CR} is crucial drive
- Shocklets directly link $\left\{ \begin{array}{l} \text{Enhanced acceleration frequency} \\ \text{"Inverse Cascade"} \end{array} \right.$

Cosmic Ray Spectrum



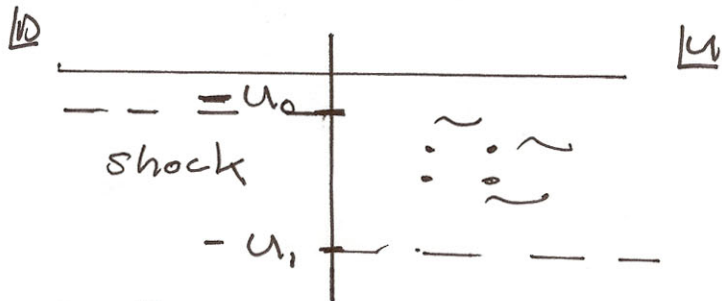
- $N \sim E^{-2.6}$ below "knee"
- $U_{CR} \sim U_{GeV} \sim U_{ISM} \sim U_{CMB} \sim U_{star}$
- $0.1 \text{ Myr} < t < 15 \text{ Myr}$
- $L/M \sim 5 E_9^{-.3} \text{ g cm}^{-2}$
- $\Rightarrow S \sim E^{-2.3}$
- $L_{CR} \sim U_{CR} M_{gas} c \lambda^{-1}$
 $\sim 3 \times 10^{33} W \sim 0.03 L_{SNR} \sim 10^{-3} L_{gal}$

Can SNR accelerate up to PeV "knee"?

→ "knee"
 → spectral break $\sim 10^{15}$
 → can DSA - accelerate particles to knee?
 - explain/encompass knee?

vs. by R. D. Blandford

Usual story of DSA - I)



Shock + Confinement
(circa late 70's)

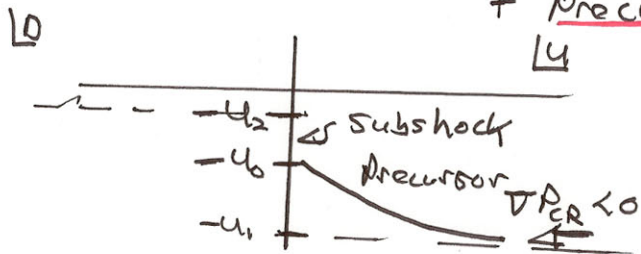
Confinement:

- resonant pitch \neq scattering by CR-emitted AW's
- resonance: $k\rho/m = \Omega_g$?
 high energy \leftrightarrow large scale \leftrightarrow $\left\{ \begin{array}{l} k\rho_g \sim 1 \\ \text{(max growth)} \end{array} \right.$
- scattering: $D \sim c^2 / v_{\text{eff}}$
 $v_{\text{eff}} \sim \Omega (\delta B / B_0)^2$
 typically $\delta B / B_0 \leq 1$ $\left\{ \begin{array}{l} \text{efficiency?} \\ v_{\text{acc}} \sim u_{\text{sh}}^2 / D \end{array} \right.$
- convection-diffn. model

$$\frac{\partial f}{\partial t} + u \frac{\partial f}{\partial x} - \frac{\partial}{\partial x} D \frac{\partial f}{\partial x} = + \frac{1}{\Omega} \frac{\partial u}{\partial z} \rho \frac{\partial f}{\partial p}$$

Usual story II - Self-Consistent Shock

+ Precursor ('80's \Rightarrow)



\rightarrow discontinuity
+ CR Cloud (P_{CR})
(\Rightarrow Flow)

Fluctuations:

- AW's



high frequency, (micro)
($kV_g \sim 1$)

-(magneto) acoustic modes \Rightarrow



"shocklet
trains/turbulence"

\rightarrow low frequency (meso)

How?

- linearly, Drury Instability

$$-\gamma_0^{\pm} = -\frac{\gamma_{CR}}{\omega} + \frac{P_{CR}}{\omega} \quad (1 + \partial \ln \rho / \partial \ln \rho)$$

(diffn) ω

$$\frac{P_0}{L_{pre}} < k V_g < 1$$

- or ---- ?

role in confinement ?

$$l_g < \lambda < L_{pre}$$

mesoscale range

→ Issues in DSA ?

6.

① - how confine high energy particles?
⇒ require populating larger scales AW range
 $k\rho/m = \Omega$

② - how improve the efficiency of acceleration ?

③ - is DSA "D" ? - if not, how represent?
physics?

① how generate AW's with $k\rho_g < 1$?
⇒ requires spectral energy transfer to long wavelength
- counter to lone of MHD Alfvén cascade

② { increased P_{CR} "inflates" shock structure
time to cross ↑ with energy
⇒ exploit precursor ?

③ speculations...

A Path Forward → Multi-Scale Interaction 7

- Two Component Turbulence (à la Langmuir) Turbulence

Acoustic modes

Alfven waves



$k > 1/L_{pre}$ {stochastic refraction}

$1/L_{pre} \ll k \ll 1/r_g$

{ linear - Drury Mech.
nonlin. - pumped by AW

{ resonant growth
scattered/modulated
by acoustics ⇒

stochastically refract/modulate
Alfven waves

decay instability



- upshot:

→ pumping of wave population
on mesoscales

↓ $k > 1/r_g$
population conversion

→ excitation of large scale "Drury" Modes by decay
(initiates need for robust Drury inst.).

→ Key Physics: Stochastic Refraction

P.

- Consider A/Fuen wave packet in field of acoustic scatterers:

$$\omega = k v_A$$



$$\frac{dk}{dt} = -\frac{\partial}{\partial x} k v_A \approx \frac{k v_A}{2} \frac{\partial (\tilde{\rho}/\rho_0)}{\partial x}$$

scatterers:
density bumps
("shocklets")

- ∴ For random ensemble shocklets, diffusion in $k \Rightarrow$ spectral spreading

i.e.

$$D_k \approx \frac{k^2 v_A^2}{4} \sum_z |\tilde{\rho}_z/\rho_0|^2 z^2 \gamma_{0z}$$

{ Diffusivity for stochastic Refraction

- ∴ - population density flux:

$$\Gamma_n = -D_n \partial \langle N \rangle / \partial k$$

\Rightarrow { populates $k \gamma_0 < 1$ from inversion.

$$- \frac{dE}{dt} < 0 \quad (\text{pop/n. imbalance})$$

Key Physics II : Ponderomotive Feedback (Stochastic Analogue NLS) 9.

- How does stochastic refraction pump large scales?

$$\frac{\partial \tilde{v}}{\partial t} = -\frac{1}{\rho_0} \frac{\partial}{\partial x} (\rho_0^2 \tilde{v} + P_{\text{rad}})$$

$$\frac{\partial \tilde{\rho}}{\partial t} = -\rho_0 \frac{\partial \tilde{v}}{\partial x}$$

$$P_{\text{rad}} = \sum_{\underline{n}} \omega_{\underline{n}} (N^+ + N^-)$$

$$\Rightarrow \omega^2 - z^2 c_s^2 = z^2 \sum_{\underline{n}} \frac{\omega_{\underline{n}}}{2\rho_0} \frac{\partial z k v_A}{\partial \omega_{\underline{n}}} \frac{\partial \langle N^{\pm} \rangle}{\partial t}$$

$$\gamma^{\pm} = \frac{z^2}{4\rho_0} \frac{v_A}{c_s} \sum_{\underline{n}} k \omega_{\underline{n}} \frac{\partial \langle N \rangle}{\partial t}$$

- $\tau_{c, \text{NL}} \sim \Delta \omega_{\underline{n}} / z^2 v_{gr}^2 + \Delta \omega_{\underline{n}}^2$
- $\partial \langle N \rangle / \partial t > 0$ required for maximal $\tau_{c, \text{NL}}$
- NL pumping 'cascades' downward (i.e. $\gamma^{\pm} \sim z^2$)

Energies:

$$\rightarrow \frac{\partial \Sigma_{AW}}{\partial t} = \int dk \omega_k \frac{\partial}{\partial k} D_k \frac{\partial \langle N \rangle_{AW}}{\partial k}$$

→ stoch. scattering by acoustics

$$\approx - \int v_{gr} D_k \left[\left| \frac{\partial \tilde{\phi}}{\partial k} \right|^2 \right] \frac{\partial \langle N \rangle_{AW}}{\partial k}$$

↳ acoustic field

$$\frac{\partial \Sigma_{Ac}}{\partial t} = \int dZ \delta_Z \left[\frac{\partial \langle N \rangle_{AW}}{\partial k} \right] \left| \frac{\partial \tilde{\phi}}{\partial k} \right|^2$$

↳ AW wave field.

Books balance → manifestly ...

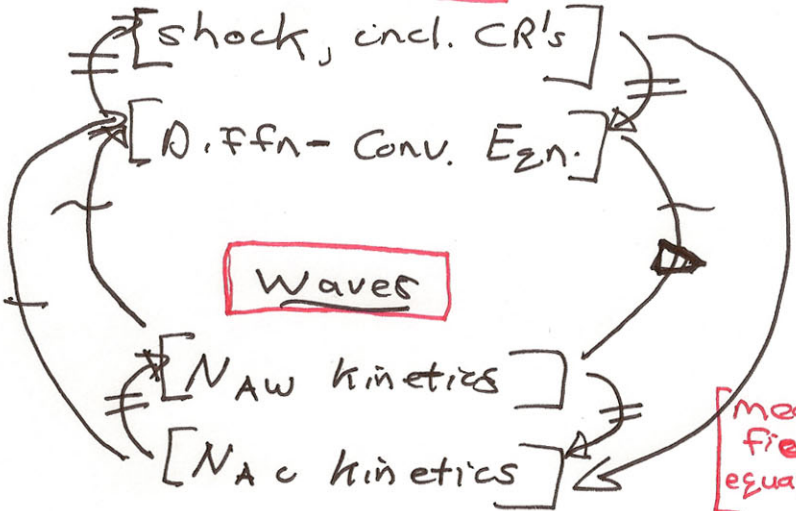
$$\partial_t (\Sigma_{AW} + \Sigma_{Ac}) = \mathcal{S}_o + \mathcal{S}_c.$$

→ akin to $\partial_t (RPHED + WED) = 0$
in quasi-linear theory ...

→ A Bit More Depth

- where does this lead us? → { multi-population system with feedback loops

Gas + C.R.'s



↗ - modulational instability feedback

↻ - CR - AW feedback for confinement

↘ - CR - shock feedback

Bottom line: evolve AW, AC populations on ~~equal~~ footing with F_{CR} , gas dynamics
→ How?

↗ - Drury process

↖ - Acoustic effect on confinement.

→ Deriving Wave Population Evolution

E_{zns} - { Mean Field Theory + Wave Kinetics }

i.e. { Consider AW wave kinetics on AC scatterer field }

$\partial_t \langle N^\pm \rangle + (u \pm v_A) \frac{\partial}{\partial x} \langle N^\pm \rangle - kv_x \frac{\partial}{\partial k} \langle N^\pm \rangle + \frac{\partial}{\partial k} \left\langle kv_A \frac{\partial}{\partial x} \tilde{N}^\pm \right\rangle = \gamma_k^\pm \langle N^\pm \rangle + \langle C(N^\pm) \rangle$

→ closure needed

- proceed a la Chapman-Enskog :

$\gamma_k^\pm N^\pm + C^\pm \{ N^\pm \} = 0$ e.o.

$L^\pm \tilde{N}^\pm = -kv_A \frac{\partial}{\partial x} \frac{\partial}{\partial k} \langle N^\pm \rangle$ 1st O.
propagator $[L^\pm = L_{linear} + \Delta \omega_H]$ $\gamma^\pm \tilde{N}^\pm + \frac{\partial C}{\partial N} N^\pm \approx -\Delta \omega_H \tilde{N}^\pm$

$$\frac{\partial}{\partial t} \langle N^\pm \rangle + u \frac{\partial}{\partial x} \langle N^\pm \rangle - \kappa u_x \frac{\partial}{\partial x} \langle N^\pm \rangle$$

$$- \frac{\partial}{\partial k} D_k \frac{\partial \langle N^\pm \rangle}{\partial k} = \gamma_k^\pm \langle N^\pm \rangle + \langle C(N^\pm) \rangle$$

$$D_k = \frac{1}{4} \kappa^2 N_A^2 \left\langle \frac{\partial_x}{\rho_0} L^{-1} \frac{\partial_x}{\rho_0} \right\rangle$$

→ $\langle N \rangle$ evolves via anisotropic mode couplings
non-isoceler

c.e.



vs



→ akin w. T.T. decay, with resonance broadening

→ recovers wave packet diffusion (no surprise)

→ finite amplitude can overcome mis-match

$$\Delta \omega \approx \left[\frac{\kappa^2 u^2}{4} \frac{\langle (\partial_x)^2 \rangle}{\rho_0^2 (\Delta k)^2} - \tilde{\Sigma}_{MM}^2 \right]^{1/2}$$

→ Observation

- like many related systems, N_{AW} map to N_{AC} ecology fed by P_{CR} c.i.e. "Predator + Prey"

n.b.: { Contrast Mean Field Electrodynamics

→ prey fed by linear resonant instability

- Prey : $N_{AW} \rightarrow \frac{dN_{AW}}{dt} = \gamma_{AW}[P_{CR}] - \gamma_{decay}[N_{AC}]N_{AW}$

↙ inverse cascade !!
process

$N_{AC} \rightarrow \frac{dN_{AC}}{dt} = \gamma_{decay}[N_{AW}]N_{AC} - \alpha N_{AC}^2 + \gamma_{Drury}N_{AC}$

↘ steepening

→ $\gamma_D \leq 0 \Rightarrow$ usual

2 species fixed point $\Rightarrow N_{AW} \leftrightarrow \gamma_D \rightarrow$ diffusive damping of Drury modes is regulator

but:

- $\gamma_D > 0 \Rightarrow$ simple story lost ---- }
- enhanced breaking
- non-stationary state }

↑ Drury drive

→ Something Specific: Effectiveness of Diffusive

Spectral Spreading in

Shock driven systems ?

- concern: effectiveness on $\tau \sim L/u$? → precursor, crossing time

- $I_k \equiv |\tilde{B}_k / B_0|^2$

where: $\partial I_k + u \partial I_k / \partial x - \frac{\partial D_k}{\partial k} \frac{\partial I_k}{\partial k} = \frac{2u}{kVA} \frac{\partial P_{CS}}{\partial x}$

spectral spreading

AW growth/drive

$D_k = \frac{k^2 VA^2}{4} \sum \frac{z^2 \Delta \omega}{z^2 v_{gr}^2 + \Delta \omega^2} \left| \frac{\tilde{\rho}_z}{\rho_0} \right|^2$

- take: $\left| \tilde{\rho}_z / \rho_0 \right|^2 \approx \overset{\text{str.}}{\int} A \left(z_0 / z \right)^2$ { shock amplitude \sqrt{A} , separation $1/z_0$ }
 ⇒ (cf. Kang, '92)

$\therefore D_k \approx \frac{3\pi}{4} A z_0 \frac{VA^2}{v_g} k^2 \equiv v_c k^2$

for $\gamma^3 = D_k v_g' / z_0 v_g^3 \gg 1$. ↳ scattering rate

→ for strong scattering limit

- then, intensity:

$$\frac{\partial}{\partial k} k^2 \frac{\partial I}{\partial k} - \frac{U_1}{r_c} \frac{\partial I}{\partial x} = -2 M_A \frac{U_1}{r_c} \frac{\partial \rho}{\partial x} \quad \left| \begin{array}{l} \text{res.} \\ \rho k = m \Omega_c \end{array} \right.$$

if $r \rightarrow 0 \Rightarrow I \sim \rho|_{\text{reson}}$, as usual

- a bit of algebra \Rightarrow

$$Q = 2 M_A \rho / k$$

$$I = \int_{-\infty}^{\infty} dT' \frac{\exp[-(1/4)(T-T')^2]}{(4\pi(T-T'))^{1/2}} \int_{-\infty}^{+\infty} dv' \exp\left[-\frac{(v-v')^2}{4(T-T')} + \frac{1}{2}(v-v')\right] \frac{\partial \rho(v')}{\partial v'}$$

Governing parameter: $\left\{ \begin{array}{l} S = r_c k \rho / U_1 = r_c \tilde{T}_{\text{conv.}} \\ \equiv \tau_{\text{prec. conv.}} / \tau_{\text{scattering}} \end{array} \right.$

$S \rightarrow 0 \Rightarrow I = Q \sim \rho_{CR}$, as usual.

$S \gg 1$?

$$\delta > 1 \Rightarrow$$

$$I(x) = \frac{2Ma}{5} E(x) \begin{cases} 1 & \rho > \rho_{\max} \\ \rho/\rho_{\max} & \rho < \rho_{\max} \end{cases}$$

ρ_{\min} ρ_{\max}
 $\rho > \rho_{\max} \rightarrow$ upper c.o. (critical spectrum)
 $\rho < \rho_{\max} \sim \rho_b \rightarrow$ lower c.o.

recall $k\rho \sim \text{const}$, so

$$\rightarrow \text{high } \rho \Rightarrow \text{low } k \Leftrightarrow I(k) \sim \text{const} \quad (\sim \text{flat spectrum})$$

$$\rightarrow \text{low } \rho \Rightarrow \text{high } k \Leftrightarrow k^2 \frac{dI}{dk} \sim \text{const}$$

(\sim const. flux)

\therefore { diffusive refraction in k
generates low k and confinement of
higher ρ particles, with $\rho > \rho_{\max}$. !

n.b. here: $E(x) = \frac{\rho_c(x)}{\rho_0 u_1^2} = \frac{4\pi}{3} \frac{mc^2}{\rho_0 u_1^2} \int_{\rho_0(x)}^{\rho_{\max}} d\rho \rho^3 f_0(\rho)$

→ what is \mathcal{S} , really?

- \mathcal{S} determined by structure, spacing, etc. of scatterers

c.e. acoustic shocks $\Delta \rho \sim \rho_0 \Rightarrow \mathcal{S} \sim \rho_0 L / M_A$
(i.e. standard scenario)

- for realistic parameters,
 $\mathcal{S} \gg 1$ confirmed.

$\sim N_s / M_A$
shocklets in train
 $\wedge \wedge \wedge$

- Generally, both short AW wave and particle N.L. dynamics depend sensitively on structure of scatterer field in non/linear regime!

n.b. studies of impact of shock coalescence in non/linear regime especially interesting!

Partial Summary

19.

- modulational decay via stochastic refraction identified as important multi-scale mechanism
- modulational decay:
 - generates larger scale AW
 - enhances shocklets (vitiates need for linear Drury mechanism)
- spectrum of density gradients $|\tilde{\rho}_z/\rho_0|^2$ is crucial
- critical parameter $S = v_0 \tau_{\text{conv}}$
 - set by scatterer configuration (i.e. spatial structure)
- demonstrated $S > 1 \Rightarrow$ spectral broadening
enhanced acceleration + conf.

Next steps:

→ self-consistent theory of Drury instability
coupled to AW spectrum

i.e. - modulate R and $\langle \sigma B^2 \rangle$, ...

- retain decay feed, etc.

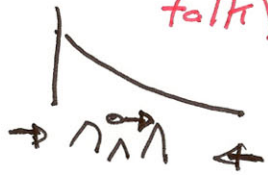
→ Analysis of coupled states / feedback loops
in shock + AW + AC system

II.) Enhanced Efficiency !!

- another weakness of DSA is low efficiency at high energy
i.e. $1/\tau_{acc} \sim U_{sh}^2 / D(p)$
- difficulty arises from long time for multiple shock crossings ?

⇒

- Malkov '06
(and following talk)



→ forego traditional Fermi process

→ explore scattering in convergent precursor flow - "Pinball Process"

→ scatterers: (magneto)acoustic shocklets in precursor flow.

- relation to this → both {enh. conf. / enh. eff} ⇒ effects of DSA driven shocklets.

→ Relation to Multi-Scale Interaction?

- M-S I generates precursor scattering field

i.e. Drury + NL decay + ? ⇒ shocklet configuration
 ⇔ acceleration

- $P_x \leftrightarrow$ knee → location of enhanced loss, in p
 → loss islands (also loss cones)
 in CR phase space, cf.

⇔ scattered configuration !

(i.e. loss cone ⇔ scale, mirror ratio)

- Orbits of CR's no longer simply diffusive..

- trapping osc.

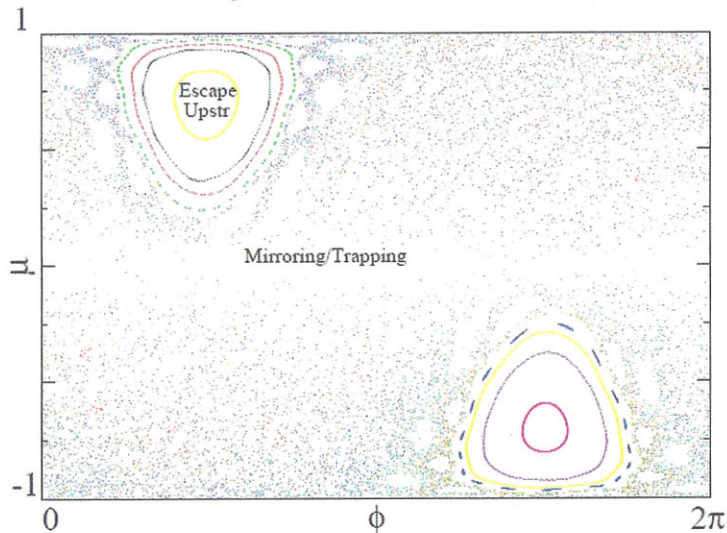
- mirroring; stochastic
 non-stoch

⇒

- Levy flights → super-diffusive

- sticking → sub-diffusive

Particle dynamics and transport in a shock train inside of CR precursor



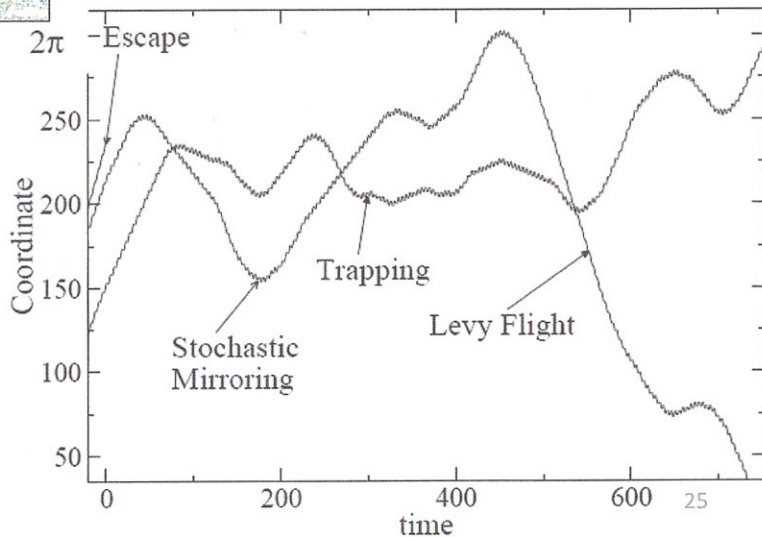
Particle trajectories

Pitch-angle/Gyro-phase

Poincare map

(Pitch-angle wrt shock normal, 45 deg here)

$$r_g(p) / L \simeq 3$$



→ A speculation?! - { convergence of attention on mesoscale scatterer field ⇒

- time to junk time-honored diffusion-convection equation....

- alternative strategy:

→ develop simple characterization of scatterer distribution of scale, separation, etc. for shocklets, 'solenoidlets' (c.f. Bell)

n.b.: linear instability won't suffice

dis ⇒ shock coalescence. ⇒ scatterer field

→ compute particle scattering using scatterer pdf model
c.e. Fractional kinetics (c.f. Zaslavsky)

$$\text{expect: } \frac{\partial}{\partial x} D \frac{\partial}{\partial x} F \rightarrow \frac{\partial}{\partial x} \int dx' K(x, x') \frac{\partial F}{\partial x'}$$

↗ nonlocal interaction kernel

memory!

Conclusions

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- ∇P_{CR} shocklets (Drury or nonlinear)
promising route toward:
 - higher energy confinement \Leftrightarrow via stochastic refraction
 - enhanced efficiency
- linear Drury mechanism not critical \Rightarrow AW pumping...
- Crucial to understand: \Leftrightarrow self-consistency!
 - scale
 - strength
 - spacing \Leftrightarrow packing fraction } of { meso scale scattering field
- need for θ_k , D , K etc. \Rightarrow spectral structure.
- replace/extend DSA to FK SA