

# Data Release of UV to Submillimeter Broadband Fluxes for Simulated Galaxies from the EAGLE Project

Peter Camps<sup>1</sup> , Ana Trčka<sup>1</sup> , James Trayford<sup>2</sup>, Maarten Baes<sup>1</sup> , Tom Theuns<sup>2</sup>, Robert A. Crain<sup>3</sup> , Stuart McAlpine<sup>2</sup>,  
Matthieu Schaller<sup>2</sup> , and Joop Schaye<sup>4</sup> 

## 1. EAGLE: simulations

- Cosmological hydrodynamic simulations using GADGET3 (Tree-PM + SPH)
  - Follow formation & evolution of galaxies
  - Adopt subgrid recipes for radiative cooling, star & BH formation & feedback from stars and AGNs
- (Schaye + 2015)



## 2. EAGLE: Public data release

- halo & galaxy catalogues (McAlpine + 2016)
- particle data (EAGLE Team 2017)



## 3. SKIRT (based on Baes + 2011)

- post-processing of galaxy data from EAGLE
- assign **dust** to star-forming gas particles
- Forward modelling: 3D radiative transfer
- Predict broadband fluxes from UV to sub-mm (Camps + 2018)

# The EAGLE project: simulating the evolution and assembly of galaxies and their environments

Schaye + 2015

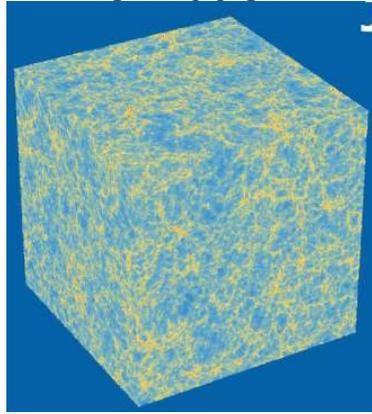
## Evolution and Assembly of GaLaxies and their Environments



- Volumes of 25 - 100 Mpc and zooms
- Particle mass  $10^5 - 10^6 M_{\odot}$  (smaller for zooms), resolves warm ISM
- Modern SPH **GADGET3**
- Includes feedback from stars and AGN (1 type each)
- Subgrid recipes depend only on local gas properties
- Hydro and cooling never turned off
- Winds develop without predetermined mass loading or velocity
- Stellar feedback efficiency calibrated to  $z = 0$  mass function and galaxy sizes
- AGN feedback efficiency calibrated to  $z = 0$  BH mass – stellar mass relation
- Many different models, spin offs

# The formation of galaxies

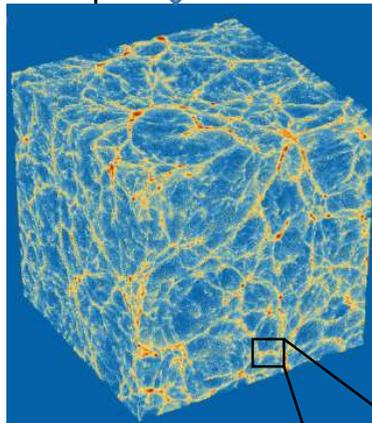
Large-scale structure formation



$100 h^{-1} \text{Mpc}$

$Z \sim 127$

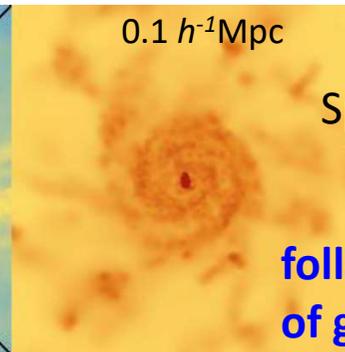
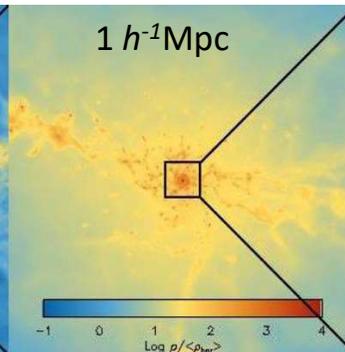
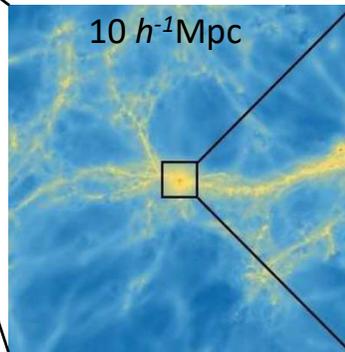
$100 h^{-1} \text{Mpc}$



$Z \sim 0$

follow the emergence of cosmic web

Schaye + 2015



follow the formation of galaxies

Formation of large-scale structure

Dark matter, intergalactic gas  
Primordial fluctuations

Gravity

Universal expansion  
Outflows

Halo of dark matter, gas

Gravity  
Radiative cooling

Gas pressure  
Centrifugal force  
Stripping  
Outflows

Formation of stars & BHs

Feedback from stars & BHs

Interstellar gas, stars, black holes

Subgrid recipes

## 6 set of simulations with GADGET 3 (N-body Tree-PM SPH code)

Identifier	L (cMpc)	N	$m_g (M_\odot)$	$m_{dm} (M_\odot)$	$\epsilon_{com}$ (ckpc)	$\epsilon_{phys}$ (pkpc)	$n_{H,0} (cm^{-3})$
Ref-L0025N0376	25	$2 \times 376^3$	$1.81 \times 10^6$	$9.70 \times 10^6$	2.66	0.70	0.67
Ref-L0025N0752	25	$2 \times 752^3$	$2.26 \times 10^5$	$1.21 \times 10^6$	1.33	0.35	0.67
Recal-L0025N0752	25	$2 \times 752^3$	$2.26 \times 10^5$	$1.21 \times 10^6$	1.33	0.35	0.25
Ref-L0050N0752	50	$2 \times 752^3$	$1.81 \times 10^6$	$9.70 \times 10^6$	2.66	0.70	0.67
AGNdT9-L0050N0752	50	$2 \times 752^3$	$1.81 \times 10^6$	$9.70 \times 10^6$	2.66	0.70	0.67
Ref-L0100N1504	100	$2 \times 1504^3$	$1.81 \times 10^6$	$9.70 \times 10^6$	2.66	0.70	0.67

-DM + Baryonic matter: gas, stars, BH

Plummer potential  
with Softening length

-from  $z=127 \rightarrow$  to  $z=0$

-LCDM cosmology:

$$\Omega_\Lambda=0.693, \Omega_m=0.307, \Omega_b=0.04825, \sigma_8=0.8288, n_s=0.9611, H_0=67.77 \text{ km/s/Mpc}$$

$$-\Delta x_{grid} = 33 - 66 h^{-1} \text{ kpc,}$$

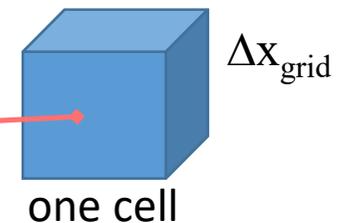
but gravitational softening lengths, smoothing length  $h \ll \Delta x_{grid}$

-subgrid physics model:

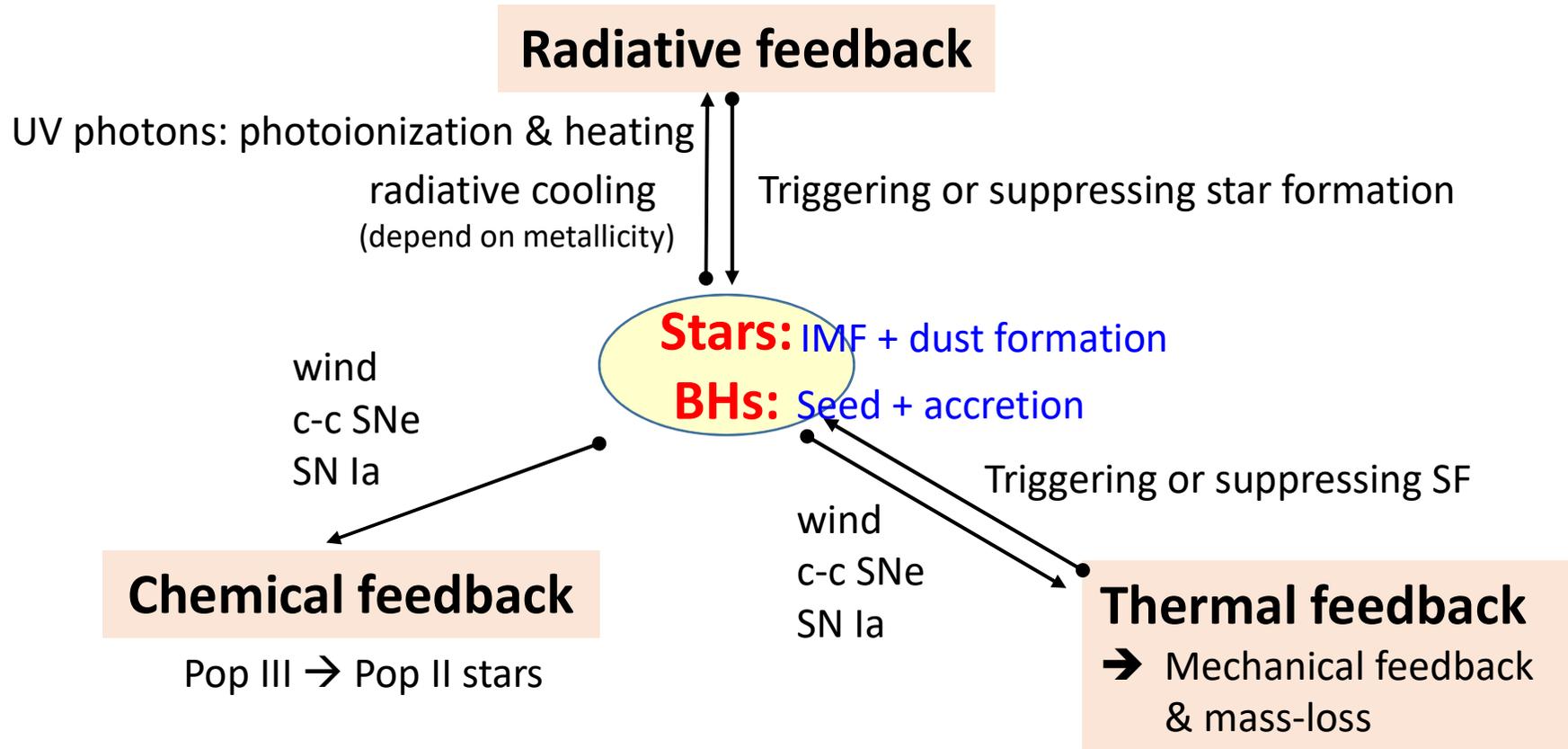
processes not resolved by the numerical scheme are implemented as subgrid source and sink terms.

→ Fine-tuning of a set of free parameters

→ calibrated against observations:



# Subgrid recipes for feedback from stars & BHs



- 1 . stochastic star formation: Kennicutt-Schmidt relations
2. Initial Mass Function by Chabrier (2003) + Dust formation
3. Seed black hole: Bondi-Hoyle accretion + feedback (AGN heating)
4. radiative cooling and photoionization heating with temperature floor
5. stellar mass loss due to winds (massive stars, SN Ia, core collapse SN)
6. thermal energy feedback + chemical enrichment from star formation

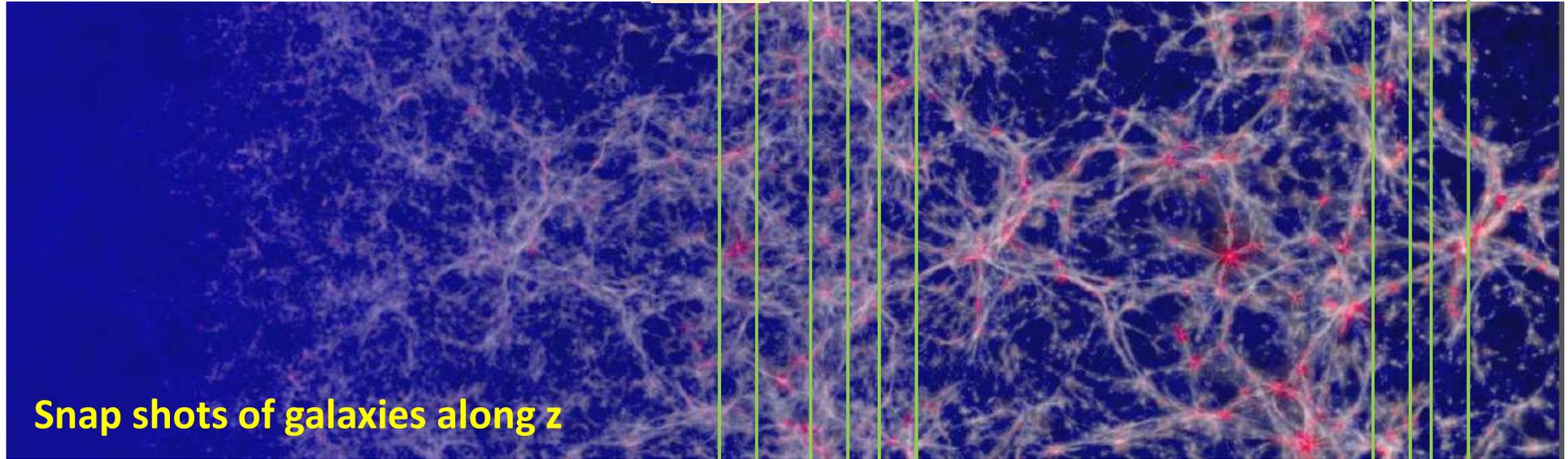
Z=127

time →

Z=20

29 snap shots

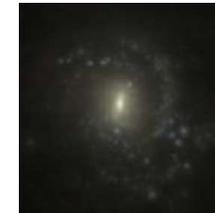
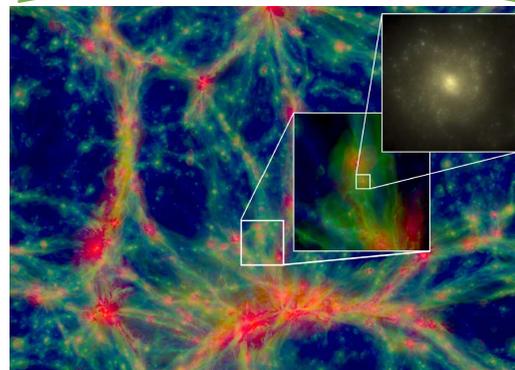
Z=0



**Structured Query Language (SQL):**  
data retrieval tool

Properties of haloes and galaxies

- stellar mass, SFR, metallicity, merger trees (evolution)
- photometric magnitudes & colors: SDSS *u*, *g*, *r* filters
- mock three-color *gri* images



# The EAGLE simulations of galaxy formation: Public release of halo and galaxy catalogues<sup>☆</sup>

S. McAlpine<sup>a,\*</sup>, J.C. Helly<sup>a</sup>, M. Schaller<sup>a</sup>, J.W. Trayford<sup>a</sup>, Y. Qu<sup>a</sup>, M. Furlong<sup>a</sup>, R.G. Bower<sup>a</sup>, R.A. Crain<sup>b</sup>, J. Schaye<sup>c</sup>, T. Theuns<sup>a</sup>, C. Dalla Vecchia<sup>d,e</sup>, C.S. Frenk<sup>a</sup>, I.G. McCarthy<sup>b</sup>, A. Jenkins<sup>a</sup>, Y. Rosas-Guevara<sup>f</sup>, S.D.M. White<sup>g</sup>, M. Baes<sup>h</sup>, P. Camps<sup>h</sup>, G. Lemson<sup>i</sup>

## EAGLE galaxy database

If you already have an account, follow this [link](#) to access the data or register via the form below.

Documentation can be found on the database itself or in the galaxy catalogue [release paper](#).

The python examples from the release paper are available here:

- To connect to the database directly via Python [here](#)
- Galaxy stellar mass function example [here](#).

**Structured Query Language (SQL)**  
+ Python

## EAGLE particle data

If you already have an account, the particle data can be accessed from this [link](#). Registration is done via the form below.

The documentation accompanying the particle data can be downloaded in the data [release paper](#).

The Python examples from the release paper are available below:

## EAGLE Public Data Release

Welcome to the EAGLE public data release.

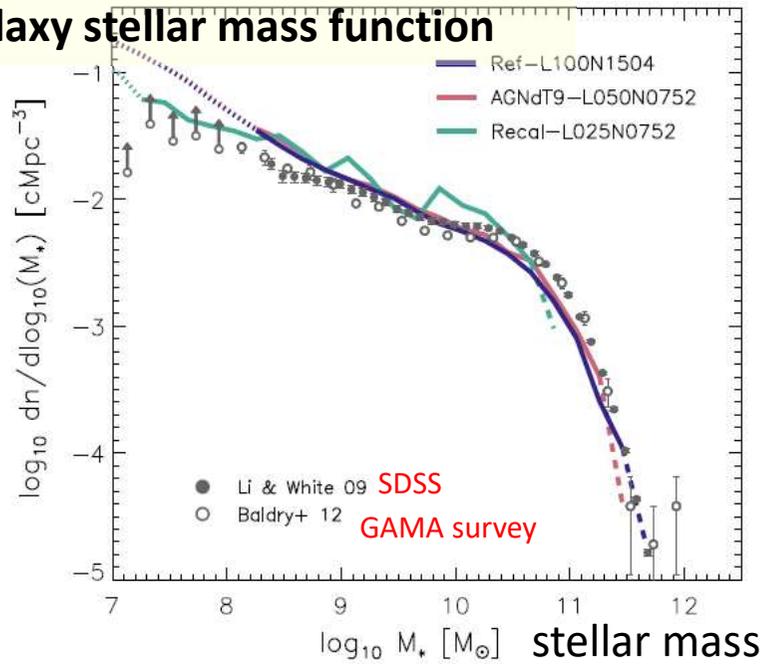
Below we provide access to the EAGLE public database, which contains galaxy properties (such as masses, star formation rates, luminosities and metallicities), merger histories and images for more than 1,000,000 simulated galaxies extracted from multiple simulations of various box sizes, numerical resolutions and physical models. Additionally, we also provide access to the particle data for each of these simulations.

Access to both the database and particle data require only a single account, which you can register for below.

<http://icc.dur.ac.uk/Eagle/database.php>

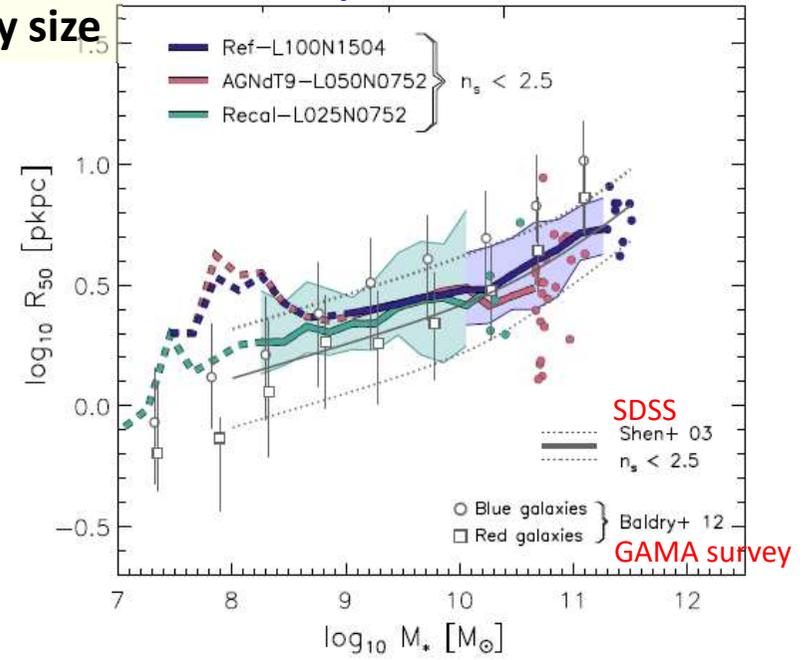
# Calibrating subgrid recipes against observations

## galaxy stellar mass function

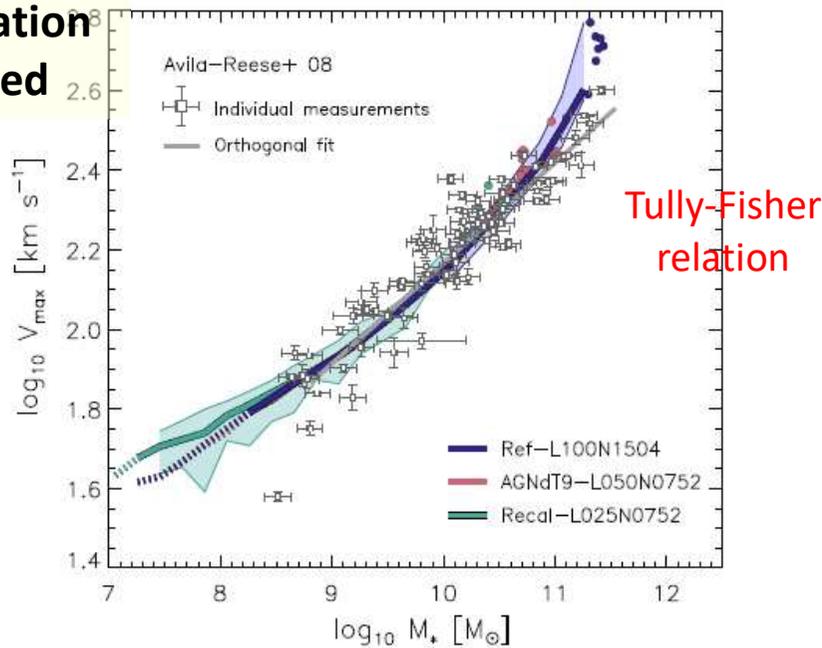


## Projected half-mass radius

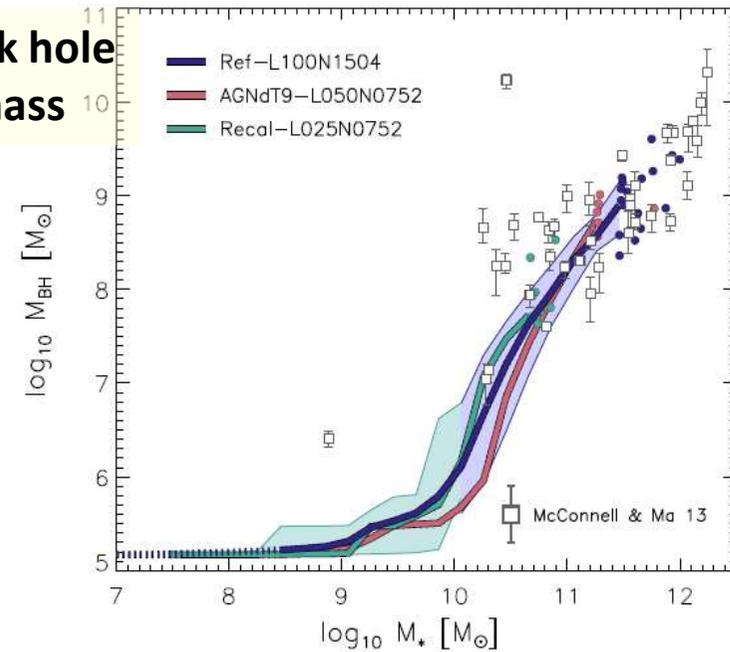
## galaxy size



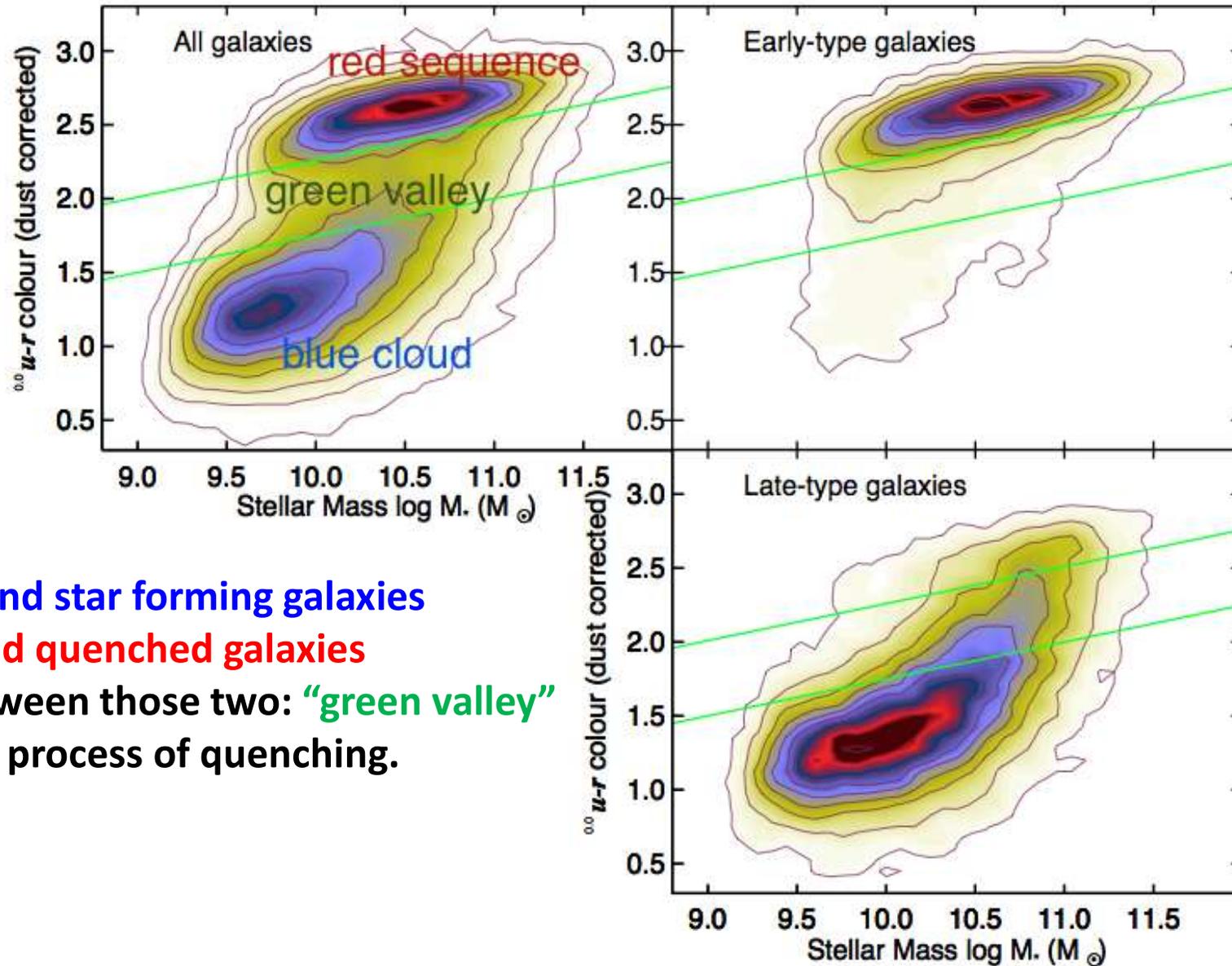
## rotation speed



## black hole mass

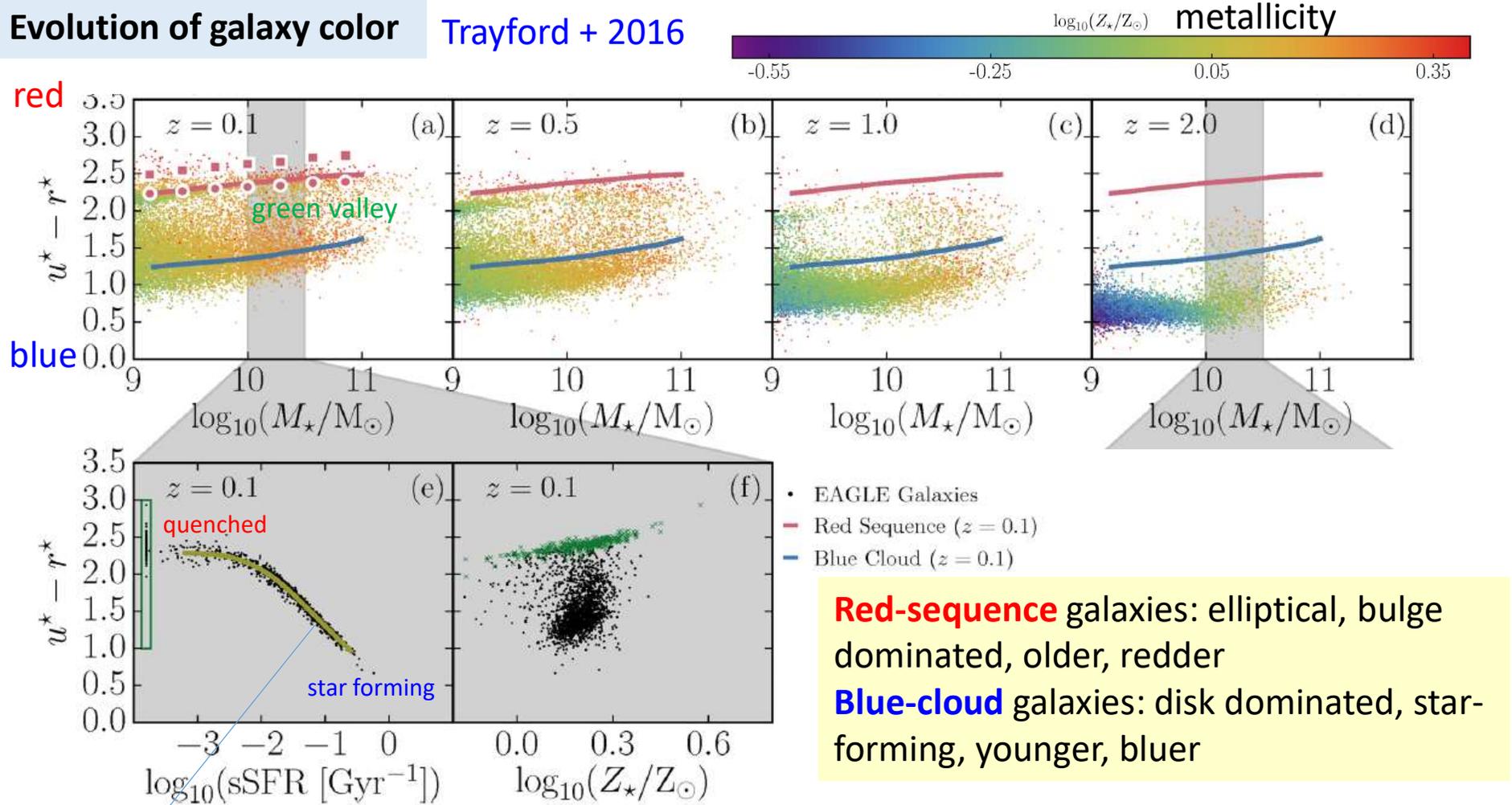


# Observational Data: SDSS+GALEX+Galaxy Zoo data



- blue and star forming galaxies
- red and quenched galaxies
- in between those two: “green valley”  
in the process of quenching.

# Evolution of galaxy color Trayford + 2016



-**Bimodality** in color at  $z=0.1$  : **red** & **blue** sequences.

-**Redder** with increasing  $M^*$ , as more massive galaxies are more **metal rich** and hence **redder**

-**red sequence** becomes bluer and less populated at higher  $z$

-**blue sequence** becomes **bluer** with increasing  $z$

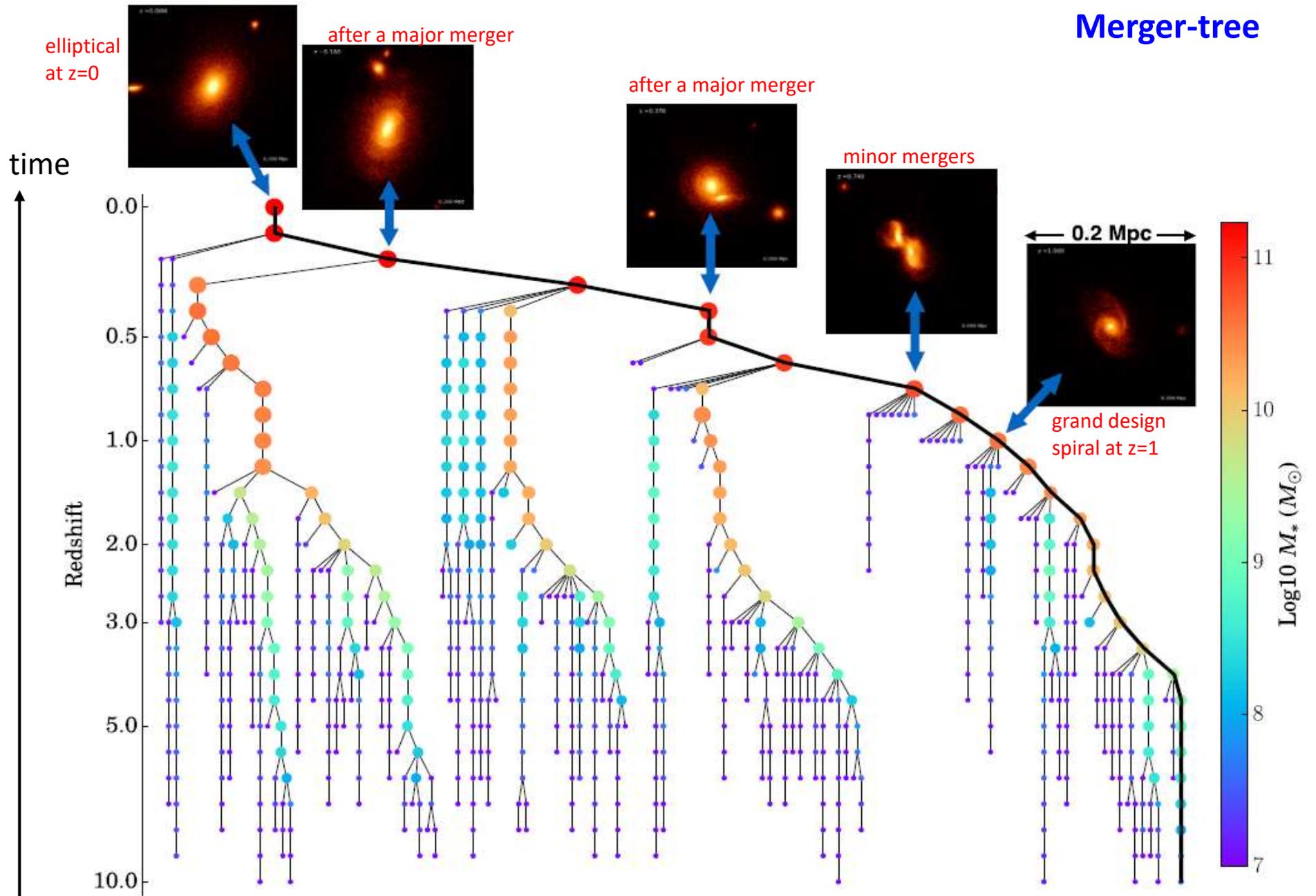
→ -**bluer** with higher SFR: strong (anti)correlation

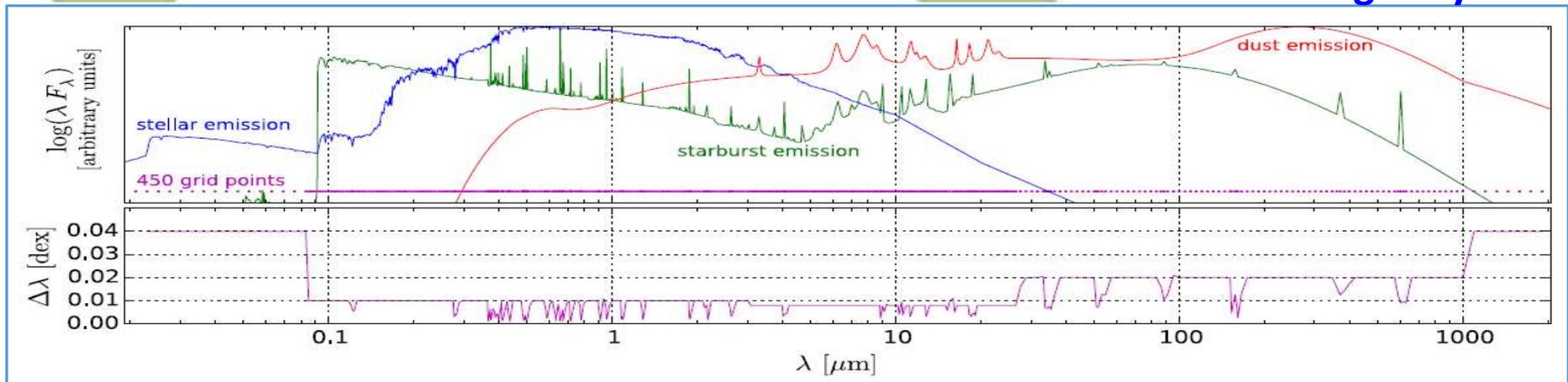
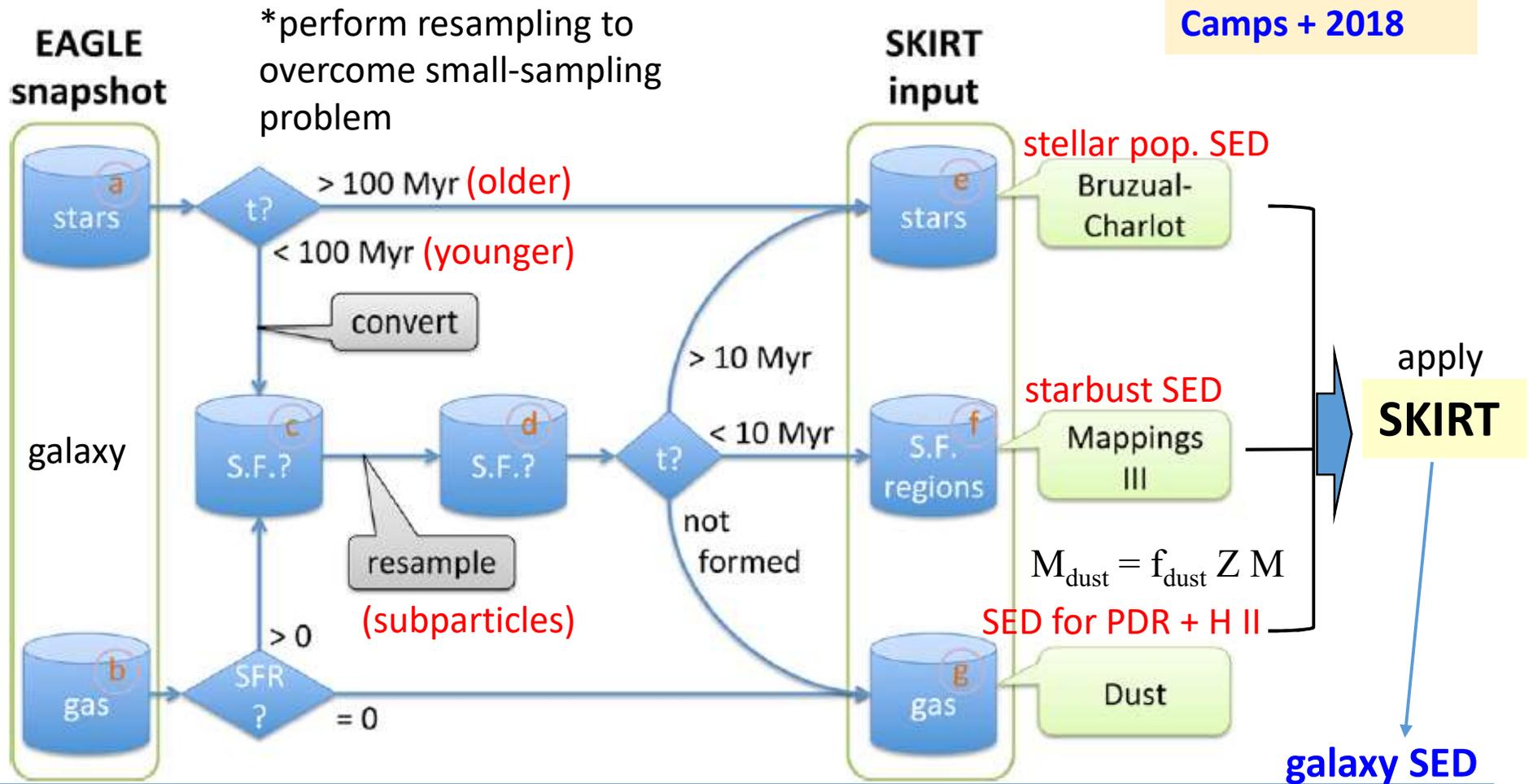
-it is not being easy being **green**:  $\Delta t_{\text{green}} < 2 \text{ Gyr}$

# Following the evolution of galaxy properties with merger history

Qu + 2017

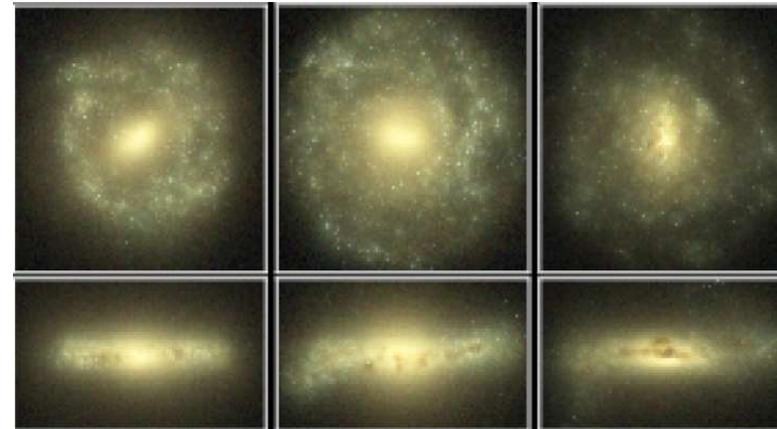
## Merger-tree



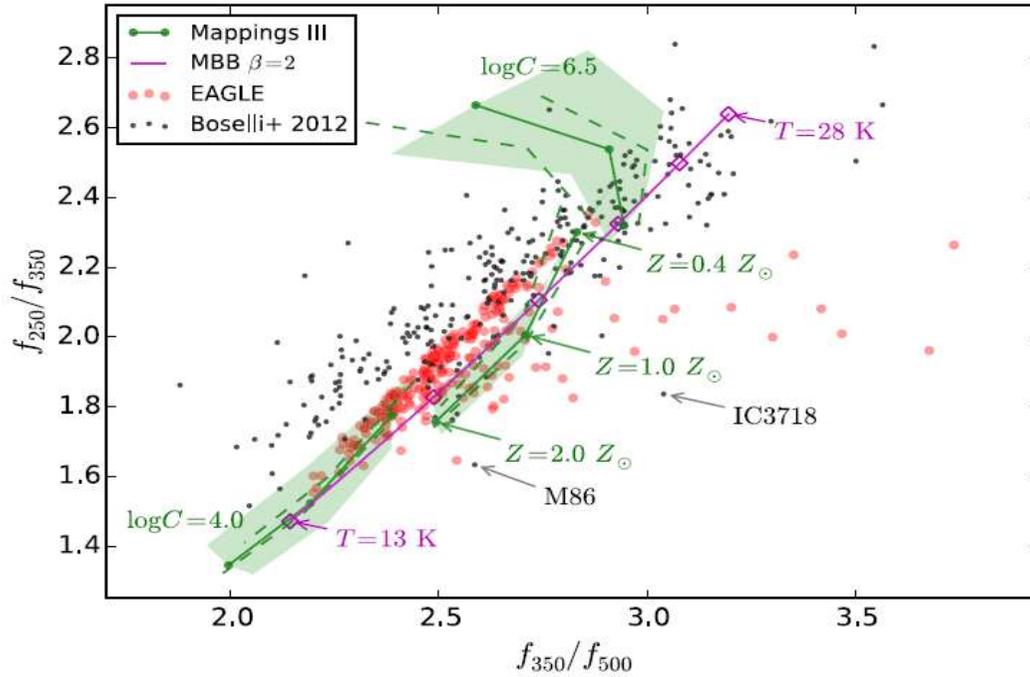


EAGLE Model	$L$ (cMpc)	$m_g$ ( $M_\odot$ )	Number of Galaxies with $M_* > 10^{8.5} M_\odot$		
			All	With Some Dust	With Resolved Dust
RefL0025N0752	25	$2.26 \times 10^5$	8279	8096 (97.8%)	7819 (94.4%)
RecalL0025N0752	25	$2.26 \times 10^5$	5954	5886 (98.9%)	5700 (95.7%)
RefL0025N0376	25	$1.81 \times 10^6$	5742	5553 (96.7%)	3871 (67.4%)
RefL0050N0752	50	$1.81 \times 10^6$	48,261	44,470 (92.1%)	31,422 (65.1%)
AGNdT9L0050N0752	50	$1.81 \times 10^6$	48,278	44,601 (92.4%)	31,231 (64.7%)
RefL0100N1504	100	$1.81 \times 10^6$	371,728	334,717 (90.0%)	236,346 (63.6%)
Total			488,242	443,323 (90.8%)	316,389 (64.8%)

-  *UV-FIR Photometry*
-  *UV-FIR Spectra*
-  *Broad band FITS files*
-  *IFU data cubes*



*GALEX* FUV/NUV (Morrissey et al. 2007); SDSS *ugriz* (Doi et al. 2010); 2MASS *JHK* (Cohen, Wheaton & Megeath 2003); *WISE* W1/W2/W3/W4 (Wright et al. 2010); *Spitzer* MIPS 24/70/160 (Rieke et al. 2004); *Herschel* PACS 70/100/160 (Poglitsch et al. 2010); and *Herschel* SPIRE 250/350/500 for extended sources (Griffin et al. 2010).



Herschel SPIRE:250/350/500  $\mu\text{m}$

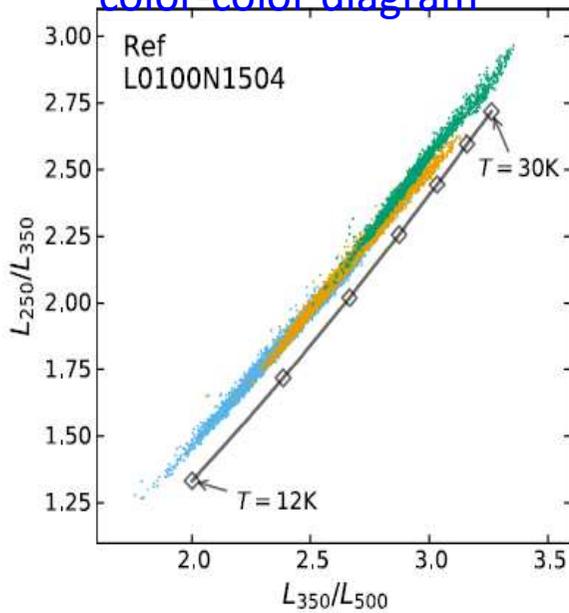
Mapping III template with different  $Z$

Modified Black Body

EAGLE galaxies

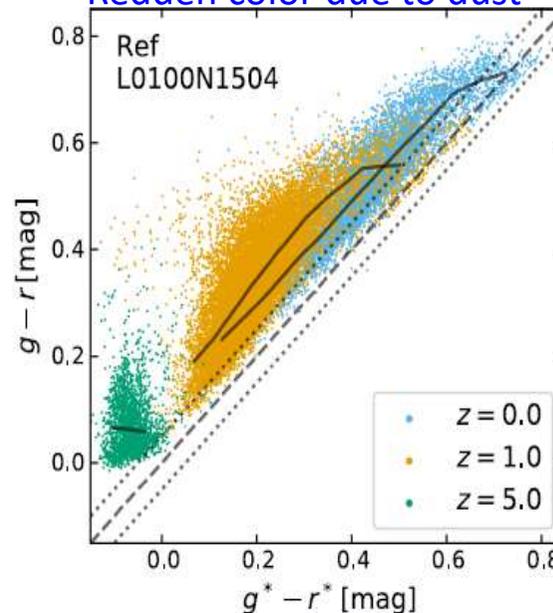
Herschel Reference Survey data

color-color diagram



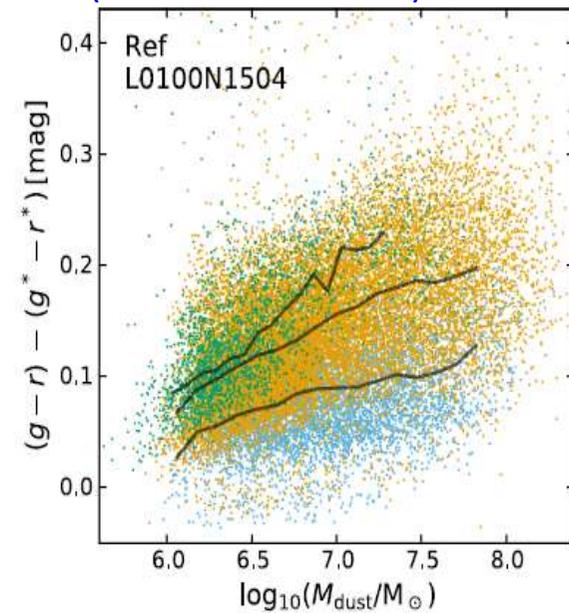
Herschel SPIRE:250/350/500  $\mu\text{m}$

Redden color due to dust



(intrinsic color)

(observed-intrinsic) color



(dust mass)

# Evolution and Assembly of GaLaxies and their Environments

## 1. SIMULATION

MNRAS 446, 521–554 (2015)

**The EAGLE project: simulating the evolution and assembly of galaxies and their environments**

Joop Schaye,<sup>1</sup>★ Robert A. Crain,<sup>1</sup> Richard G. Bower,<sup>2</sup> Michelle Furlong,<sup>2</sup>

## 2. Galaxy Catalogues

Astronomy and Computing 15 (2016) 72–89

The EAGLE simulations of galaxy formation: Public release of halo and galaxy catalogues<sup>☆</sup>

S. McAlpine<sup>a,\*</sup>, J.C. Helly<sup>a</sup>, M. Schaller<sup>a</sup>, J.W. Trayford<sup>a</sup>, Y. Qu<sup>a</sup>, M. Furlong<sup>a</sup>, R.G. Bower<sup>a</sup>,

<sup>a</sup>Department of Physics and Astronomy, University of Southampton, Highfield, Southampton, SO9 5NH, UK

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## 3. Revised Catalogues with Dust + Radiative Transfer

THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES, 234:20 (15pp), 2018 February

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Matthieu Schaller<sup>2</sup> , and Joop Schaye<sup>4</sup> 

➔ EAGLE provides a powerful tool for understanding and interpreting a wide range of observation measurements.

# No. of papers with “EAGLE” in the title since 2015: ~ 60

## Query Results from the ADS Database

Selected and retrieved 60 abstracts.

#	Bibcode Authors	Score Title	Date	List of Links Access Control Help
1	<a href="#">2018MNRAS.tmp..578R</a> Rahmati, Alireza; Oppenheimer, Benjamin D.	1.000 The metallicity distribution of HI systems in the <b>EAGLE</b> cosmological simulations	03/2018	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">U</a>
2	<a href="#">2018MNRAS.475.1288S</a> Scholtz, J.; Alexander, D. M.; Harrison, C. M.; Rosario, D. J.; McAlpine, S.; Mullaney, J. R.; Stanley, F.; Simpson, J.; Theuns, T.; Bower, R. G.; and 3 coauthors	1.000 Identifying the subtle signatures of feedback from distant AGN using ALMA observations and the <b>EAGLE</b> hydrodynamical simulations	03/2018	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">D</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
3	<a href="#">2018MNRAS.474.3746A</a> Armitage, Thomas J.; Barnes, David J.; Kay, Scott T.; Bahé, Yannick M.; Dalla Vecchia, Claudio; Crain, Robert A.; Theuns, Tom	1.000 The Cluster- <b>EAGLE</b> project: velocity bias and the velocity dispersion-mass relation of cluster galaxies	03/2018	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">U</a>
4	<a href="#">2018MNRAS.474.3403T</a> Tagore, Amitpal S.; Barnes, David J.; Jackson, Neal; Kay, Scott T.; Schaller, Matthieu; Schaye, Joop; Theuns, Tom	1.000 Reducing biases on H $\alpha$ measurements using strong lensing and galaxy dynamics: results from the <b>EAGLE</b> simulation	03/2018	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
5	<a href="#">2018ApJS..234..20C</a> Camps, Peter; Trcka, Ana; Trayford, James; Baes, Maarten; Theuns, Tom; Crain, Robert A.; McAlpine, Stuart; Schaller, Matthieu; Schaye, Joop	1.000 Data Release of UV to Submillimeter Broadband Fluxes for Simulated Galaxies from the <b>EAGLE</b> Project	02/2018	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
6	<a href="#">2018MNRAS.473..380T</a> Tescari, E.; Cortese, L.; Power, C.; Wyithe, J. S. B.; Ho, I.-T.; Crain, R. A.; Bland-Hawthorn, J.; Croom, S. M.; Kewley, L. J.; Schaye, J.; and 16 coauthors	1.000 The SAMI Galaxy Survey: understanding observations of large-scale outflows at low redshift with <b>EAGLE</b> simulations	01/2018	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">S</a> <a href="#">U</a>
7	<a href="#">2018arXiv180108781B</a> Barnes, Luke A.; Elahi, Pascal J.; Salcido, Jaime; Bower, Richard G.; Lewis, Geraint F.; Theuns, Tom; Schaller, Matthieu; Crain, Robert A.; Schaye, Joop	1.000 Galaxy Formation Efficiency and the Multiverse Explanation of the Cosmological Constant with <b>EAGLE</b> Simulations	01/2018	<a href="#">A</a> <a href="#">X</a> <a href="#">R</a> <a href="#">U</a>
8	<a href="#">2018AAS...23114904K</a> Kulier, Andrea; Padilla, Nelson; Schaye, Joop; Crain, Robert; Schaller, Matthieu; Bower, Richard; Theuns, Tom; Paillas, Enrique	1.000 Examining the effect of galaxy evolution on the stellar-halo mass relation in the <b>EAGLE</b> simulation	01/2018	<a href="#">A</a> <a href="#">U</a>
9	<a href="#">2017MNRAS.472.3354D</a> De Rossi, Maria Emilia; Bower, Richard G.; Font, Andreea S.; Schaye, Joop; Theuns, Tom	1.000 Galaxy metallicity scaling relations in the <b>EAGLE</b> simulations	12/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
10	<a href="#">2017MNRAS.472L..45C</a> Correa, Camila A.; Schaye, Joop; Clauwens, Bart; Bower, Richard G.; Crain, Robert A.; Schaller, Matthieu; Theuns, Tom; Thob, Adrien C. R.	1.000 The relation between galaxy morphology and colour in the <b>EAGLE</b> simulation	11/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
11	<a href="#">2017MNRAS.472..919K</a> Katsianis, A.; Blanc, G.; Lagos, C. P.; Tejos, N.; Bower, R. G.; Alavi, A.; Gonzalez, V.; Theuns, T.; Schaller, M.; Lopez, S.	1.000 The evolution of the star formation rate function in the <b>EAGLE</b> simulations: a comparison with UV, IR and H $\alpha$ observations from $z \sim 8$ to $z \sim 0$	11/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
12	<a href="#">2017MNRAS.471.2856V</a> Velliscig, Marco; Cacciato, Marcello; Hoekstra, Henk; Schaye, Joop; Heymans, Catherine; Hildebrandt, Hendrik; Loveday, Jon; Norberg, Peder; Sifón, Cristóbal; Schneider, Peter; and 7 coauthors	1.000 Galaxy-galaxy lensing in <b>EAGLE</b> : comparison with data from 180 deg $^2$ of the KiDS and GAMA surveys	11/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">D</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
13	<a href="#">2017MNRAS.471L..11D</a> Desmond, Harry; Mao, Yao-Yuan; Wechsler, Risa H.; Crain, Robert A.; Schaye, Joop	1.000 On the galaxy-halo connection in the <b>EAGLE</b> simulation	10/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
14	<a href="#">2017MNRAS.471.1088B</a> Barnes, David J.; Kay, Scott T.; Bahé, Yannick M.; Dalla Vecchia, Claudio; McCarthy, Ian G.; Schaye, Joop; Bower, Richard G.; Jenkins, Adrian; Thomas, Peter A.; Schaller, Matthieu; and 3 coauthors	1.000 The Cluster- <b>EAGLE</b> project: global properties of simulated clusters with resolved galaxies	10/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">S</a> <a href="#">U</a>
15	<a href="#">2017MNRAS.470.4434P</a> Paillas, Enrique; Lagos, Claudia D. P.; Padilla, Nelson; Tissera, Patricia; Helly, John; Schaller, Matthieu	1.000 Baryon effects on void statistics in the <b>EAGLE</b> simulation	10/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
16	<a href="#">2017MNRAS.470.1771A</a> Artale, M. Celeste; Pedrosa, Susana E.; Trayford, James W.; Theuns, Tom; Farrow, Daniel J.; Norberg, Peder; Zehavi, Idit; Bower, Richard G.; Schaller, Matthieu	1.000 Small-scale galaxy clustering in the <b>eagle</b> simulation	09/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>
17	<a href="#">2017MNRAS.470..771T</a> Trayford, James W.; Camps, Peter; Theuns, Tom; Baes, Maarten; Bower, Richard G.; Crain, Robert A.; Gunawardhana, Madusha L. P.; Schaller, Matthieu; Schaye, Joop; Frenk, Carlos S.	1.000 Optical colours and spectral indices of $z = 0.1$ <b>eagle</b> galaxies with the 3D dust radiative transfer code skirt	09/2017	<a href="#">A</a> <a href="#">E</a> <a href="#">F</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">S</a> <a href="#">U</a>
18	<a href="#">2017arXiv170907570</a> Oppenheimer, Benjamin D.; Schaye, Joop; Crain, Robert A.; Werk, Jessica K.; Richings, Alexander J.	1.000 The multiphase circumgalactic medium traced by low metal ions in <b>EAGLE</b> zoom simulations	09/2017	<a href="#">A</a> <a href="#">X</a> <a href="#">R</a> <a href="#">C</a> <a href="#">U</a>



# The Illustris Simulation

$3 \times 1820^3 = 18.1 \times 10^9$   
cells / particles / tracers

106.5 Mpc boxsize

$M_{\text{baryon}} = 1.26 \times 10^5 M_{\odot}$   
 $M_{\text{dm}} = 6.26 \times 10^5 M_{\odot}$

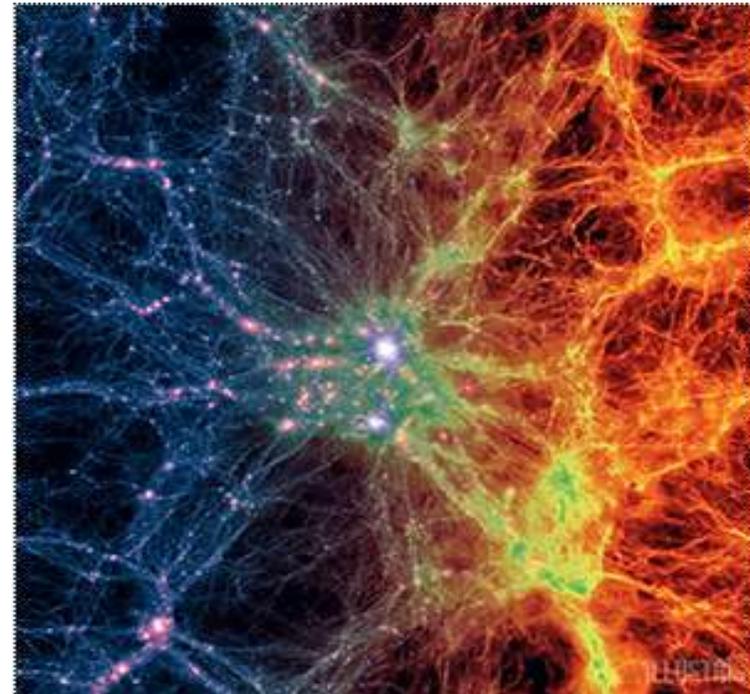
~50 pc smallest cell size

16 (+3) million CPU hours

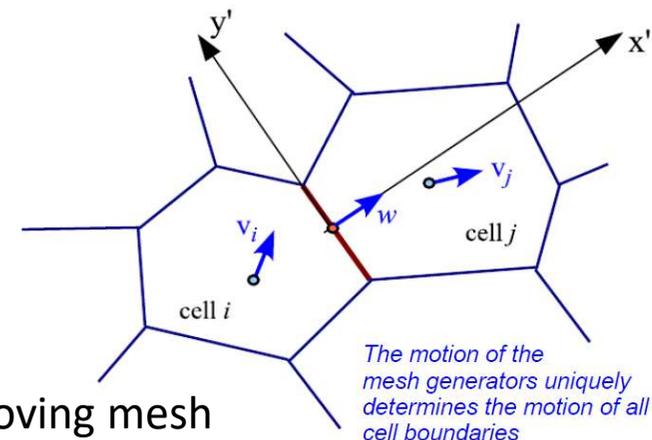
[www.illustris-project.org](http://www.illustris-project.org)

**primordial & metal line cooling**  
**+ self-shielding**  
**stellar evolution**  
**stellar feedback**  
**gas recycling**  
**chemical enrichment**  
**black hole growth**  
**black hole feedback:**  
**quasar, radiative and radio bubbles**

AREPO codel (Springel 2010): unstructured moving mesh



Large scale projection through the Illustris volume at  $z=0$ . Shows dark matter density (left) transitioning to gas density (right).



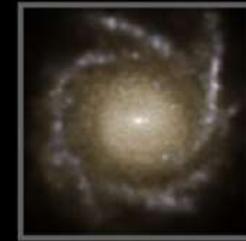
*The motion of the mesh generators uniquely determines the motion of all cell boundaries*

# The Illustris simulation reproduces the morphological mix of galaxies

## SIMULATED HUBBLE TUNING FORK DIAGRAM



**ellipticals**



**disk galaxies**

**irregular**

