

ENERGY ESTIMATOR OF
ULTRA-HIGH ENERGY
COSMIC RAYS(UHECRS)
USING GEANT4 SIMULATION

2010-09-10

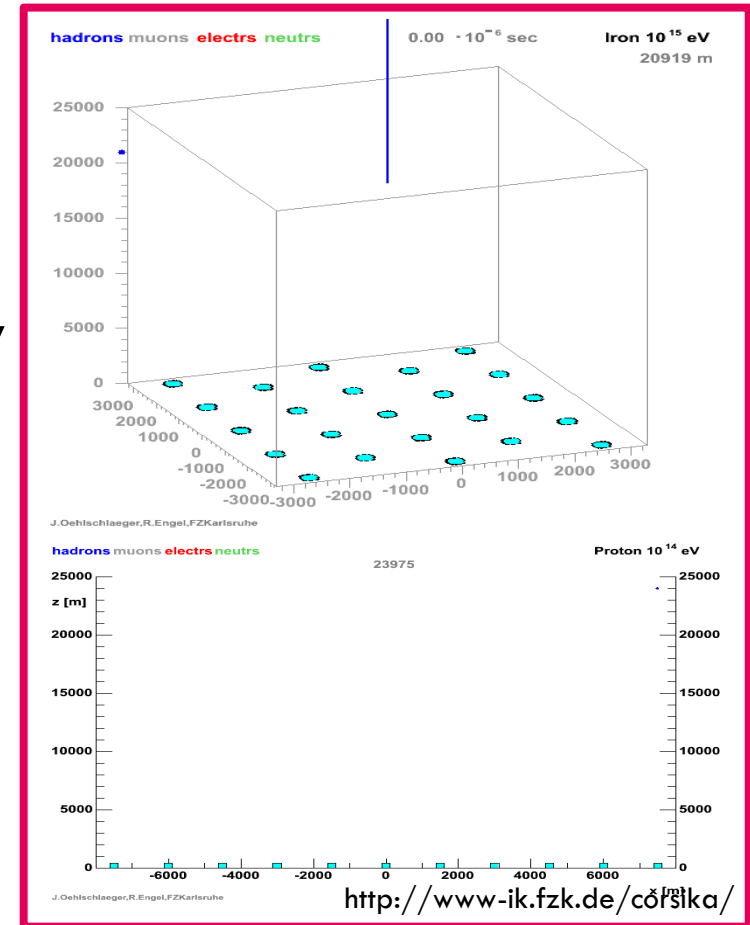
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UHECRs(Ultra-High Energy Cosmic Rays)

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- Intensity of UHECRs
1 event
per 1 km² area per century
for particles with energies $*E_0 > 10^{18} \text{ eV}$
 $* E_0$ is primary cosmic ray energy
- However,
the shower cascade in air
→ Extensive Air shower(EAS)
→ billions of sub-particles
on the ground



Ground Array Method

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- Sample the charged secondary shower particles as they reach the ground
 - distributed surface detectors measure the energy loss generated by particles
 - the primary energy from the particle density

Ground Array ? - TA Surface Detectors



KNAG Meeting 2010-09-10

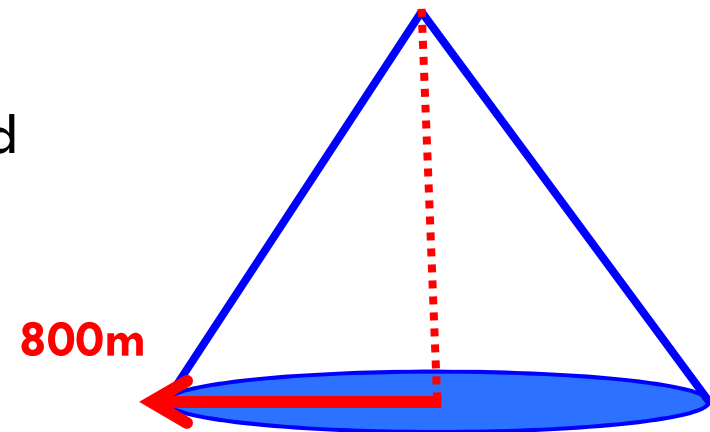
A Good Energy Estimator, $S(800)$

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- **$S(800)$** means the particle density at 800m from the shower core!
⇒ by measuring energy deposition on surface detector

- **$S(\text{distance from shower core})$, for example, $S(800)$**
 - Fluctuation of the local number density of shower particles far from the core is significantly small.

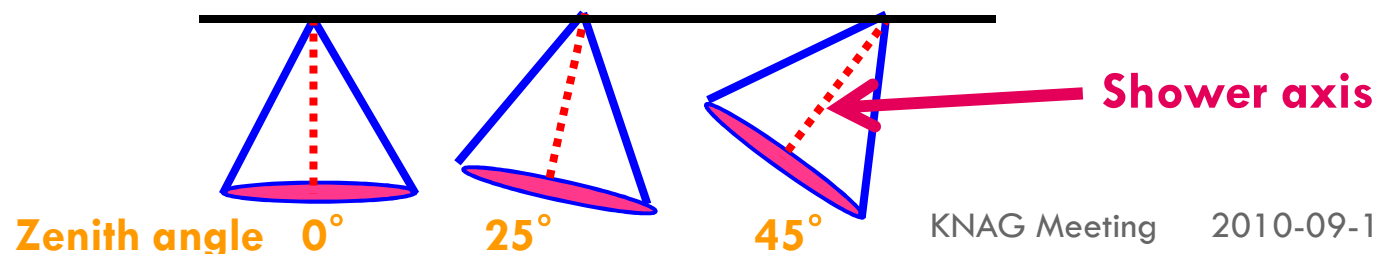
- The past cosmic ray experiments used
 - $S(600)$ in AGASA
 - $S(1000)$ in Auger
 - **$S(800)$ in TA ?**



Monte Carlo Simulations

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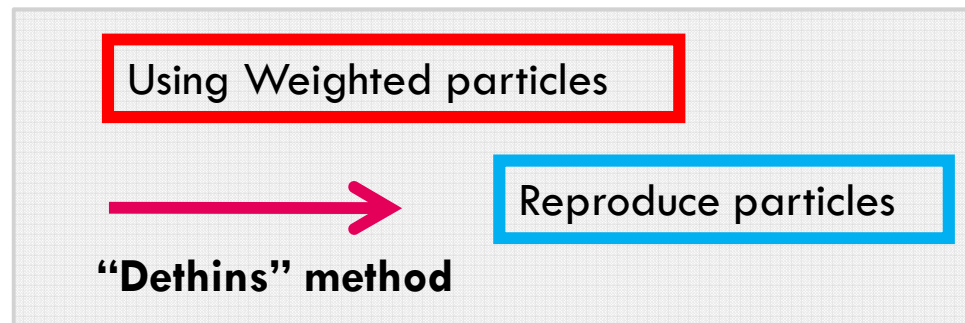
- CORSIKA and COSMOS simulation
- Hadronic Interaction Model
- Cut energies
 - EM component 500 keV
 - The other component 50 MeV
- **Generate AS event (p, Fe)**
 - Primary energy $10^{18.5} \sim 10^{20.25}$ eV 0.25 step in \log_{10}
 - Zenith angle $1.00 \sim 0.70$ 0.05 steps in cosine



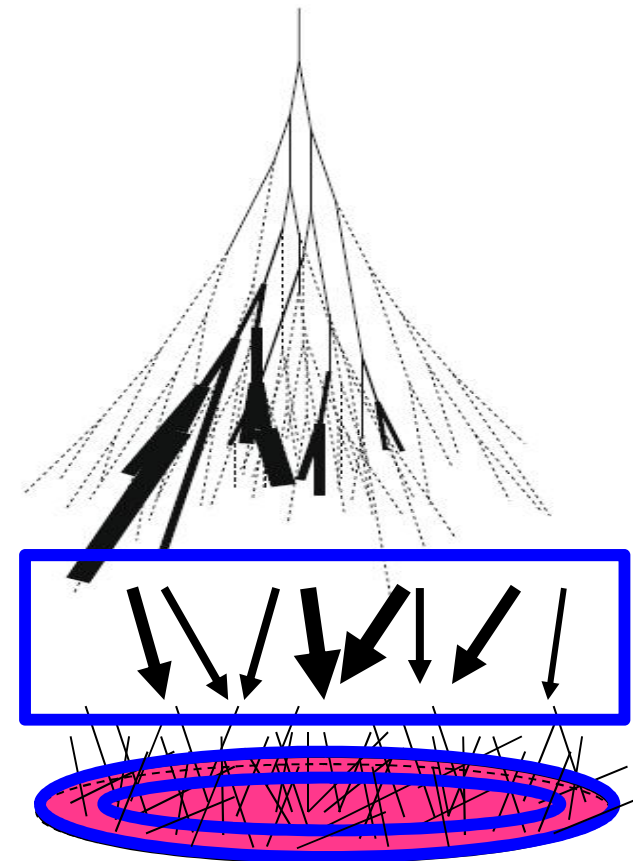
Particle Energy Spectrum

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- Generate Air shower
 - “thinning method”
- Using sub-particles



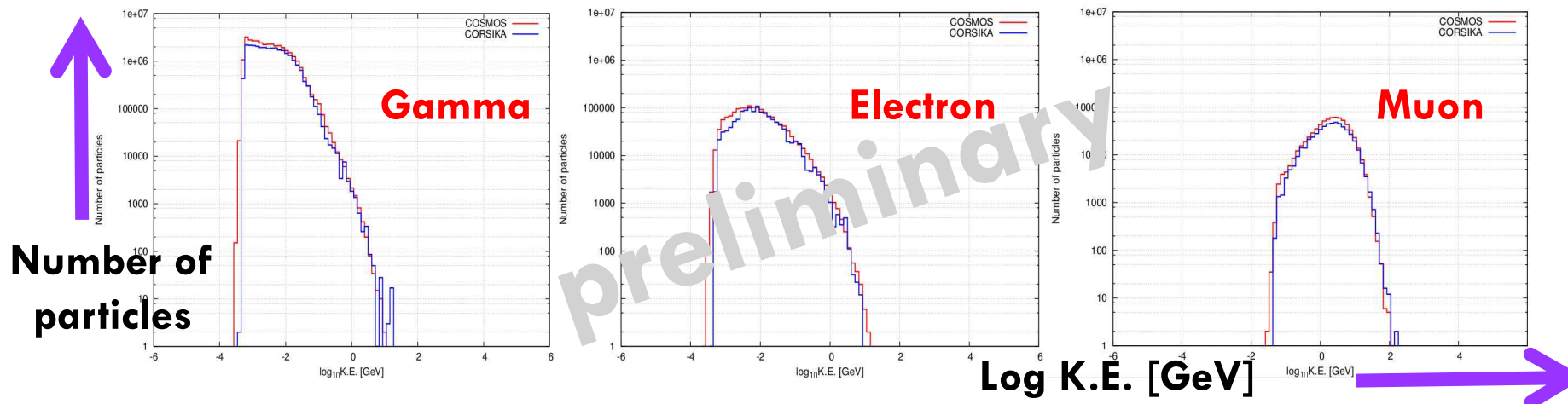
- Acquire the particles falling in the ring (near 800m)



800m ←

Particle Energy Spectrum in the ring(750~850m)

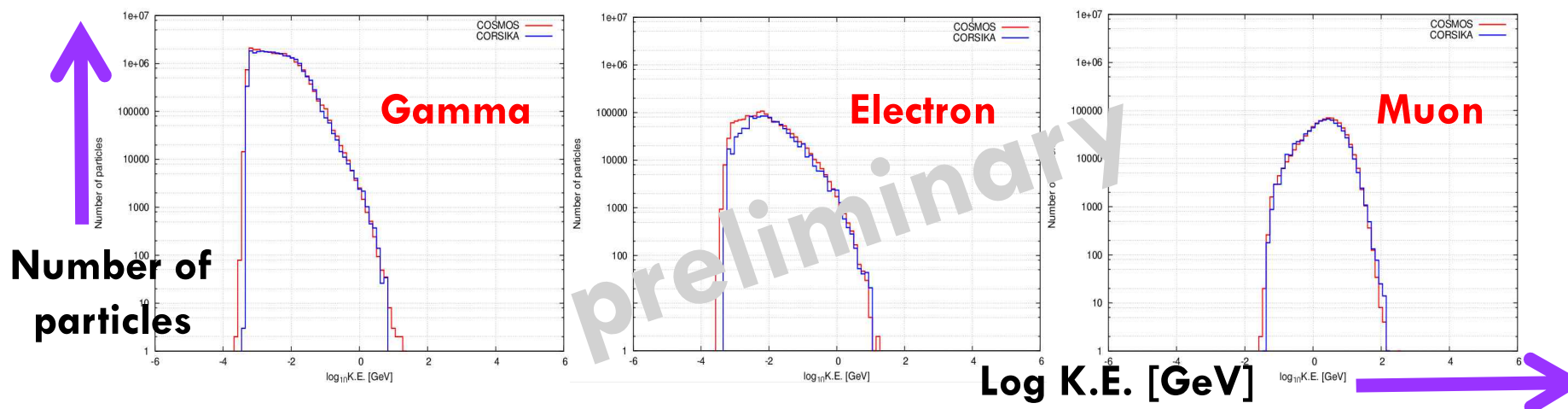
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| Proton Primaries | | COSMOS Sim. | CORSIKA Sim. | Difference |
|------------------|------------------------|--------------|--------------|------------|
| Gamma | Total energy | 2.618819e+05 | 2.202126e+05 | 19% |
| | Total no. of particles | 3.498243e+07 | 2.936467e+07 | 19% |
| Electron | Total energy | 3.895400e+04 | 3.075245e+04 | 27% |
| | Total no. of particles | 1.516596e+06 | 1.226405e+06 | 24% |
| Muon | Total energy | 1.906912e+06 | 1.613041e+06 | 18% |
| | Total no. of particles | 6.532425e+05 | 5.301190e+05 | 23% |

Particle Energy Spectrum in the ring(750~850m)

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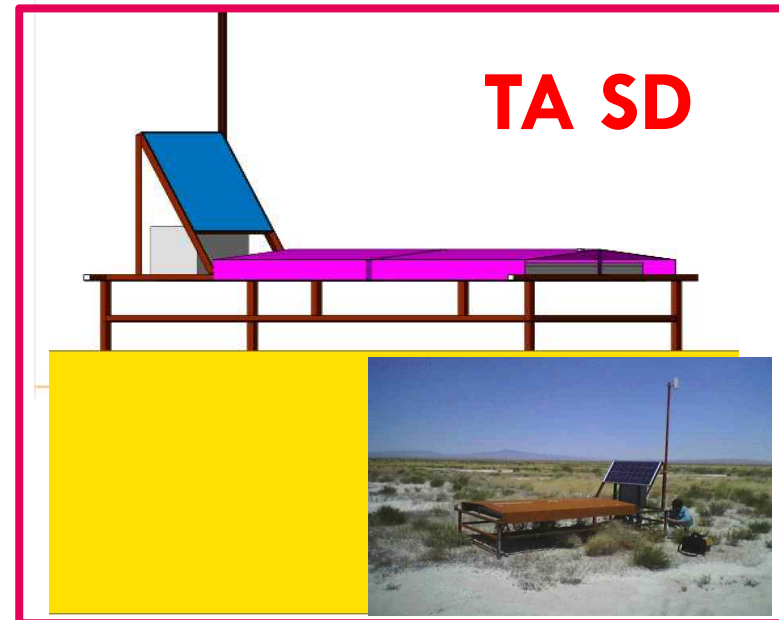
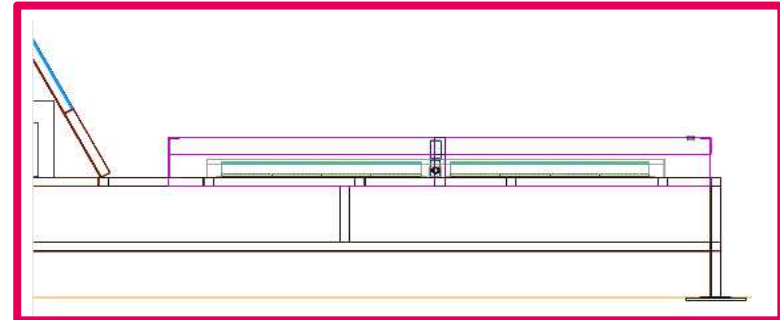
| Iron Primaries | | COSMOS Sim. | CORSIKA Sim. | Difference |
|----------------|------------------------|--------------|--------------|------------|
| Gamma | Total energy | 2.043813e+05 | 2.095098e+05 | 2.5% |
| | Total no. of particles | 2.382721e+07 | 2.493880e+07 | 4.5% |
| Electron | Total energy | 3.913562e+04 | 3.568663e+04 | 9.7% |
| | Total no. of particles | 1.303323e+06 | 1.167950e+06 | 1.2% |
| Muon | Total energy | 2.569609e+06 | 2.249879e+06 | 1.4% |
| | Total no. of particles | 7.453085e+05 | 7.046200e+05 | 5.8% |

Geant4 Simulations

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- I used TA(Telescope Array) SD(Surface Detector) set which is made by one of our TA group collaborators

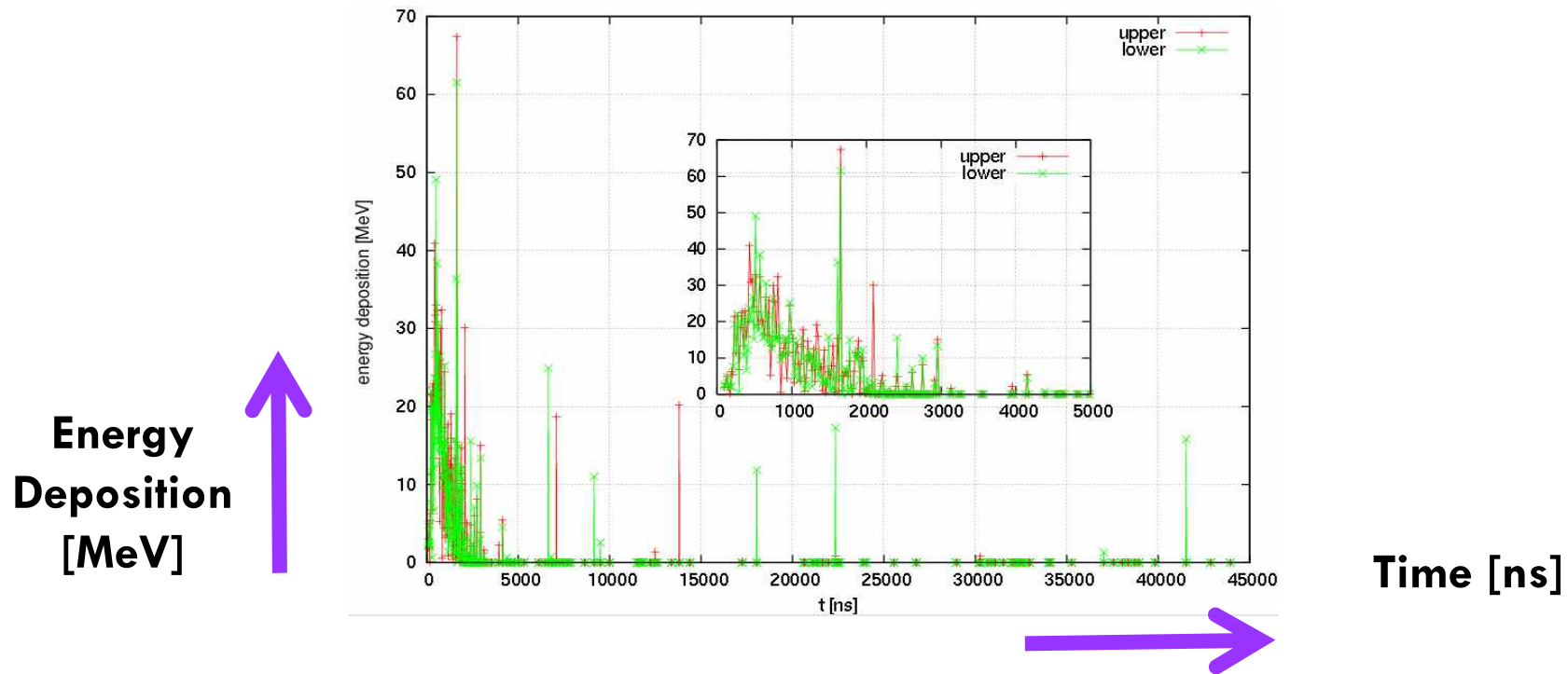
- What we need
 - ✓ Particle ID
 - ✓ charge of particle
 - ✓ arrival time
 - ✓ momentum
 - ✓ energy of particle
 - ✓ position
 - ✓ incident angle



Detector Response

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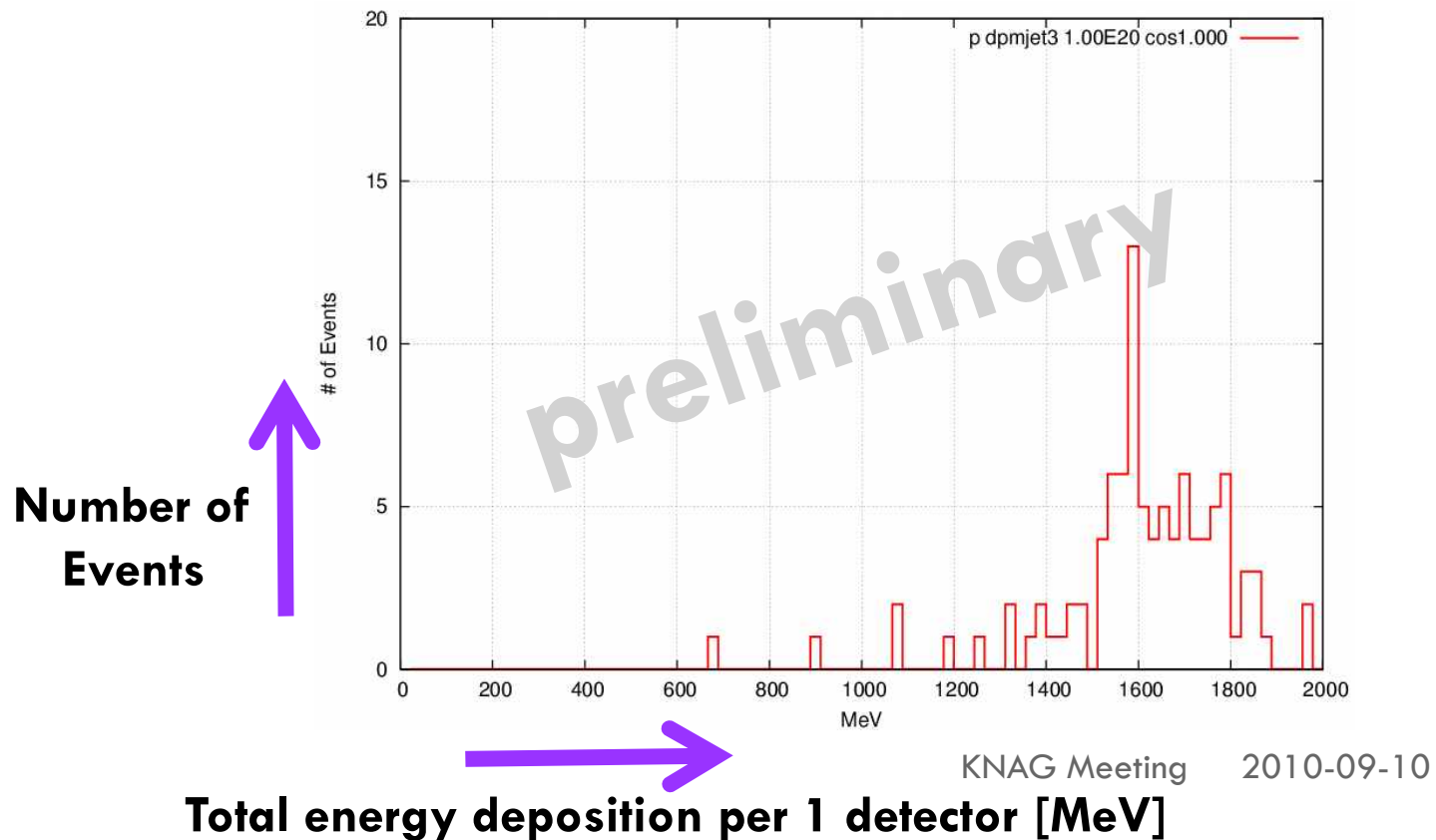
- 3 m² area, 2 layers
 - Upper & Lower layer are separated optically
- ← Using GEANT4 simulation



Energy Deposition S(800)

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- Proton primaries
- 10^{20} eV with a zenith angle of 0°

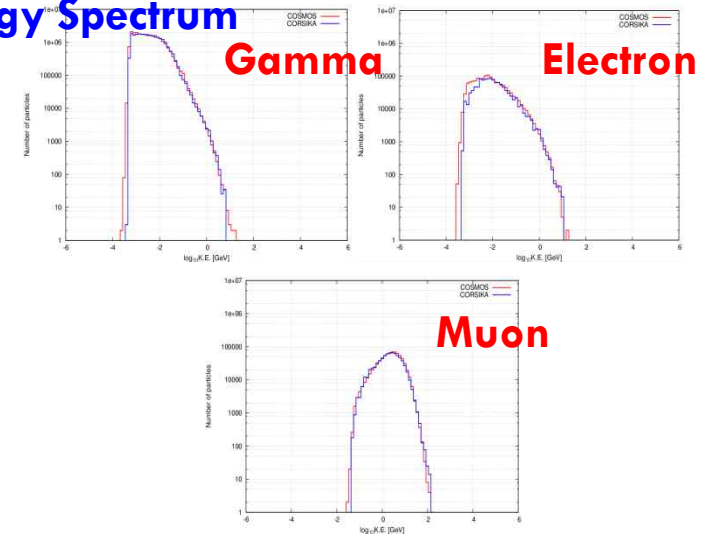


Summary

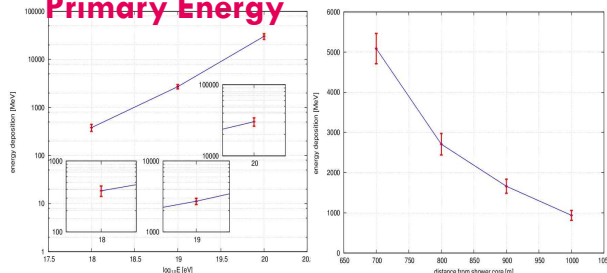
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- Particle Energy Spectrum near 800m from core
- ✓ Energy of particle arriving at the ground COSMOS Sim. > CORSIKA Sim.
- ✓ In Iron shower, CORSIKA has a lot of low energy particles than COSMOS.

Energy Spectrum



Dependence of Primary Energy



Diff. distance from shower core

- Energy Deposition
- ✓ The energy deposition far from the core is indeed stably proportional to primary energy.
- ✓ The fluctuation of the local energy deposition far from the core is significantly small.

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Back-Up Slides

- ✓ Current CORSIKA/COSMOS simulation library

Current CORSIKA/COSMOS simulation library

Table1. CORSIKA simulation library

Table2. COSMOS simulation library

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| Primary energy Cos(Zenith angle) | 10 ^{18.5} eV | 10 ^{18.75} eV | 10 ¹⁹ eV | 10 ^{19.25} eV | 10 ^{19.50} eV | 10 ^{19.75} eV | 10 ²⁰ eV | 10 ^{20.25} eV |
|-------------------------------------|-----------------------|------------------------|---------------------|------------------------|------------------------|------------------------|---------------------|------------------------|
| 1.00 | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe |
| 0.95 | | | | | | | | |
| 0.90 | | | | | | | | |
| 0.85 | | | | | | | | |
| 0.80 | | | | | | | | |
| 0.75 | | | | | | | | |
| 0.70 | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe |

| Primary energy Cos(Zenith angle) | 10 ^{18.5} eV | 10 ^{18.75} eV | 10 ¹⁹ eV | 10 ^{19.25} eV | 10 ^{19.50} eV | 10 ^{19.75} eV | 10 ²⁰ eV | 10 ^{20.25} eV |
|-------------------------------------|-----------------------|------------------------|---------------------|------------------------|------------------------|------------------------|---------------------|------------------------|
| 1.00 | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe |
| 0.95 | | | | | 10p | | | |
| 0.90 | 50p,50Fe | 40p | 40p | 40p | 40p | 30p | 30p | 8p |
| 0.85 | | | | | 10p | | | |
| 0.80 | 10p,10Fe | 10p,10Fe | 10p,10Fe | 10p,10Fe | 10p,10Fe | 10p,10Fe | 10p,10Fe | 10p,10Fe |
| 0.75 | | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | | 50p,50Fe | |
| 0.70 | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe | 50p,50Fe |